

LOOKING FOR THE HIGGS BOSON WITHOUT (TOO MUCH) PREJUDICE

► ROBERTO CONTINO
(UNIVERSITY OF ROME "LA SAPIENZA")

The starting point (and main assumption):

► Evidence for Electroweak Symmetry Breaking

massive (light) spin-1 particles $\{W_\mu^\pm, Z_\mu\}$

$$\{\chi^\pm, \chi^0\}$$

$$\{W_\mu^T, Z_\mu^T\}$$

Longitudinal polarizations =
Nambu-Goldstone bosons of $SU(2)_L \times U(1)_Y \rightarrow U(1)_{em}$

strongly coupled at $E \sim 4\pi f = 4\pi \left(\frac{m_V}{g}\right)$

Transverse polarizations =
gauge fields

elementary up to $E \gg 4\pi \left(\frac{m_V}{g}\right)$

Top-Down Approach

UV

PERTURBATIVE MODEL

$SU(2)_L \times U(1)_Y$ Linearly Realized



IR

LOW-ENERGY
PHENOMENOLOGY

verify model's predictions

in this talk I will rather follow a

Bottom-Up Approach

Scenario #1
no linear regime

UV strong dynamics
(new resonances ρ , ...)



IR effective theory of χ^i

in this talk I will rather follow a

Bottom-Up Approach

Scenario #1

no linear regime

UV

strong dynamics
(new resonances ρ , ...)



IR

effective theory of χ^i

Scenario #2

$SU(2)_L \times U(1)_Y$ linear
+ perturbativity

UV

weakly coupled theory

$$[\chi^i, \phi^a]$$



IR

effective theory of χ^i

in this talk I will rather follow a

Bottom-Up Approach

Scenario #1

no linear regime

UV

strong dynamics
(new resonances ρ , ...)

IR

effective theory of χ^i



Scenario #2

$SU(2)_L \times U(1)_Y$ linear
+ perturbativity

UV

weakly coupled theory

$$[\chi^i, \phi^a]$$

Higgs bosons

IR

effective theory of χ^i



in this talk I will rather follow a

Bottom-Up Approach

Scenario #1

no linear regime

UV strong dynamics
(new resonances ρ, \dots)

IR effective theory of χ^i



Scenario #2

$SU(2)_L \times U(1)_Y$ linear
+ perturbativity

weakly coupled theory

$[\chi^i, \phi^a]$

Higgs bosons

UV

IR

effective theory of χ^i



Scenario #3

$SU(2)_L \times U(1)_Y$ linear
+ strong dynamics

strong dynamics
(new resonances ρ, \dots)

$[\chi^i, \phi^a]$



IR effective theory of χ^i

in this talk I will rather follow a

Bottom-Up Approach

- ▶ If a light scalar is discovered at the LHC, we want to experimentally determine which scenario (#1, #2 or #3) is realized

in this talk I will rather follow a

Bottom-Up Approach

- ▶ If a light scalar is discovered at the LHC, we want to experimentally determine which scenario (#1, #2 or #3) is realized

Notice: the light scalar might be an “impostor” (light dilaton ?), have nothing to do with EWBS, so that we would be in scenario #1 (no linear regime)

Rules (CCWZ):

[Coleman,Wess,Zumino PRD 117 (1969) 2239

Callan,Coleman,Wess,Zumino PRD 117 (1969) 2247]

- [1.] Any $U(1)_{em}$ locally-invariant Lagrangian can be dressed up with NG-bosons and rewritten as manifestly $SU(2)_L \times U(1)_Y$ invariant

Ex: $\text{Tr}[W_\mu^2] \longrightarrow \text{Tr}[(D_\mu \Sigma)^\dagger (D_\mu \Sigma)] \quad (\Sigma = \exp(i\chi/v))$

$$W_\mu^+ W_\mu^- Z_{\mu\nu} \longrightarrow \text{Tr}[(D_\mu \Sigma)^\dagger \sigma^a W_{\mu\nu}^a (D_\nu \Sigma)]$$

- [2.] Fields must come in multiplets of $U(1)_{em}$
(i.e. not necessarily of $SU(2)_L \times U(1)_Y$)

Possible Extra Rules:

- ▶ The EWSB sector has an approximate custodial $SU(2)_c$ invariance:

$$SU(2)_L \times U(1)_Y \rightarrow U(1)_{em} \quad \Rightarrow \quad SU(2)_L \times SU(2)_R \rightarrow SU(2)_c$$

or equivalently

$$SO(4) \rightarrow SO(3)$$

Possible Extra Rules:

- The EWSB sector has an approximate custodial $SU(2)_c$ invariance:



see Grojean's talk for a bottom-up approach w/o custodial symmetry

$$SU(2)_L \times U(1)_Y \rightarrow U(1)_{em} \quad \Rightarrow \quad SU(2)_L \times SU(2)_R \rightarrow SU(2)_c$$

or equivalently

$$SO(4) \rightarrow SO(3)$$

Possible Extra Rules:

- The EWSB sector has an approximate custodial $SU(2)_c$ invariance:



see Grojean's talk for a bottom-up approach w/o custodial symmetry

$$SU(2)_L \times U(1)_Y \rightarrow U(1)_{\text{em}} \quad \Rightarrow \quad SU(2)_L \times SU(2)_R \rightarrow SU(2)_c$$

or equivalently

$$SO(4) \rightarrow SO(3)$$

- The Higgs is a (pseudo) NG boson of a larger symmetry breaking

Ex: $SO(5) \rightarrow SO(4)$

$$H = [\chi^i, h]$$

Additional invariance
(broken by spurions only):

$$h(x) \rightarrow h(x) + \alpha$$

$$T_h \in \text{Alg} \left\{ \frac{SO(5)}{SO(4)} \right\}$$

(Higgs shift symmetry)

Chiral Lagrangian for a light Higgs

Assumptions:

- ▶ Higgs boson is a scalar, singlet of $U(1)_{\text{em}}$ ($SU(2)_c$)
- ▶ No extra light particles
- ▶ No tree-level FCNC (mediated by the Higgs)

Chiral Lagrangian for a light Higgs

Assumptions:

- ▶ Higgs boson is a scalar, singlet of $U(1)_{\text{em}}$ ($SU(2)_c$)
- ▶ No extra light particles
- ▶ No tree-level FCNC (mediated by the Higgs)

Chiral expansion parameter = $\left(\frac{\partial_\mu}{\Lambda}\right)$ $\Lambda \lesssim 4\pi v = \Lambda_s$

The cutoff can be made larger if
the Higgs partly unitarizes all
scattering amplitudes

Ex: Higgs is a pNG boson

$$\Lambda \lesssim 4\pi f = \frac{4\pi v}{\sqrt{\xi}} \quad \xi = \left(\frac{v}{f}\right)^2$$

Chiral Lagrangian for a light Higgs

$$\begin{aligned}\mathcal{L} = & \frac{1}{2}(\partial_\mu h)^2 - \frac{1}{2}m_h^2 h^2 - \frac{d_3}{6} \left(\frac{3m_h^2}{v} \right) h^3 - \frac{d_4}{24} \left(\frac{3m_h^2}{v^2} \right) h^4 \dots \\ & - \left(m_W^2 W_\mu W_\mu + \frac{1}{2} m_Z^2 Z_\mu Z_\mu \right) \left(1 + 2a \frac{h}{v} + b \frac{h^2}{v^2} + \dots \right) \\ & - \sum_{\psi=u,d,l} m_{\psi^{(i)}} \bar{\psi}^{(i)} \psi^{(i)} \left(1 + c_\psi \frac{h}{v} + c_{2\psi} \frac{h^2}{v^2} + \dots \right) \\ & + \frac{g^2}{16\pi^2} \left(c_{WW} W_{\mu\nu}^+ W_{\mu\nu}^- + c_{ZZ} Z_{\mu\nu}^2 + c_{Z\gamma} Z_{\mu\nu} \gamma_{\mu\nu} \right) \frac{h}{v} + \dots \\ & + \frac{g^2}{16\pi^2} \left[\gamma_{\mu\nu}^2 \left(c_{\gamma\gamma} \frac{h}{v} + \dots \right) + G_{\mu\nu}^2 \left(c_{gg} \frac{h}{v} + c_{2gg} \frac{h^2}{v^2} \dots \right) \right] \\ & + \frac{g^2}{16\pi^2} \left[\frac{c_{hhgg}}{\Lambda^2} G_{\mu\nu}^2 \frac{(\partial_\rho h)^2}{v^2} + \frac{c'_{hhgg}}{\Lambda^2} G_{\mu\rho} G_{\rho\nu} \frac{\partial_\mu h \partial_\nu h}{v^2} + \dots \right] \\ & + \dots\end{aligned}$$

Chiral Lagrangian for a light Higgs

$$\mathcal{L} = \frac{1}{2}(\partial_\mu h)^2 - \frac{1}{2}m_h^2 h^2 - \frac{d_3}{6} \left(\frac{3m_h^2}{v} \right) h^3 - \frac{d_4}{24} \left(\frac{3m_h^2}{v^2} \right) h^4 \dots$$

$$- \left(m_W^2 W_\mu W_\mu + \frac{1}{2} m_Z^2 Z_\mu Z_\mu \right) \left(1 + 2a \frac{h}{v} + b \frac{h^2}{v^2} + \dots \right)$$

$$- \sum_{\psi=u,d,l} m_{\psi^{(i)}} \bar{\psi}^{(i)} \psi^{(i)} \left(1 + c_\psi \frac{h}{v} + c_{2\psi} \frac{h^2}{v^2} + \dots \right)$$

$$+ \frac{g^2}{16\pi^2} \left(c_{WW} W_{\mu\nu}^+ W_{\mu\nu}^- + c_{ZZ} Z_{\mu\nu}^2 + c_{Z\gamma} Z_{\mu\nu} \gamma_{\mu\nu} \right) \frac{h}{v} + \dots$$

$$+ \frac{g^2}{16\pi^2} \left[\gamma_{\mu\nu}^2 \left(c_{\gamma\gamma} \frac{h}{v} + \dots \right) + G_{\mu\nu}^2 \left(c_{gg} \frac{h}{v} + c_{2gg} \frac{h^2}{v^2} \dots \right) \right]$$

$$+ \frac{g^2}{16\pi^2} \left[\frac{c_{hhgg}}{\Lambda^2} G_{\mu\nu}^2 \frac{(\partial_\rho h)^2}{v^2} + \frac{c'_{hhgg}}{\Lambda^2} G_{\mu\rho} G_{\rho\nu} \frac{\partial_\mu h \partial_\nu h}{v^2} + \dots \right]$$

+ ...

$\mathbf{o(p^2)}$ terms

Chiral Lagrangian for a light Higgs

$$\mathcal{L} = \frac{1}{2}(\partial_\mu h)^2 - \frac{1}{2}m_h^2 h^2 - \frac{d_3}{6} \left(\frac{3m_h^2}{v} \right) h^3 - \frac{d_4}{24} \left(\frac{3m_h^2}{v^2} \right) h^4 \dots$$

$$- \left(m_W^2 W_\mu W_\mu + \frac{1}{2} m_Z^2 Z_\mu Z_\mu \right) \left(1 + 2a \frac{h}{v} + b \frac{h^2}{v^2} + \dots \right)$$

$$- \sum_{\psi=u,d,l} m_{\psi^{(i)}} \bar{\psi}^{(i)} \psi^{(i)} \left(1 + c_\psi \frac{h}{v} + c_{2\psi} \frac{h^2}{v^2} + \dots \right)$$

$$+ \frac{g^2}{16\pi^2} \left(c_{WW} W_{\mu\nu}^+ W_{\mu\nu}^- + c_{ZZ} Z_{\mu\nu}^2 + c_{Z\gamma} Z_{\mu\nu} \gamma_{\mu\nu} \right) \frac{h}{v} + \dots$$

$$+ \frac{g^2}{16\pi^2} \left[\gamma_{\mu\nu}^2 \left(c_{\gamma\gamma} \frac{h}{v} + \dots \right) + G_{\mu\nu}^2 \left(c_{gg} \frac{h}{v} + c_{2gg} \frac{h^2}{v^2} \dots \right) \right]$$

$$+ \frac{g^2}{16\pi^2} \left[\frac{c_{hhgg}}{\Lambda^2} G_{\mu\nu}^2 \frac{(\partial_\rho h)^2}{v^2} + \frac{c'_{hhgg}}{\Lambda^2} G_{\mu\rho} G_{\rho\nu} \frac{\partial_\mu h \partial_\nu h}{v^2} + \dots \right]$$

+ ...

o(p²) terms

o(p⁴) terms

Chiral Lagrangian for a light Higgs

$$\mathcal{L} = \frac{1}{2}(\partial_\mu h)^2 - \frac{1}{2}m_h^2 h^2 - \frac{d_3}{6} \left(\frac{3m_h^2}{v} \right) h^3 - \frac{d_4}{24} \left(\frac{3m_h^2}{v^2} \right) h^4 \dots$$

$$- \left(m_W^2 W_\mu W_\mu + \frac{1}{2} m_Z^2 Z_\mu Z_\mu \right) \left(1 + 2a \frac{h}{v} + b \frac{h^2}{v^2} + \dots \right)$$

$$- \sum_{\psi=u,d,l} m_{\psi^{(i)}} \bar{\psi}^{(i)} \psi^{(i)} \left(1 + c_\psi \frac{h}{v} + c_{2\psi} \frac{h^2}{v^2} + \dots \right)$$

$$+ \frac{g^2}{16\pi^2} \left(c_{WW} W_{\mu\nu}^+ W_{\mu\nu}^- + c_{ZZ} Z_{\mu\nu}^2 + c_{Z\gamma} Z_{\mu\nu} \gamma_{\mu\nu} \right) \frac{h}{v} + \dots$$

$$+ \frac{g^2}{16\pi^2} \left[\gamma_{\mu\nu}^2 \left(c_{\gamma\gamma} \frac{h}{v} + \dots \right) + G_{\mu\nu}^2 \left(c_{gg} \frac{h}{v} + c_{2gg} \frac{h^2}{v^2} + \dots \right) \right]$$

$$+ \frac{g^2}{16\pi^2} \left[\frac{c_{hhgg}}{\Lambda^2} G_{\mu\nu}^2 \frac{(\partial_\rho h)^2}{v^2} + \frac{c'_{hhgg}}{\Lambda^2} G_{\mu\rho} G_{\rho\nu} \frac{\partial_\mu h \partial_\nu h}{v^2} + \dots \right]$$

+ ...

o(p²) terms

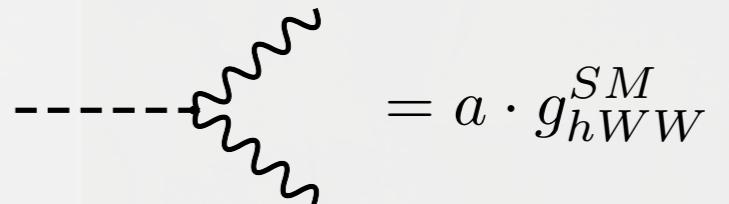
o(p⁴) terms

o(p⁶) terms

Chiral Lagrangian for a light Higgs

Controls the hWW , hZZ couplings

$$\begin{aligned}
 \mathcal{L} = & \frac{1}{2}(\partial_\mu h)^2 - \frac{1}{2}m_h^2 h^2 - \frac{d_3}{6} \left(\frac{3m_h^2}{v} \right) h^3 - \frac{d_4}{24} \left(\frac{3m_h^2}{v^2} \right) h^4 \dots \\
 & - \left(m_W^2 W_\mu W_\mu + \frac{1}{2} m_Z^2 Z_\mu Z_\mu \right) \left(1 + 2a \frac{h}{v} + b \frac{h^2}{v^2} + \dots \right) \\
 & - \sum_{\psi=u,d,l} m_{\psi^{(i)}} \bar{\psi}^{(i)} \psi^{(i)} \left(1 + c_\psi \frac{h}{v} + c_{2\psi} \frac{h^2}{v^2} + \dots \right) \\
 & + \frac{g^2}{16\pi^2} \left(c_{WW} W_{\mu\nu}^+ W_{\mu\nu}^- + c_{ZZ} Z_{\mu\nu}^2 + c_{Z\gamma} Z_{\mu\nu} \gamma_{\mu\nu} \right) \frac{h}{v} + \dots \\
 & + \frac{g^2}{16\pi^2} \left[\gamma_{\mu\nu}^2 \left(c_{\gamma\gamma} \frac{h}{v} + \dots \right) + G_{\mu\nu}^2 \left(c_{gg} \frac{h}{v} + c_{2gg} \frac{h^2}{v^2} \dots \right) \right] \\
 & + \frac{g^2}{16\pi^2} \left[\frac{c_{hhgg}}{\Lambda^2} G_{\mu\nu}^2 \frac{(\partial_\rho h)^2}{v^2} + \frac{c'_{hhgg}}{\Lambda^2} G_{\mu\rho} G_{\rho\nu} \frac{\partial_\mu h \partial_\nu h}{v^2} + \dots \right] \\
 & + \dots
 \end{aligned}$$



Chiral Lagrangian for a light Higgs

Controls the $h\psi\psi$ coupling

$$-\mathcal{L} = \frac{1}{2}(\partial_\mu h)^2 = \frac{1}{2}c_{\psi^2} g_{h\psi\psi}^{SM} \frac{d_3}{6}$$

Controls the $h\gamma\gamma$, $h\gamma Z$ couplings

$$= a \cdot g_{hWW}^{SM}$$

$$\left(\frac{3m_h^2}{v} \right) h^3 - \frac{d_4}{24} \left(\frac{3m_h^2}{v^2} \right) h^4 \dots$$

$$- \left(m_W^2 W_\mu W_\mu + \frac{1}{2} m_Z^2 Z_\mu Z_\mu \right) \left(1 + 2a \frac{h}{v} + b \frac{h^2}{v^2} + \dots \right)$$

$$- \sum_{\psi=u,d,l} m_{\psi^{(i)}} \bar{\psi}^{(i)} \psi^{(i)} \left(1 + c_\psi \frac{h}{v} + c_{2\psi} \frac{h^2}{v^2} + \dots \right)$$

$$+ \frac{g^2}{16\pi^2} \left(c_{WW} W_{\mu\nu}^+ W_{\mu\nu}^- + c_{ZZ} Z_{\mu\nu}^2 + c_{Z\gamma} Z_{\mu\nu} \gamma_{\mu\nu} \right) \frac{h}{v} + \dots$$

$$+ \frac{g^2}{16\pi^2} \left[\gamma_{\mu\nu}^2 \left(c_{\gamma\gamma} \frac{h}{v} + \dots \right) + G_{\mu\nu}^2 \left(c_{gg} \frac{h}{v} + c_{2gg} \frac{h^2}{v^2} \dots \right) \right]$$

$$+ \frac{g^2}{16\pi^2} \left[\frac{c_{hhgg}}{\Lambda^2} G_{\mu\nu}^2 \frac{(\partial_\rho h)^2}{v^2} + \frac{c'_{hhgg}}{\Lambda^2} G_{\mu\rho} G_{\rho\nu} \frac{\partial_\mu h \partial_\nu h}{v^2} + \dots \right]$$

+ ...

Chiral Lagrangian for a light Higgs

Controls the $h\psi\psi$ coupling

$$\text{---} \mathcal{L} = \frac{1}{2} (\partial_\mu h)^2 = \frac{1}{2} c_{\psi\psi} g_{h\psi\psi}^{SM} \frac{d_3}{6}$$

$$- \left(m_W^2 W_\mu W_\mu + \frac{1}{2} m_Z^2 Z_\mu Z_\mu \right) \left(1 + 2a \frac{h}{v} + b \frac{h^2}{v^2} + \dots \right)$$

$$- \sum_{\psi=u,d,l} m_{\psi^{(i)}} \bar{\psi}^{(i)} \psi^{(i)} \left(1 + c_\psi \frac{h}{v} + c_{2\psi} \frac{h^2}{v^2} + \dots \right)$$

$$+ \frac{g^2}{16\pi^2} \left(c_{WW} W_{\mu\nu}^+ W_{\mu\nu}^- + c_{ZZ} Z_{\mu\nu}^2 + c_{Z\gamma} Z_{\mu\nu} \gamma_{\mu\nu} \right) \frac{h}{v} + \dots$$

$$+ \frac{g^2}{16\pi^2} \left[\gamma_{\mu\nu}^2 \left(c_{\gamma\gamma} \frac{h}{v} + \dots \right) + G_{\mu\nu}^2 \left(c_{gg} \frac{h}{v} + c_{2gg} \frac{h^2}{v^2} \dots \right) \right]$$

$$+ \frac{g^2}{16\pi^2} \left[\frac{c_{hhgg}}{\Lambda^2} G_{\mu\nu}^2 \frac{(\partial_\rho h)^2}{v^2} + \frac{c'_{hhgg}}{\Lambda^2} G_{\mu\rho} G_{\rho\nu} \frac{\partial_\mu h \partial_\nu h}{v^2} + \dots \right]$$

+ ...

Chiral Lagrangian for a light Higgs

Controls the hWW, hZZ couplings

$$\text{---} \quad = a \cdot g_{hWW}^{SM}$$

$O(p^4)$ correction to hW couplings

$$\text{---} \otimes \quad c_{ij} \sim \left(\frac{g_* v}{\Lambda} \right)$$

Chiral Lagrangian for a light Higgs

Controls the $h\psi\psi$ coupling

$$\mathcal{L} = \frac{1}{2}(\partial_\mu h)^2 = \frac{1}{2}c_{\psi}\cdot g_{h\psi\psi}^{SM} \frac{d_3}{6}$$

$$- \left(m_W^2 W_\mu W_\mu + \frac{1}{2} m_Z^2 Z_\mu Z_\mu \right) \left(1 + 2a \frac{h}{v} + b \frac{h^2}{v^2} + \dots \right)$$

$$- \sum_{\psi=u,d,l} m_{\psi^{(i)}} \bar{\psi}^{(i)} \psi^{(i)} \left(1 + c_{\psi} \frac{h}{v} + c_{2\psi} \frac{h^2}{v^2} + \dots \right)$$

$$+ \frac{g^2}{16\pi^2} \left(c_{WW} W_{\mu\nu}^+ W_{\mu\nu}^- + c_{ZZ} Z_{\mu\nu}^2 + c_{Z\gamma} Z_{\mu\nu} \gamma_{\mu\nu} \right) \frac{h}{v} + \dots$$

$$+ \frac{g^2}{16\pi^2} \left[\gamma_{\mu\nu}^2 \left(c_{\gamma\gamma} \frac{h}{v} + \dots \right) + G_{\mu\nu}^2 \left(c_{gg} \frac{h}{v} + c_{2gg} \frac{h^2}{v^2} \dots \right) \right]$$

$$+ \frac{g^2}{16\pi^2} \left[\frac{c_{hhgg}}{\Lambda^2} G_{\mu\nu}^2 \frac{(\partial_\rho h)^2}{v^2} + \frac{c'_{hhgg}}{\Lambda^2} \right]$$

+ ...

Chiral Lagrangian for a light Higgs

Controls the hWW, hZZ couplings

$$= a \cdot g_{hWW}^{SM}$$

$O(p^4)$ correction to hW couplings

$$= \frac{c_{ij}}{\Lambda} \sim \left(\frac{g_* v}{\Lambda} \right)$$

Modify gg single production and $\gamma\gamma$ decay

$$G_{\mu\rho} G_{\nu\sigma} \frac{\partial_\mu h \partial_\nu h}{v^2} + \dots$$

$$c_{ij} \sim \left(\frac{g_* v}{\Lambda} \right) \left(\frac{\lambda^2}{g_*^2} \right)$$

Chiral Lagrangian for a light Higgs

Controls the $h\psi\psi$ coupling

$$\text{---} \mathcal{L} = \frac{1}{2} (\partial_\mu h)^2 = \frac{1}{2} \frac{m_h^2}{v^2} h^2 = c_\psi \cdot g_{h\psi\psi}^{SM} \frac{d_3}{6}$$

$$\frac{3m_h^2}{v} \Big) h^3 - \frac{d_4}{24} \left(\frac{3m_h^2}{v^2} \right) h^4 \dots$$

Controls the hWW, hZZ couplings

$$\text{---} \mathcal{L} = a \cdot g_{hWW}^{SM}$$

$$- \left(m_W^2 W_\mu W_\mu + \frac{1}{2} m_Z^2 Z_\mu Z_\mu \right) \left(1 + 2a \frac{h}{v} + b \frac{h^2}{v^2} + \dots \right)$$

$$- \sum_{\psi=u,d,l} m_{\psi^{(i)}} \bar{\psi}^{(i)} \psi^{(i)} \left(1 + c_\psi \frac{h}{v} + c_{2\psi} \frac{h^2}{v^2} + \dots \right)$$

$$+ \frac{g^2}{16\pi^2} \left(c_{WW} W_{\mu\nu}^+ W_{\mu\nu}^- + c_{ZZ} Z_{\mu\nu}^2 + c_{Z\gamma} Z_{\mu\nu} \gamma_{\mu\nu} \right) \frac{h}{v} + \dots$$

$$+ \frac{g^2}{16\pi^2} \left[\gamma_{\mu\nu}^2 \left(c_{\gamma\gamma} \frac{h}{v} + \dots \right) + G_{\mu\nu}^2 \left(c_{gg} \frac{h}{v} + c_{2gg} \frac{h^2}{v^2} \dots \right) \right]$$

$$+ \frac{g^2}{16\pi^2} \left[\frac{c_{hhgg}}{\Lambda^2} G_{\mu\nu}^2 \frac{(\partial_\rho h)^2}{v^2} + \frac{c'_{hhgg}}{\Lambda^2} \right]$$

+ ...

$O(p^4)$ correction to hW couplings

$$\text{---} \otimes \mathcal{L}$$

$$c_{ij} \sim \left(\frac{g_* v}{\Lambda} \right)$$

Modify gg single production and $\gamma\gamma$ decay

$$\text{---} \frac{\partial_\mu h \partial_\nu h}{v^2} + \dots$$

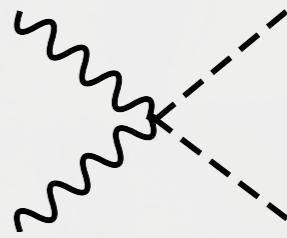
$$c_{ij} \sim \left(\frac{g_* v}{\Lambda} \right) \left(\frac{\lambda^2}{g_*^2} \right)$$

extra suppression
for a pNG Higgs

Chiral Lagrangian for a light Higgs

$$\begin{aligned}
\mathcal{L} = & \frac{1}{2}(\partial_\mu h)^2 - \frac{1}{2}m_h^2 h^2 - \frac{d_3}{6} \left(\frac{3m_h^2}{v} \right) h^3 - \frac{d_4}{24} \left(\frac{3m_h^2}{v^2} \right) h^4 \dots \\
& - \left(m_W^2 W_\mu W_\mu + \frac{1}{2} m_Z^2 Z_\mu Z_\mu \right) \left(1 + 2a \frac{h}{v} + b \frac{h^2}{v^2} + \dots \right) \\
& - \sum_{\psi=u,d,l} m_{\psi^{(i)}} \bar{\psi}^{(i)} \psi^{(i)} \left(1 + c_\psi \frac{h}{v} + c_{2\psi} \frac{h^2}{v^2} + \dots \right) \\
& + \frac{g^2}{16\pi^2} \left(c_{WW} W_{\mu\nu}^+ W_{\mu\nu}^- + c_{ZZ} Z_{\mu\nu}^2 + c_{Z\gamma} Z_{\mu\nu} \gamma_{\mu\nu} \right) \frac{h}{v} + \dots \\
& + \frac{g^2}{16\pi^2} \left[\gamma_{\mu\nu}^2 \left(c_{\gamma\gamma} \frac{h}{v} + \dots \right) + G_{\mu\nu}^2 \left(c_{gg} \frac{h}{v} + c_{2gg} \frac{h^2}{v^2} \dots \right) \right] \\
& + \frac{g^2}{16\pi^2} \left[\frac{c_{hhgg}}{\Lambda^2} G_{\mu\nu}^2 \frac{(\partial_\rho h)^2}{v^2} + \frac{c'_{hhgg}}{\Lambda^2} G_{\mu\rho} G_{\rho\nu} \frac{\partial_\mu h \partial_\nu h}{v^2} + \dots \right] \\
& + \dots
\end{aligned}$$

Contributes to $WW \rightarrow hh$



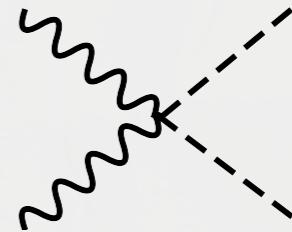
Chiral Lagrangian for a light Higgs

Contributes to $gg \rightarrow hh$

$$\mathcal{L} = \frac{1}{2}(\partial_\mu h)^2 - \frac{1}{2}m_h^2 h^2 + \frac{1}{6}t^3 \left(\frac{3m_h^2}{v} \right) h^3 - \frac{d_4}{24} \left(\frac{3m_h^2}{v^2} \right) h^4 \dots$$

$\left(m_W^2 W_\mu W_\mu + \frac{1}{2} m_Z^2 Z_\mu Z_\mu \right)$

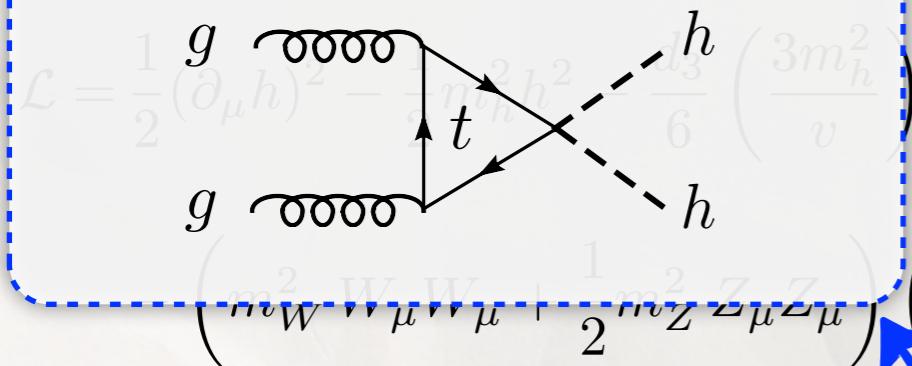
Contributes to $WW \rightarrow hh$



$$\begin{aligned}
& - \sum_{\psi=u,d,l} m_{\psi^{(i)}} \bar{\psi}^{(i)} \psi^{(i)} \left(1 + c_\psi \frac{h}{v} + c_{2\psi} \frac{h^2}{v^2} + \dots \right) \\
& + \frac{g^2}{16\pi^2} \left(c_{WW} W_\mu^+ W_\mu^- + c_{ZZ} Z_\mu^2 + c_{Z\gamma} Z_\mu \gamma_\mu \right) \frac{h}{v} + \dots \\
& + \frac{g^2}{16\pi^2} \left[\gamma_{\mu\nu}^2 \left(c_{\gamma\gamma} \frac{h}{v} + \dots \right) + G_{\mu\nu}^2 \left(c_{gg} \frac{h}{v} + c_{2gg} \frac{h^2}{v^2} + \dots \right) \right] \\
& + \frac{g^2}{16\pi^2} \left[\frac{c_{hhgg}}{\Lambda^2} G_{\mu\nu}^2 \frac{(\partial_\rho h)^2}{v^2} + \frac{c'_{hhgg}}{\Lambda^2} G_{\mu\rho} G_{\rho\nu} \frac{\partial_\mu h \partial_\nu h}{v^2} + \dots \right] \\
& + \dots
\end{aligned}$$

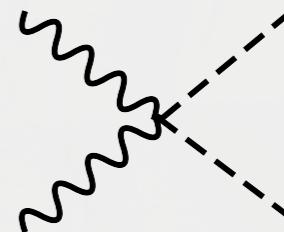
Chiral Lagrangian for a light Higgs

Contributes to $gg \rightarrow hh$



$$h^3 - \frac{d_4}{24} \left(\frac{3m_h^2}{v^2} \right) h^4 \dots$$

Contributes to $WW \rightarrow hh$



$$- \sum_{\psi=u,d,l} m_{\psi^{(i)}} \bar{\psi}^{(i)} \psi^{(i)} \left(1 + c_\psi \frac{h}{v} + c_{2\psi} \frac{h^2}{v^2} + \dots \right)$$

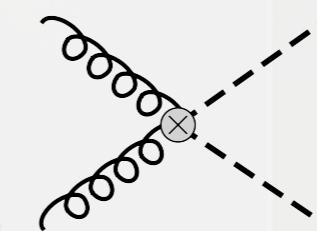
$$+ \frac{g^2}{16\pi^2} \left(c_{WW} W_{\mu\nu}^+ W_{\mu\nu}^- + c_{ZZ} Z_{\mu\nu}^2 + c_{Z\gamma} Z_{\mu\nu} \gamma_{\mu\nu} \right)$$

$$+ \frac{g^2}{16\pi^2} \left[\gamma_{\mu\nu}^2 \left(c_{\gamma\gamma} \frac{h}{v} + \dots \right) + G_{\mu\nu}^2 \left(c_{gg} \frac{h}{v} + c_{2gg} \frac{h^2}{v^2} + \dots \right) \right]$$

$$+ \frac{g^2}{16\pi^2} \left[\frac{c_{hhgg}}{\Lambda^2} G_{\mu\nu}^2 \frac{(\partial_\rho h)^2}{v^2} + \frac{c'_{hhgg}}{\Lambda^2} G_{\mu\rho} G_{\rho\nu} \frac{\partial_\mu h \partial_\nu h}{v^2} + \dots \right]$$

+ ...

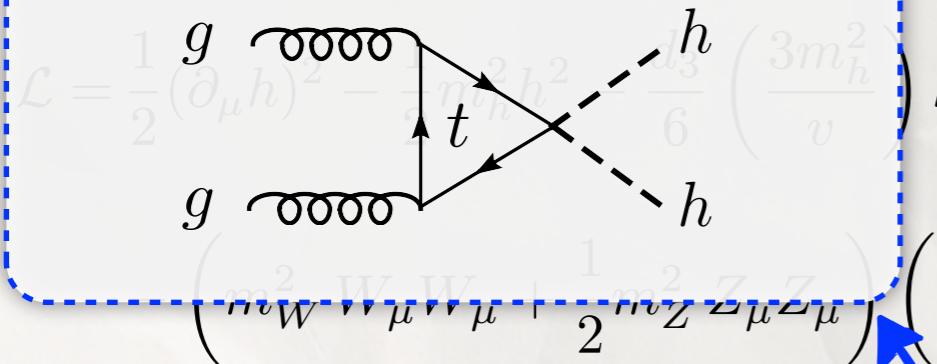
$O(p^4)$ contribution to $gg \rightarrow hh$



$$c_{2hh} \sim \left(\frac{g_*^2 v^2}{\Lambda^2} \right) \left(\frac{\lambda^2}{g_*^2} \right)$$

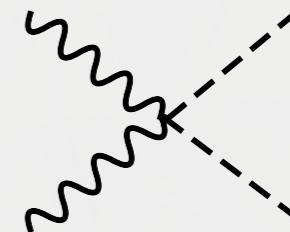
Chiral Lagrangian for a light Higgs

Contributes to $gg \rightarrow hh$

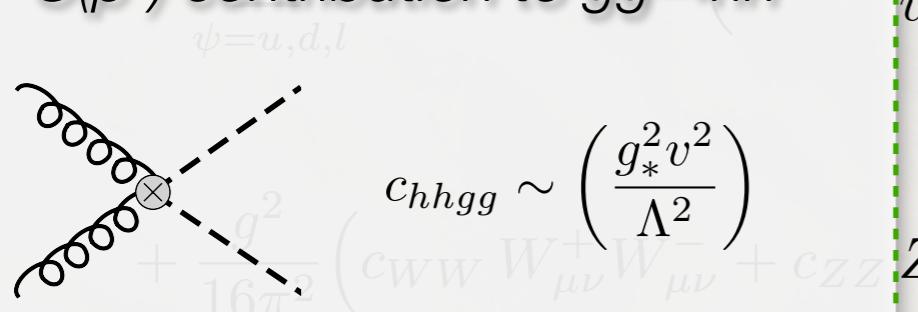


$$h^3 - \frac{d_4}{24} \left(\frac{3m_h^2}{v^2} \right) h^4 \dots$$

Contributes to $WW \rightarrow hh$

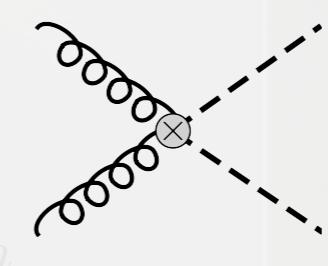


$O(p^6)$ contribution to $gg \rightarrow hh$



$$\frac{h}{v} + c_{2\psi} \frac{h^2}{v^2} + \dots$$

$O(p^4)$ contribution to $gg \rightarrow hh$



$$c_{2hh} \sim \left(\frac{g_*^2 v^2}{\Lambda^2} \right) \left(\frac{\lambda^2}{g_*^2} \right)$$

$$+ \frac{g^2}{16\pi^2} \left[\gamma_{\mu\nu}^2 \left(c_{\gamma\gamma} \frac{h}{v} + \dots \right) + G_{\mu\nu}^2 \left(c_{gg} \frac{h}{v} + c_{2gg} \frac{h^2}{v^2} + \dots \right) \right]$$

$$+ \frac{g^2}{16\pi^2} \left[\frac{c_{hhgg}}{\Lambda^2} G_{\mu\nu}^2 \frac{(\partial_\rho h)^2}{v^2} + \frac{c'_{hhgg}}{\Lambda^2} G_{\mu\rho} G_{\rho\nu} \frac{\partial_\mu h \partial_\nu h}{v^2} + \dots \right]$$

+ ...

How to use the Chiral Lagrangian

- ▶ Rules of chiral expansion:

LO = tree-level $O(p^2)$

NLO = 1-loop $O(p^2)$ + tree-level $O(p^4)$

NNLO = 2-loops $O(p^2)$ + 1-loop $O(p^4)$ + tree-level $O(p^6)$

⋮
⋮

- ▶ QCD corrections (expansion in α_s) factorize

How to use the Chiral Lagrangian

- Rules of chiral expansion:

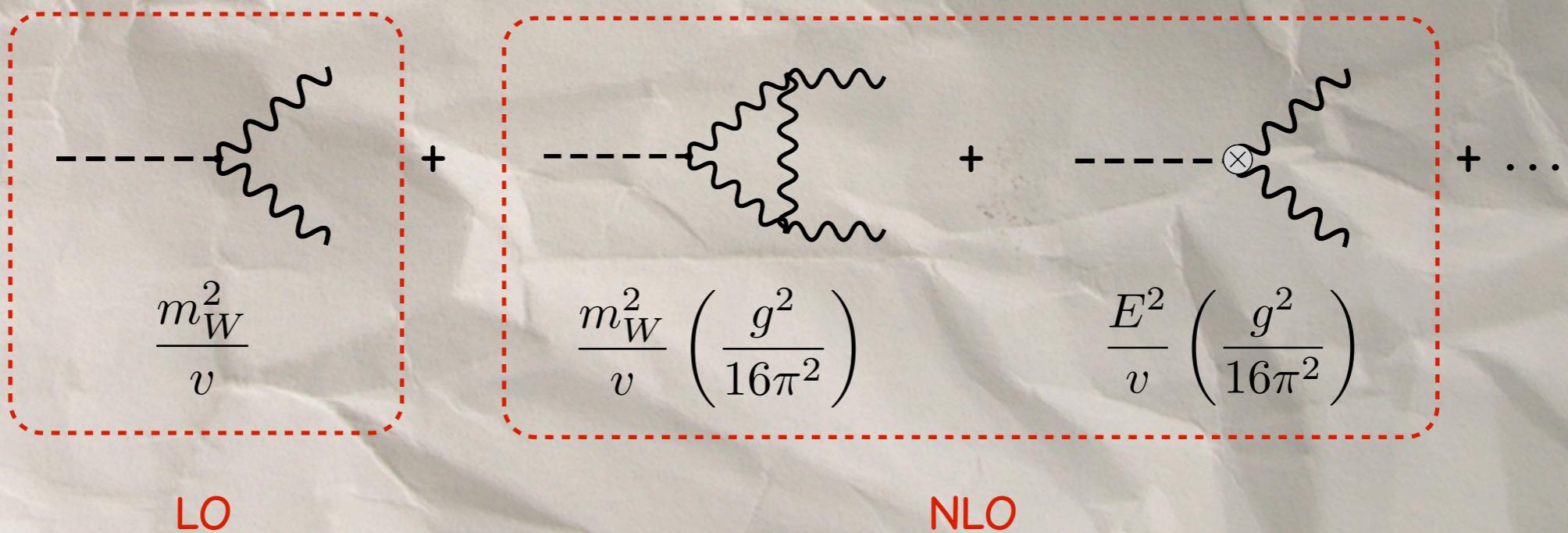
LO = tree-level $O(p^2)$

NLO = 1-loop $O(p^2)$ + tree-level $O(p^4)$

NNLO = 2-loops $O(p^2)$ + 1-loop $O(p^4)$ + tree-level $O(p^6)$

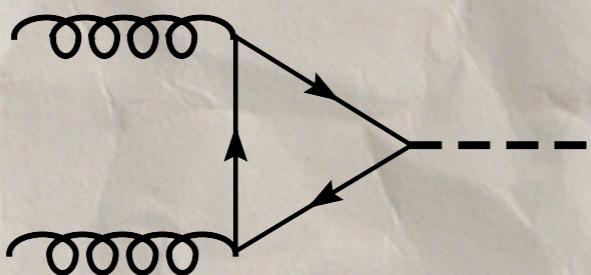
⋮
⋮

Ex: $h \rightarrow WW$

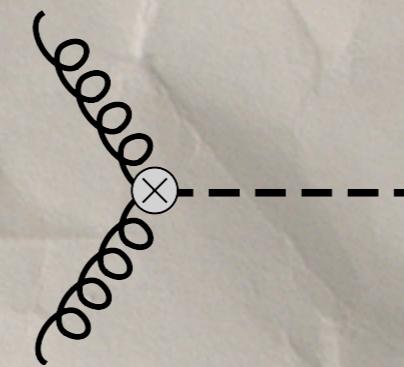


Single Higgs production via gluon fusion

in models with partial compositeness and pNG Higgs



+

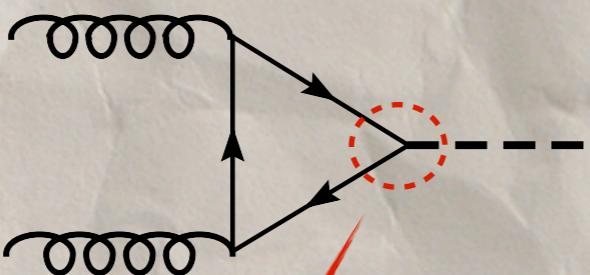


$$c_t \frac{m_h^2}{v} \frac{g_s^2}{16\pi^2} F \left(\frac{m_h^2}{m_t^2} \right)$$

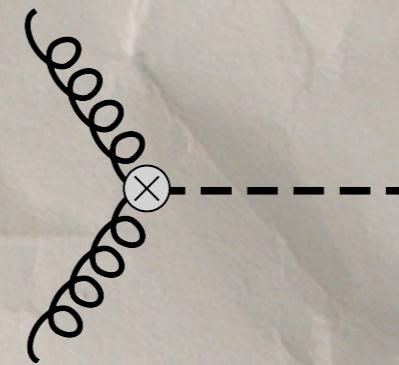
$$c_{gg} \frac{m_h^2}{v} \frac{g_s^2}{16\pi^2}$$

Single Higgs production via gluon fusion

in models with partial compositeness and pNG Higgs



+



$$c_t \frac{m_h^2}{v} \frac{g_s^2}{16\pi^2} F\left(\frac{m_h^2}{m_t^2}\right)$$

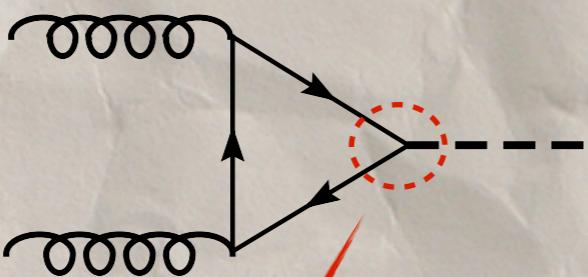
$$c_{gg} \frac{m_h^2}{v} \frac{g_s^2}{16\pi^2}$$

$c_t = 1 + O(\xi) + O\left[\left(\frac{g_*^2 v^2}{M^2}\right) \left(\frac{\lambda^2}{g_*^2}\right)\right]$

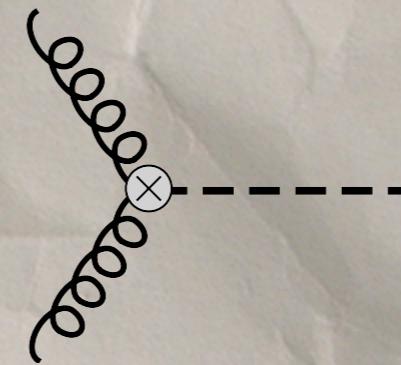


Single Higgs production via gluon fusion

in models with partial compositeness and pNG Higgs



+



$$c_t \frac{m_h^2}{v} \frac{g_s^2}{16\pi^2} F\left(\frac{m_h^2}{m_t^2}\right)$$

$$c_{gg} \frac{m_h^2}{v} \frac{g_s^2}{16\pi^2}$$

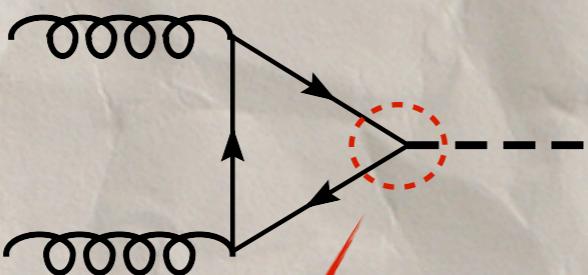


$$c_t = 1 + O(\xi) + O\left[\left(\frac{g_*^2 v^2}{M^2}\right) \left(\frac{\lambda^2}{g_*^2}\right)\right]$$

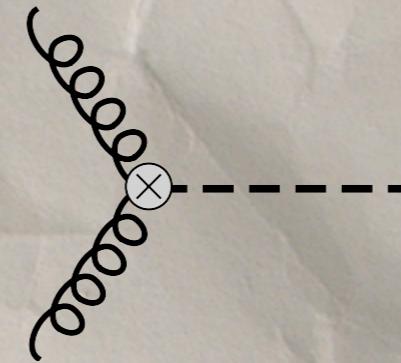
from Higgs
non-linearities

Single Higgs production via gluon fusion

in models with partial compositeness and pNG Higgs



+



$$c_t \frac{m_h^2}{v} \frac{g_s^2}{16\pi^2} F\left(\frac{m_h^2}{m_t^2}\right)$$

$$c_{gg} \frac{m_h^2}{v} \frac{g_s^2}{16\pi^2}$$

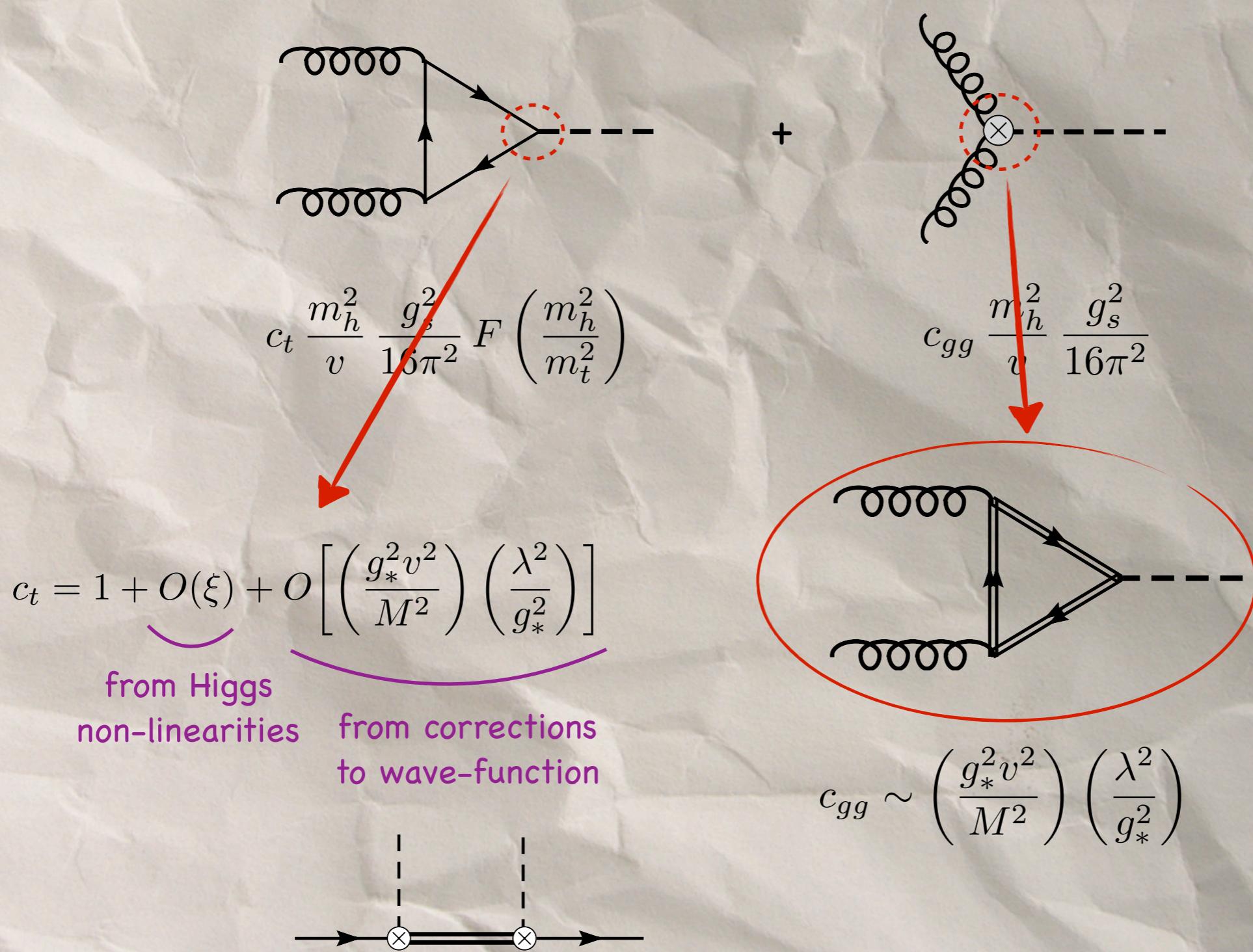
$$c_t = 1 + O(\xi) + O\left[\left(\frac{g_*^2 v^2}{M^2}\right) \left(\frac{\lambda^2}{g_*^2}\right)\right]$$

from Higgs non-linearities from corrections to wave-function



Single Higgs production via gluon fusion

in models with partial compositeness and pNG Higgs



Single Higgs production via gluon fusion

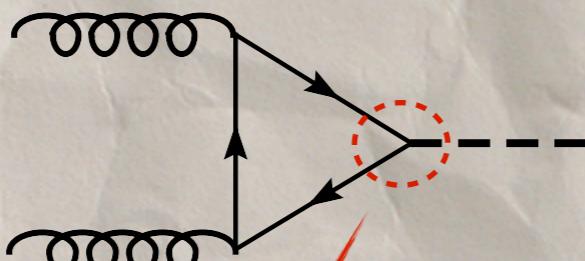
in models with partial compositeness

In minimal pNG Higgs models with partial compositeness loops of heavy fermions exactly cancel the wave function correction

Falkowski, PRD 77 (2008) 055018

Low, Rattazzi, Vichi, JHEP 1004 (2010) 126

Azatov, Galloway, PRD 85 (2012) 055013

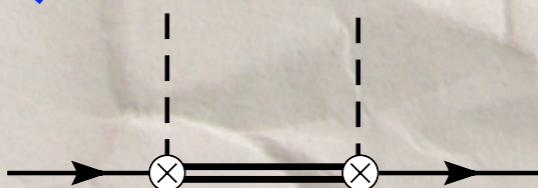


+

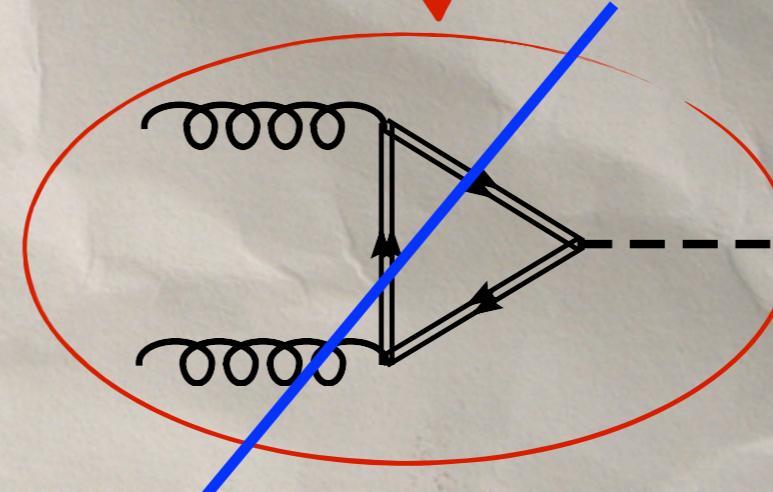
$$c_t \frac{m_h^2}{v} \frac{g_s^2}{16\pi^2} F\left(\frac{m_h^2}{m_t^2}\right)$$

$$c_t = 1 + O(\xi) + O\left[\left(\frac{g_*^2 v^2}{M^2}\right) \left(\frac{\lambda^2}{g_*^2}\right)\right]$$

from Higgs non-linearities from corrections to wave-function



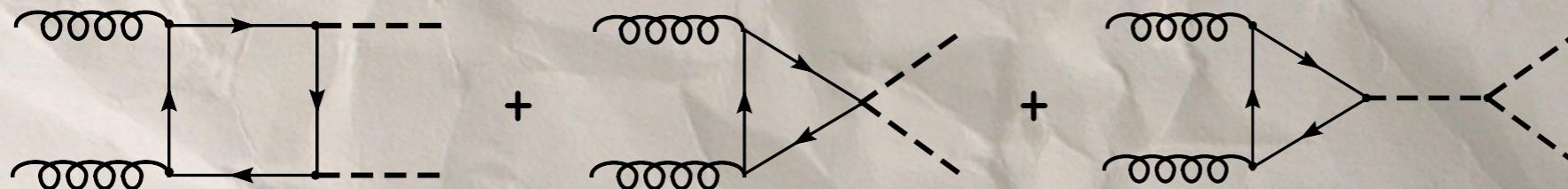
$$c_{gg} \frac{m_h^2}{v} \frac{g_s^2}{16\pi^2}$$



$$c_{gg} \sim \left(\frac{g_*^2 v^2}{M^2}\right) \left(\frac{\lambda^2}{g_*^2}\right)$$

Double Higgs production via gluon fusion

in models with partial compositeness and pNG Higgs

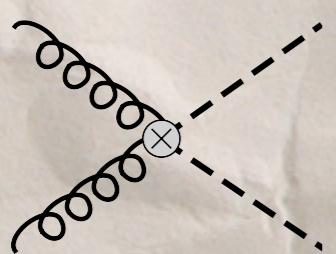


$$\frac{\hat{s}}{v^2} \frac{g_s^2}{16\pi^2} \left[c_t^2 F_{\square} \left(\frac{\hat{s}}{m_t^2} \right) + \left(c_{2t} + c d_3 \frac{m_h^2}{\hat{s}} \right) F_{\Delta} \left(\frac{\hat{s}}{m_t^2} \right) \right]$$

$$c_t = 1 + O(\xi) + O \left[\left(\frac{g_*^2 v^2}{M^2} \right) \left(\frac{\lambda^2}{g_*^2} \right) \right]$$

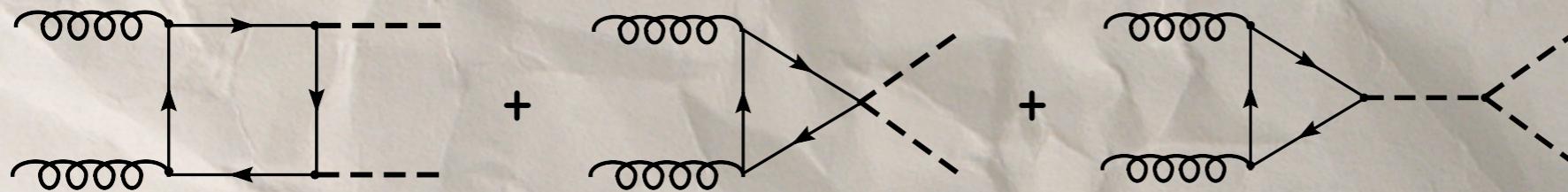
$$c_{2t} = O(\xi) + O \left[\left(\frac{g_*^2 v^2}{M^2} \right) \left(\frac{\lambda^2}{g_*^2} \right) \right]$$

$$\frac{\hat{s}}{v^2} \frac{g_s^2}{16\pi^2} \left(c_{2gg} + \frac{\hat{s}}{\Lambda^2} c_{hhgg} + \dots \right)$$



Double Higgs production via gluon fusion

in models with partial compositeness and pNG Higgs

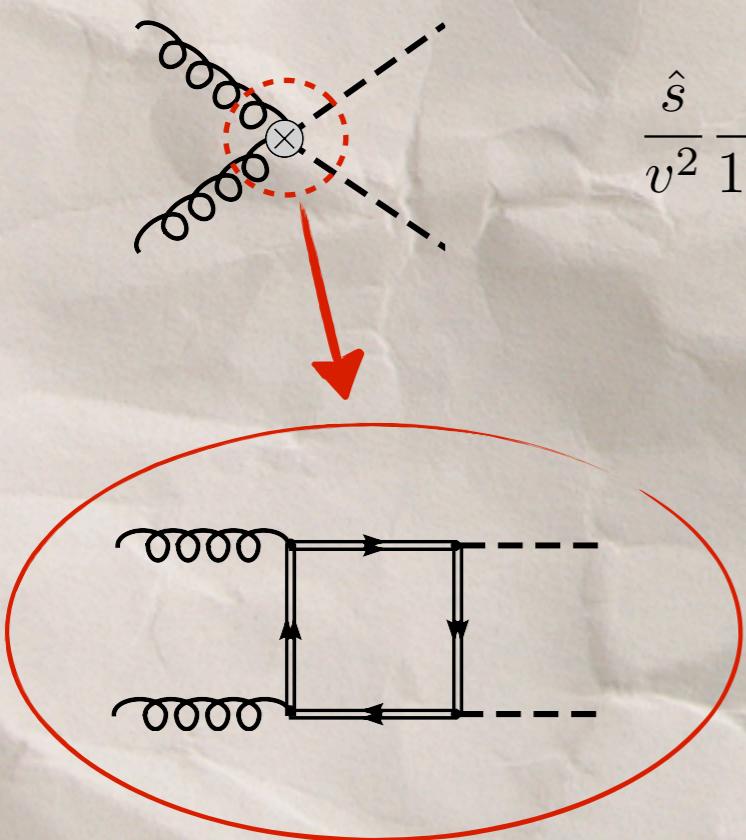


$$\frac{\hat{s}}{v^2} \frac{g_s^2}{16\pi^2} \left[c_t^2 F_{\square} \left(\frac{\hat{s}}{m_t^2} \right) + \left(c_{2t} + c d_3 \frac{m_h^2}{\hat{s}} \right) F_{\Delta} \left(\frac{\hat{s}}{m_t^2} \right) \right]$$

$$c_t = 1 + O(\xi) + O \left[\left(\frac{g_*^2 v^2}{M^2} \right) \left(\frac{\lambda^2}{g_*^2} \right) \right]$$

$$c_{2t} = O(\xi) + O \left[\left(\frac{g_*^2 v^2}{M^2} \right) \left(\frac{\lambda^2}{g_*^2} \right) \right]$$

$$\frac{\hat{s}}{v^2} \frac{g_s^2}{16\pi^2} \left(c_{2gg} + \frac{\hat{s}}{\Lambda^2} c_{hhgg} + \dots \right)$$

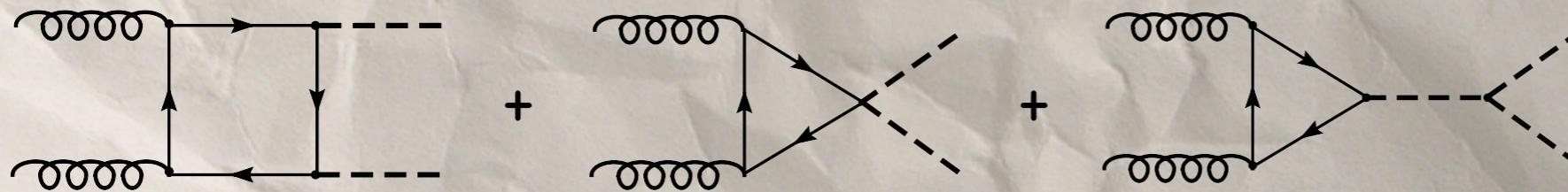


$$c_{2gg} \sim \left(\frac{g_*^2 v^2}{M^2} \right) \left(\frac{\lambda^2}{g_*^2} \right)$$

$$c_{hhgg} \sim \left(\frac{g_*^2 v^2}{M^2} \right)$$

Double Higgs production via gluon fusion

in models with partial compositeness and pNG Higgs



$$\frac{\hat{s}}{v^2} \frac{g_s^2}{16\pi^2} \left[c_t^2 F_{\square} \left(\frac{\hat{s}}{m_t^2} \right) + \left(c_{2t} + c d_3 \frac{m_h^2}{\hat{s}} \right) F_{\Delta} \left(\frac{\hat{s}}{m_t^2} \right) \right]$$

$$c_t = 1 + O(\xi) + O \left[\left(\frac{g_*^2 v^2}{M^2} \right) \left(\frac{\lambda^2}{g_*^2} \right) \right]$$

$$c_{2t} = O(\xi) + O \left[\left(\frac{g_*^2 v^2}{M^2} \right) \left(\frac{\lambda^2}{g_*^2} \right) \right]$$

$$\frac{\hat{s}}{v^2} \frac{g_s^2}{16\pi^2} \left(c_{2gg} + \frac{\hat{s}}{\Lambda^2} c_{hhgg} + \dots \right)$$

$$c_{2gg} \sim \left(\frac{g_*^2 v^2}{M^2} \right) \left(\frac{\lambda^2}{g_*^2} \right)$$

$$c_{hhgg} \sim \left(\frac{g_*^2 v^2}{M^2} \right)$$

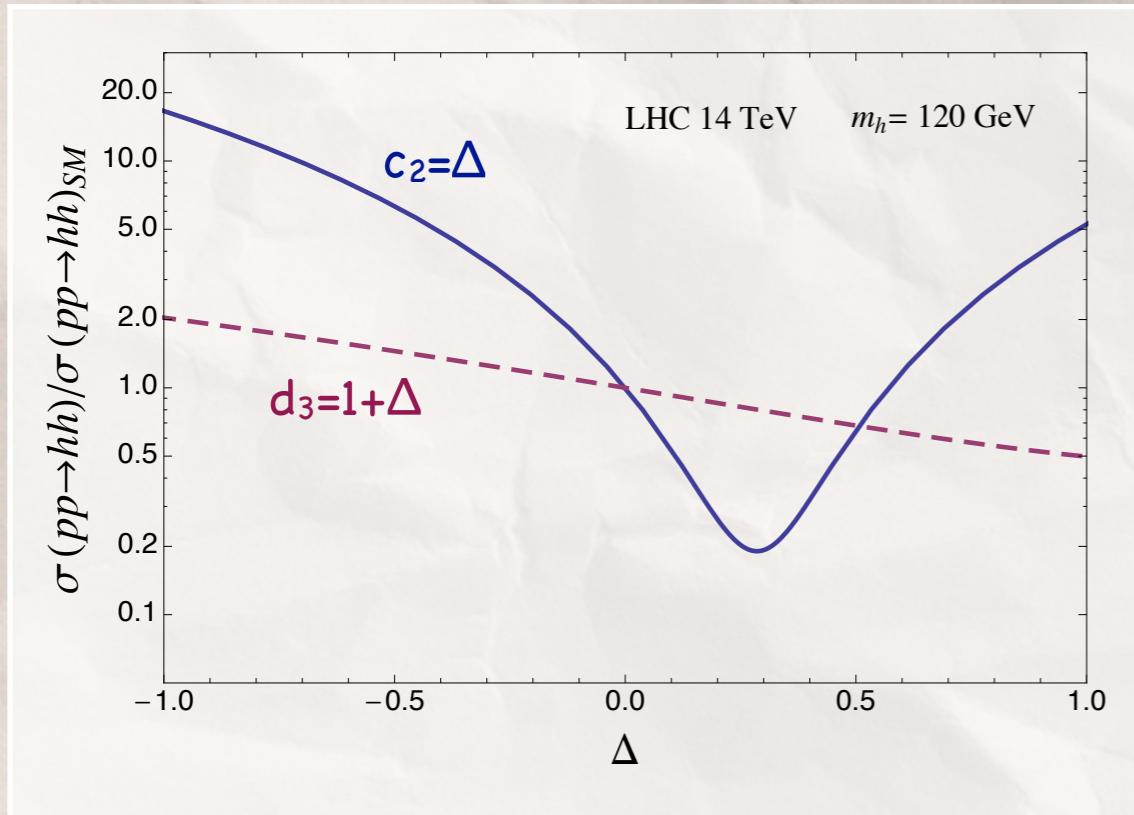
In minimal pNG Higgs models with partial compositeness loops of heavy fermions cancel the wave function correction only at zero Higgs momentum

The $O(p^6)$ terms might be numerically important

Double Higgs production via gluon fusion

- ▶ $\sigma(gg \rightarrow hh)$ much more sensitive on new $tthh$ couplings c_2 than on trilinear d_3

[First noticed by:
Dib, Rosenfeld, Zerwekh, JHEP 0605 (2006) 074
Grober and Muhlleitner, JHEP 1106 (2011) 020]



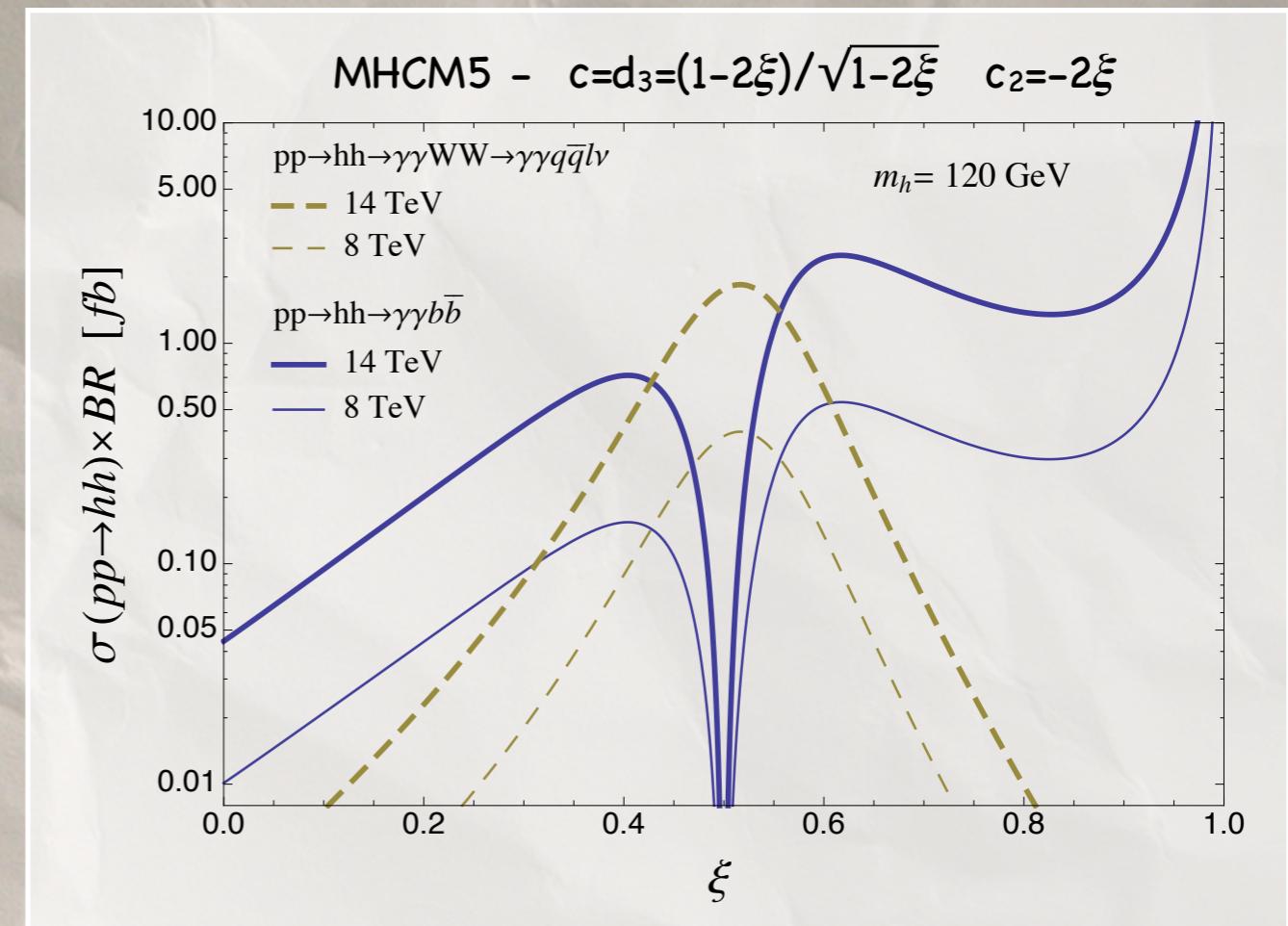
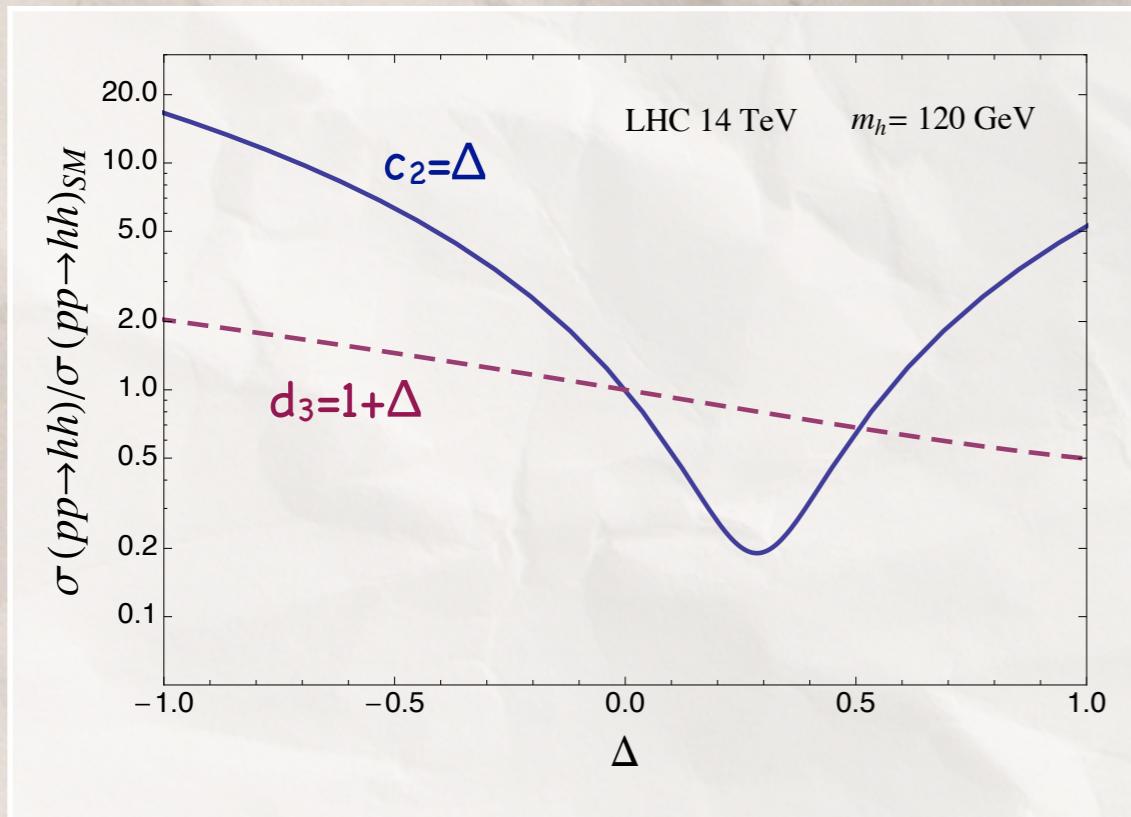
results from:

R.C., Ghezzi, Moretti, Panico, Piccinini, Wulzer,
arXiv:1205.5444

Double Higgs production via gluon fusion

- $\sigma(gg \rightarrow hh)$ much more sensitive on new tthh couplings c_2 than on trilinear d_3

[First noticed by:
 Dib, Rosenfeld, Zerwekh, JHEP 0605 (2006) 074
 Grober and Muhlleitner, JHEP 1106 (2011) 020]

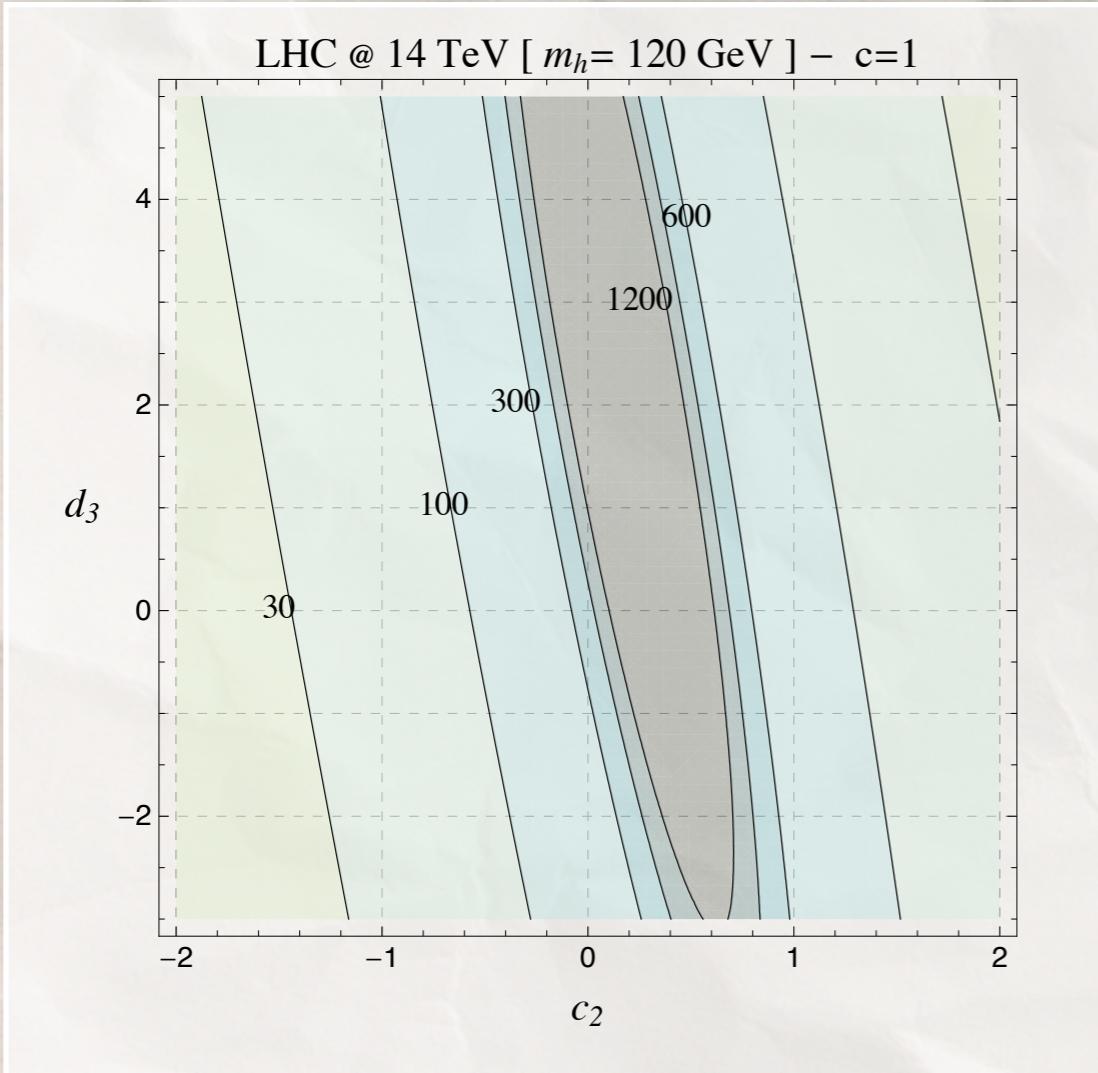


- If $BR(h) \simeq BR(h)_{SM}$ best channel is $hh \rightarrow bb\gamma\gamma$
 [Baur, Plehn, Rainwater, PRD 69 (2004) 053004]
- $\xi = 0.15 \rightarrow \sigma(gg \rightarrow hh) \times BR \sim 3 [\sigma(gg \rightarrow hh) \times BR]_{SM}$
- If $c \simeq 0$ (fermiophobic Higgs) a very promising channel is $hh \rightarrow WW\gamma\gamma \rightarrow Wqq\bar{l}\nu\gamma\gamma$
 $\xi = 0.5$ ($c=0$), $\sqrt{s}=8\text{TeV} \rightarrow \sigma(gg \rightarrow hh) \times BR \sim 0.7\text{fb}$

results from:
 R.C., Ghezzi, Moretti, Panico, Piccinini, Wulzer,
 arXiv:1205.5444

Extracting c_2 from $gg \rightarrow hh \rightarrow bb\gamma\gamma$

Discovery luminosity (fb^{-1})



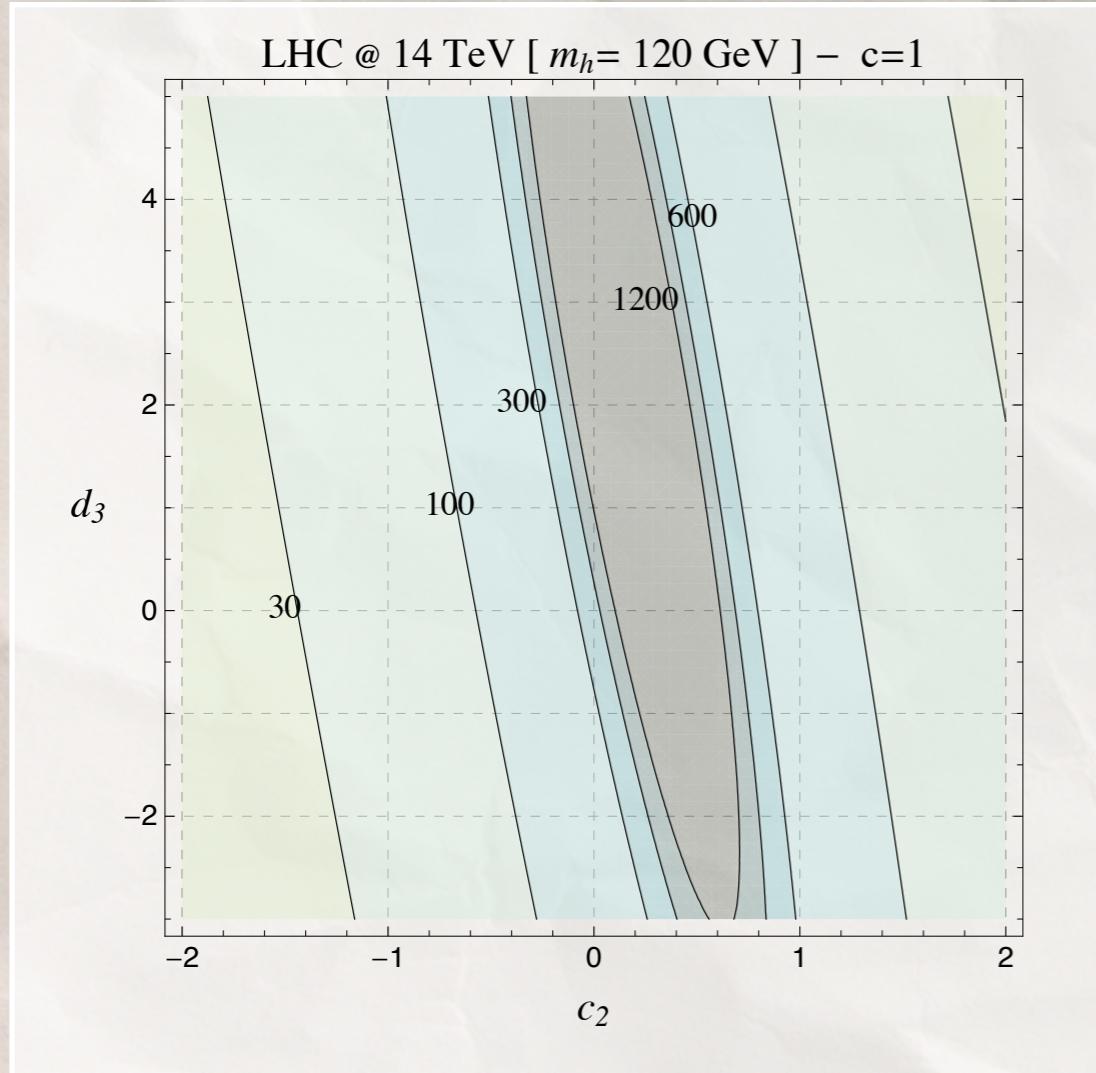
Ex: $\sqrt{s}=14 \text{ TeV}$ $L=300 \text{ fb}^{-1} \rightarrow c_2 > 0.8 \quad c_2 < -0.2$

results from:

R.C., Ghezzi, Moretti, Panico, Piccinini, Wulzer,
arXiv:1205.5444

Extracting c_2 from $gg \rightarrow hh \rightarrow bb\gamma\gamma$

Discovery luminosity (fb^{-1})

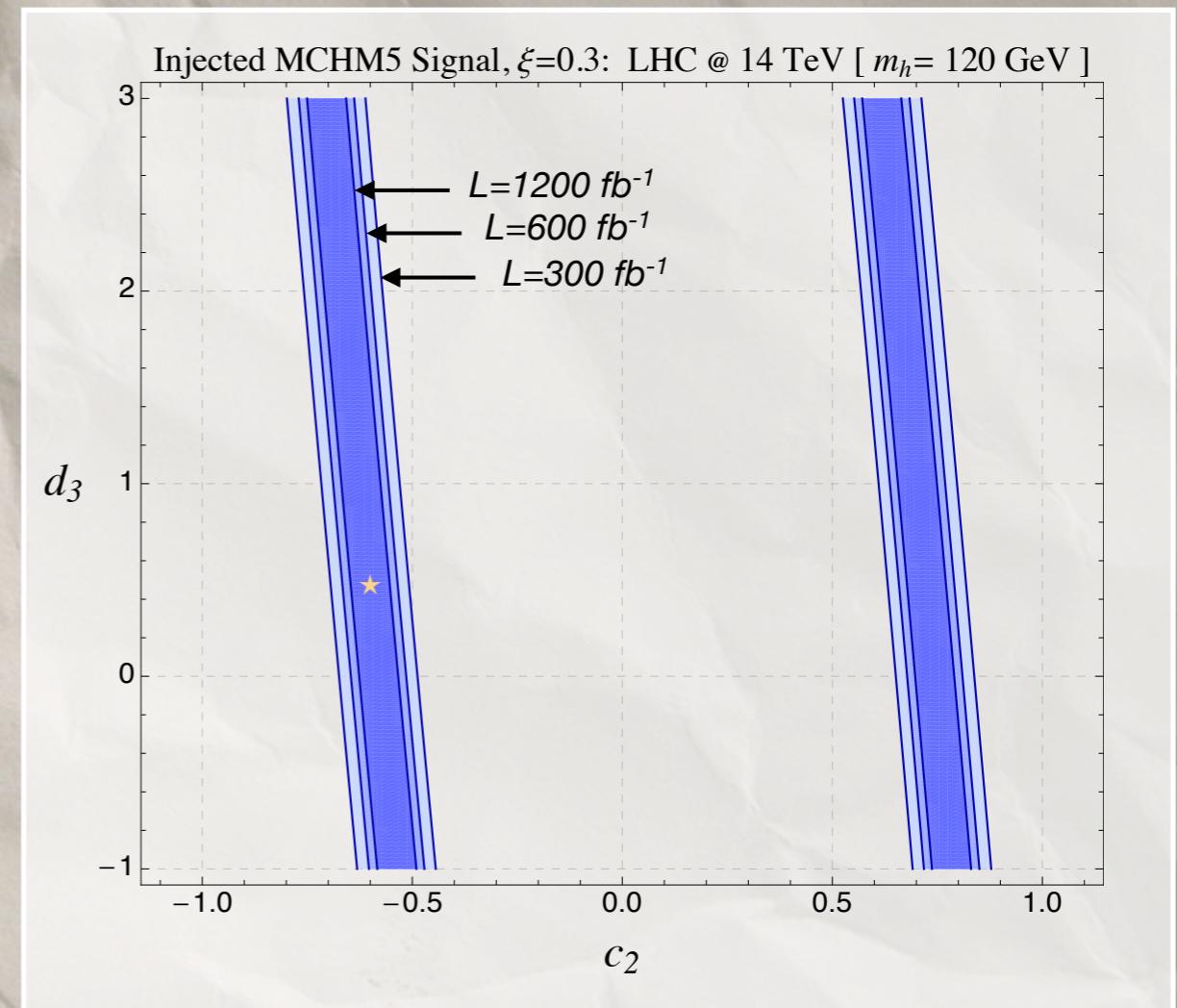


Ex: $\sqrt{s}=14\text{TeV}$ $L=300\text{fb}^{-1}$ $\rightarrow c_2 > 0.8 \quad c_2 < -0.2$

results from:

R.C., Ghezzi, Moretti, Panico, Piccinini, Wulzer,
arXiv:1205.5444

Precision on couplings (curves at 68% prob.)



Ex: Injected $\xi=0.3$ ($c=d_3=0.48 \quad c_2=-0.6$)

$\Delta c_2/c_2 = 15-20\%$

Fit of Higgs couplings with current data

- ▶ Current information made public by experimental collaborations is (often) not sufficient to allow theorists to make a rigorous fit

Need:

- [1] cut efficiencies for each Higgs production mode and event category
- [2] likelihoods functions for each event category

Fit of Higgs couplings with current data

- ▶ Current information made public by experimental collaborations is (often) not sufficient to allow theorists to make a rigorous fit

Need:

- [1] cut efficiencies for each Higgs production mode and event category
- [2] likelihoods functions for each event category

- ▶ Even with full info at disposal one should combine channels by taking into account all correlations (systematics, theory errors, etc.)

This means: marginalizing over $O(100)$ nuisance parameters !

Fit of Higgs couplings with current data

- ▶ Current information made public by experimental collaborations is (often) not sufficient to allow theorists to make a rigorous fit

Need:

- [1] cut efficiencies for each Higgs production mode and event category
- [2] likelihoods functions for each event category

- ▶ Even with full info at disposal one should combine channels by taking into account all correlations (systematics, theory errors, etc.)

This means: marginalizing over $O(100)$ nuisance parameters !

- ▶ Best if fit is done by experimentalists; theorists can give support on how to perform calculations (with chiral Lagrangian)

Fit of Higgs couplings with current data

- ▶ For the impatient: I'll show a fit performed with reasonable simplifying assumptions.

We concentrate on leading effects and keep only **two parameters: (a,c)**

results from: Azatov, R.C., Galloway, JHEP 1204 (2012) 127

Fit of Higgs couplings with current data

- ▶ For the impatient: I'll show a fit performed with reasonable simplifying assumptions.

We concentrate on leading effects and keep only **two parameters: (a,c)**

results from: Azatov, R.C., Galloway, JHEP 1204 (2012) 127



*see talks by Azatov for
more details*

Fit of Higgs couplings with current data

- ▶ For the impatient: I'll show a fit performed with reasonable simplifying assumptions.

We concentrate on leading effects and keep only **two parameters: (a,c)**

results from: Azatov, R.C., Galloway, JHEP 1204 (2012) 127



see talks by Azatov for more details

- ▶ Similar fits done by several other groups:

Carmi, Falkowski, Kuflik, Volansky, arXiv:1202.3144

Espinosa, Grojean, Muhlleitner, Trott, arXiv:1202.3697

Giardino, Kannike, Raidal, Strumia, arXiv:1203.4254

Ellis and You, arXiv:1204.0464

Farina, Grojean, Salvioni, arXiv:1205.0011

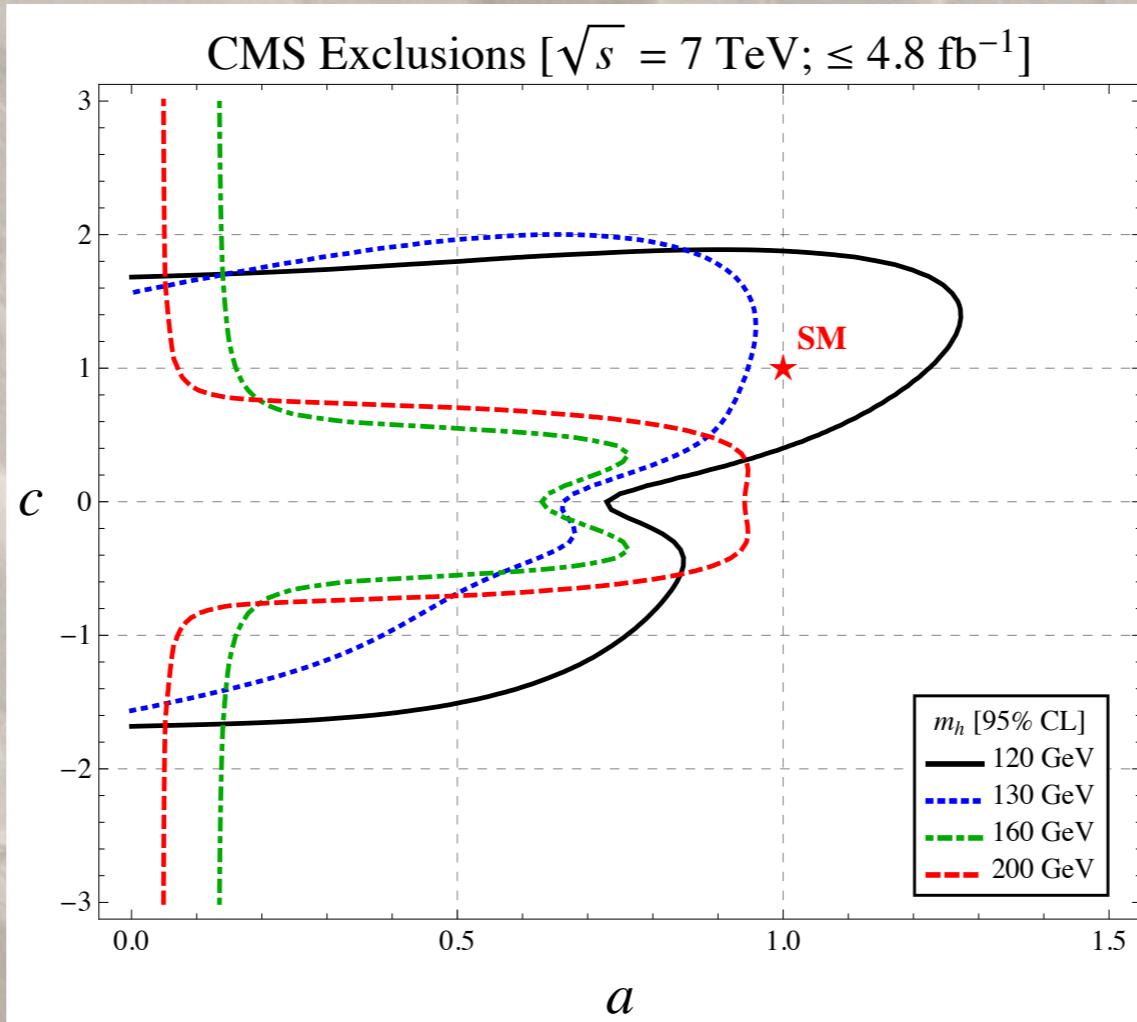
Klute, Lafaye, Plehn, Rauch, Zerwas, arXiv:1205.2699

⋮
⋮
⋮

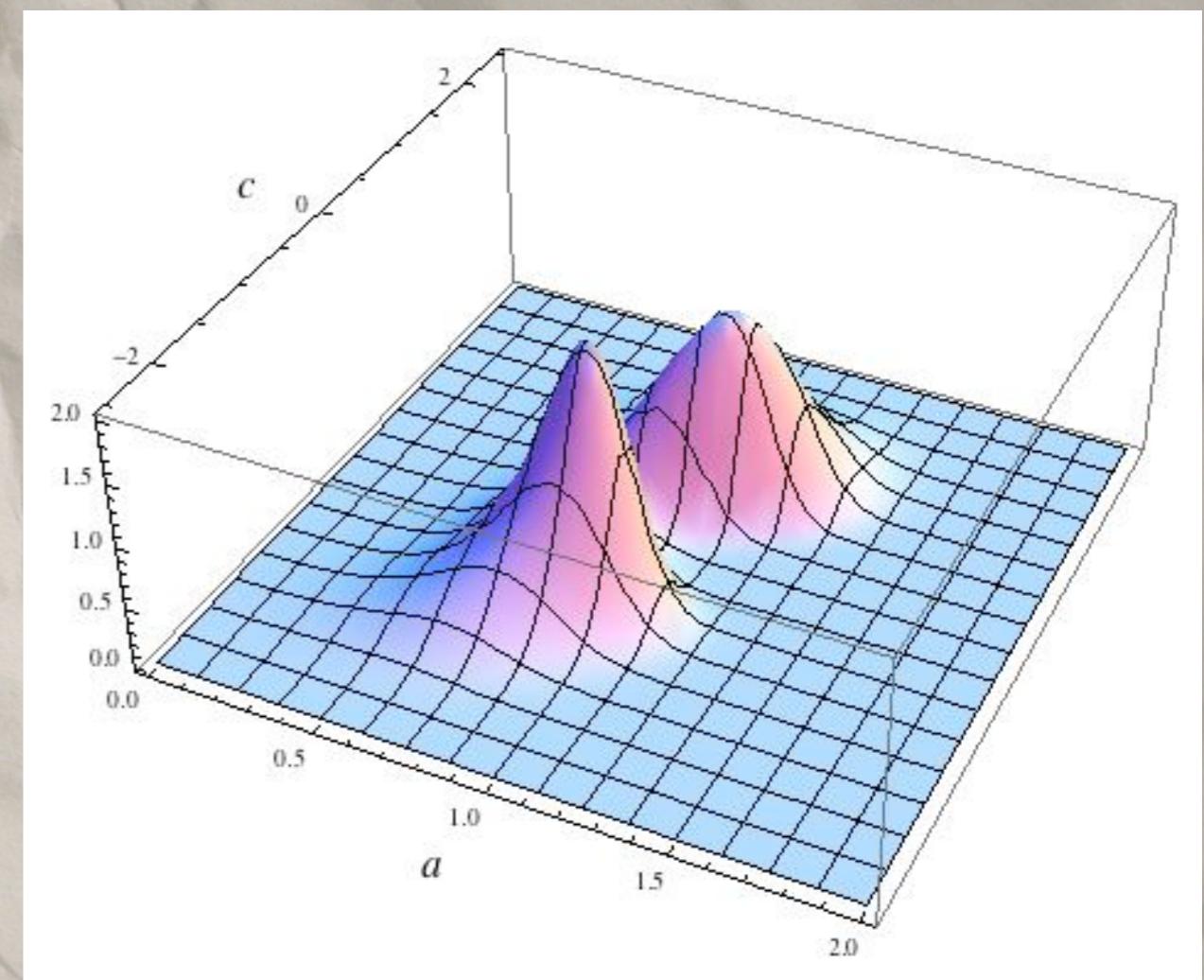
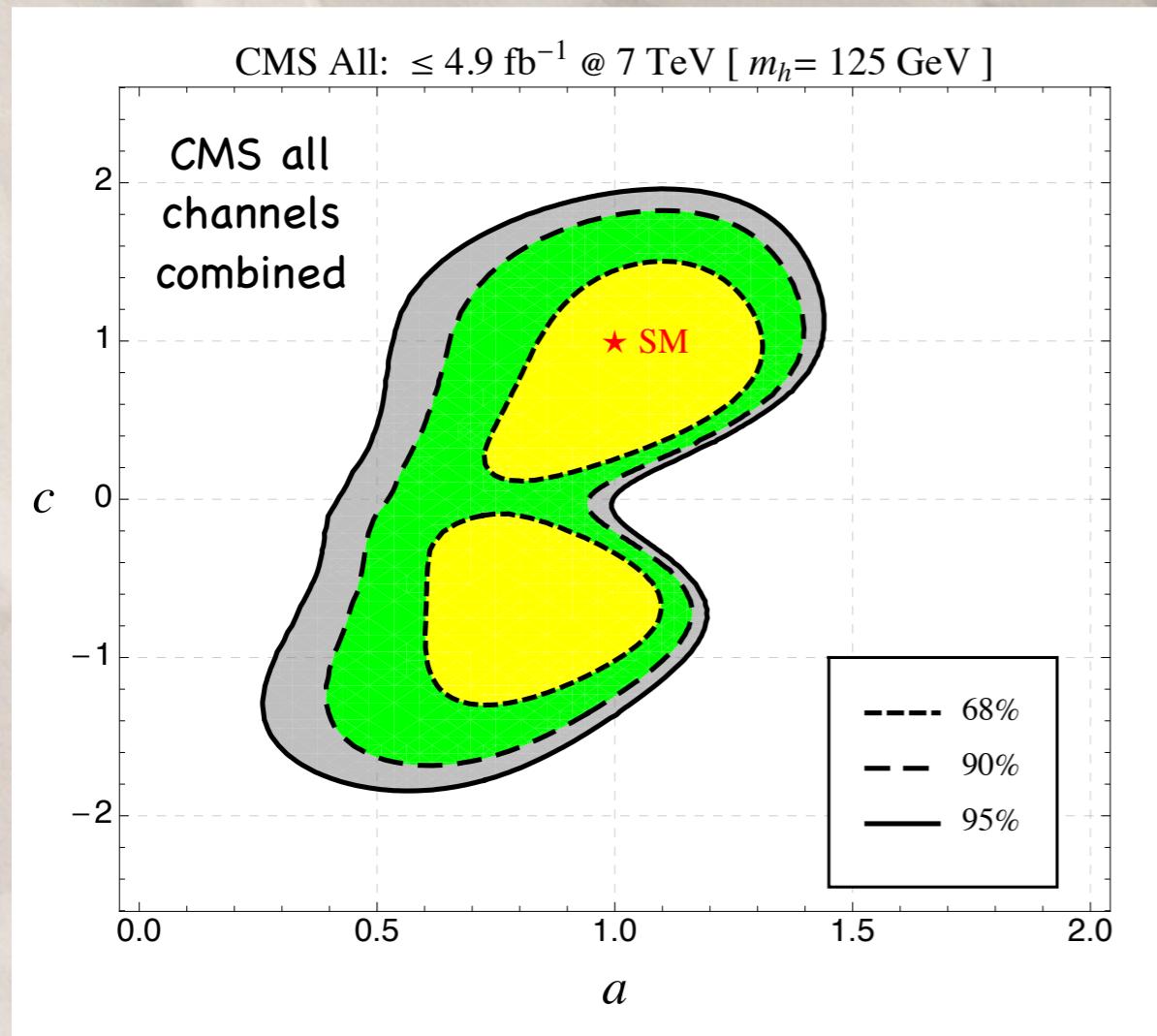


see talks by Falkowski, Grojean, Strumia

95% Exclusion for various mH

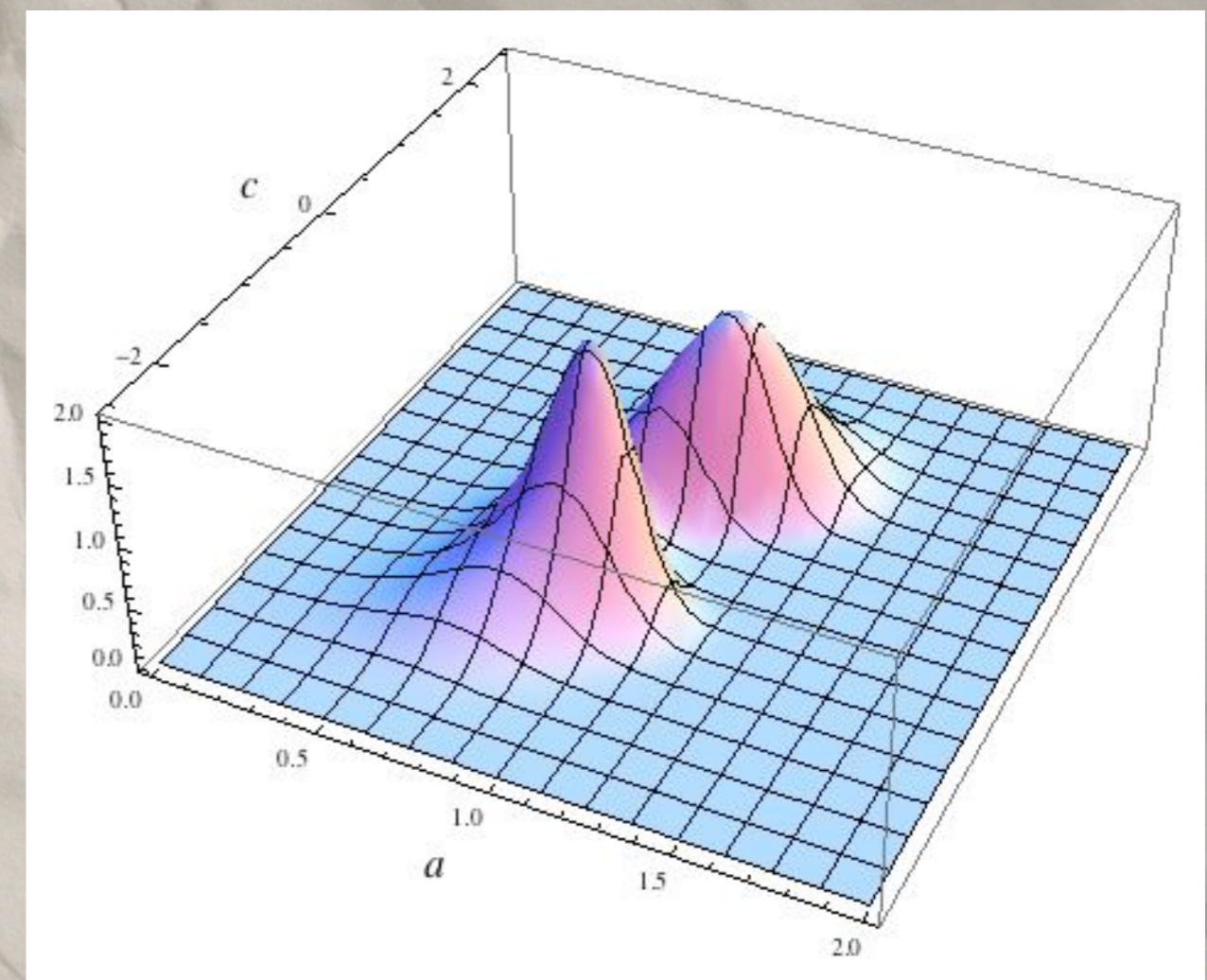
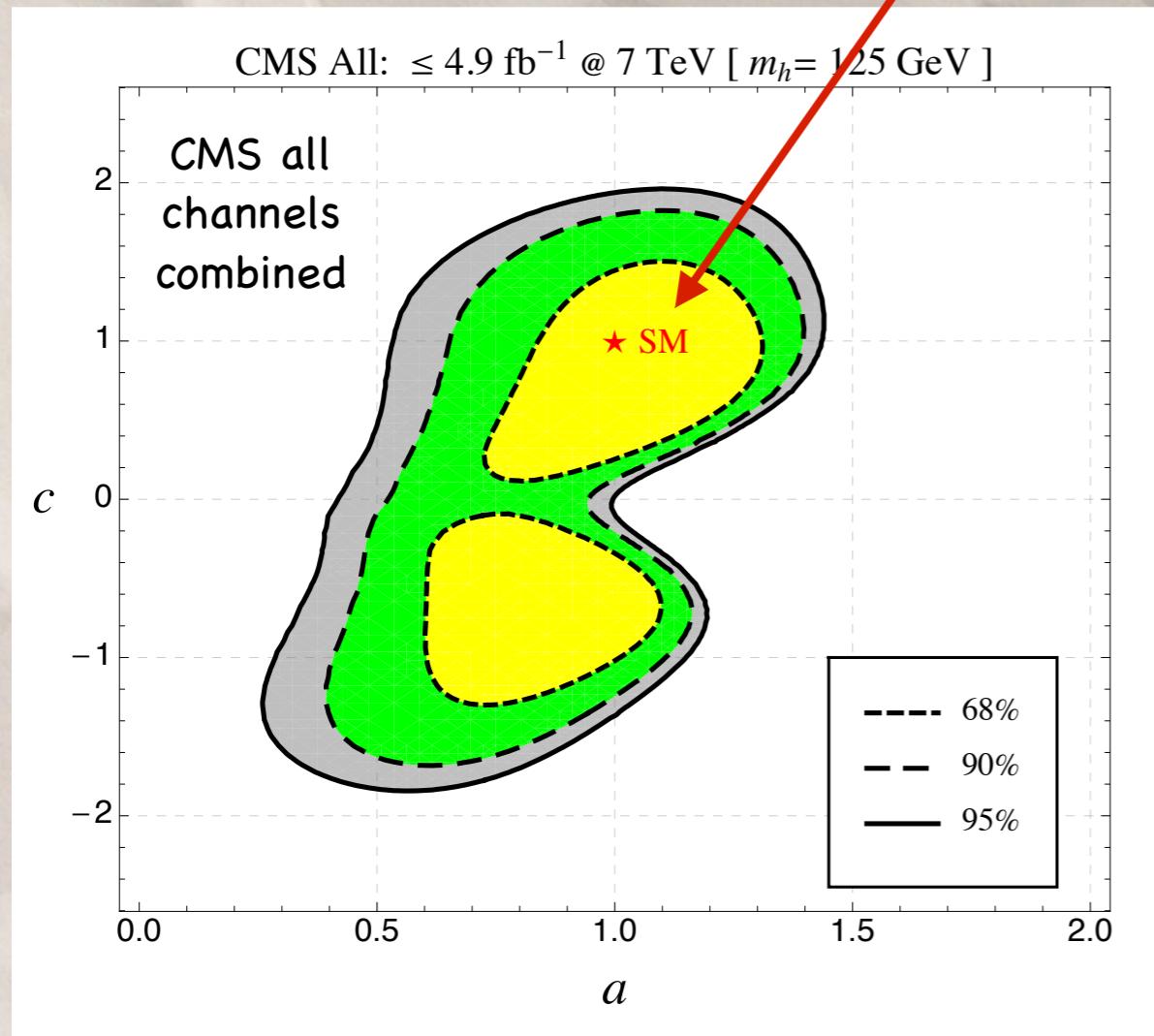


Fit at $m_H=125\text{GeV}$

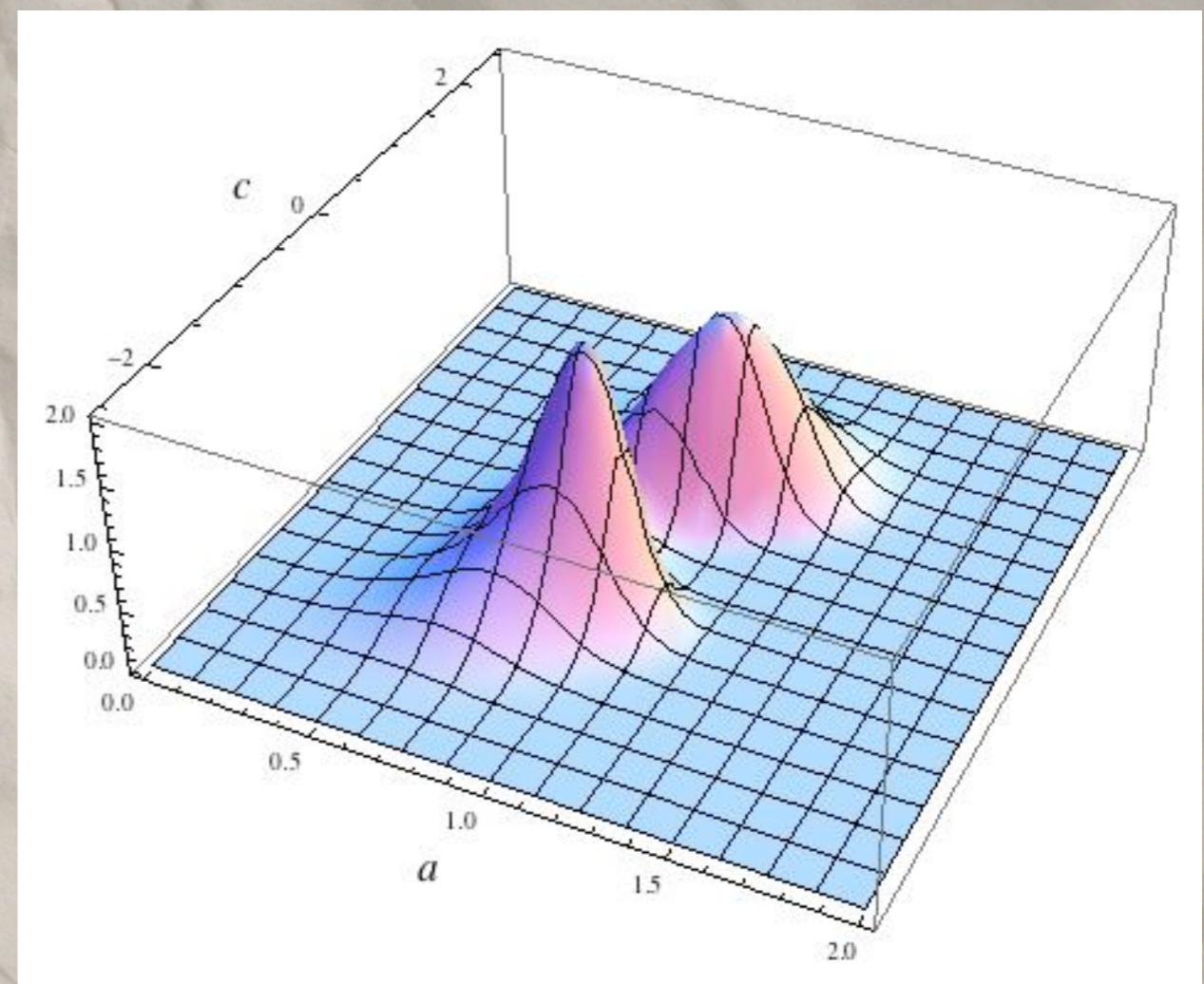
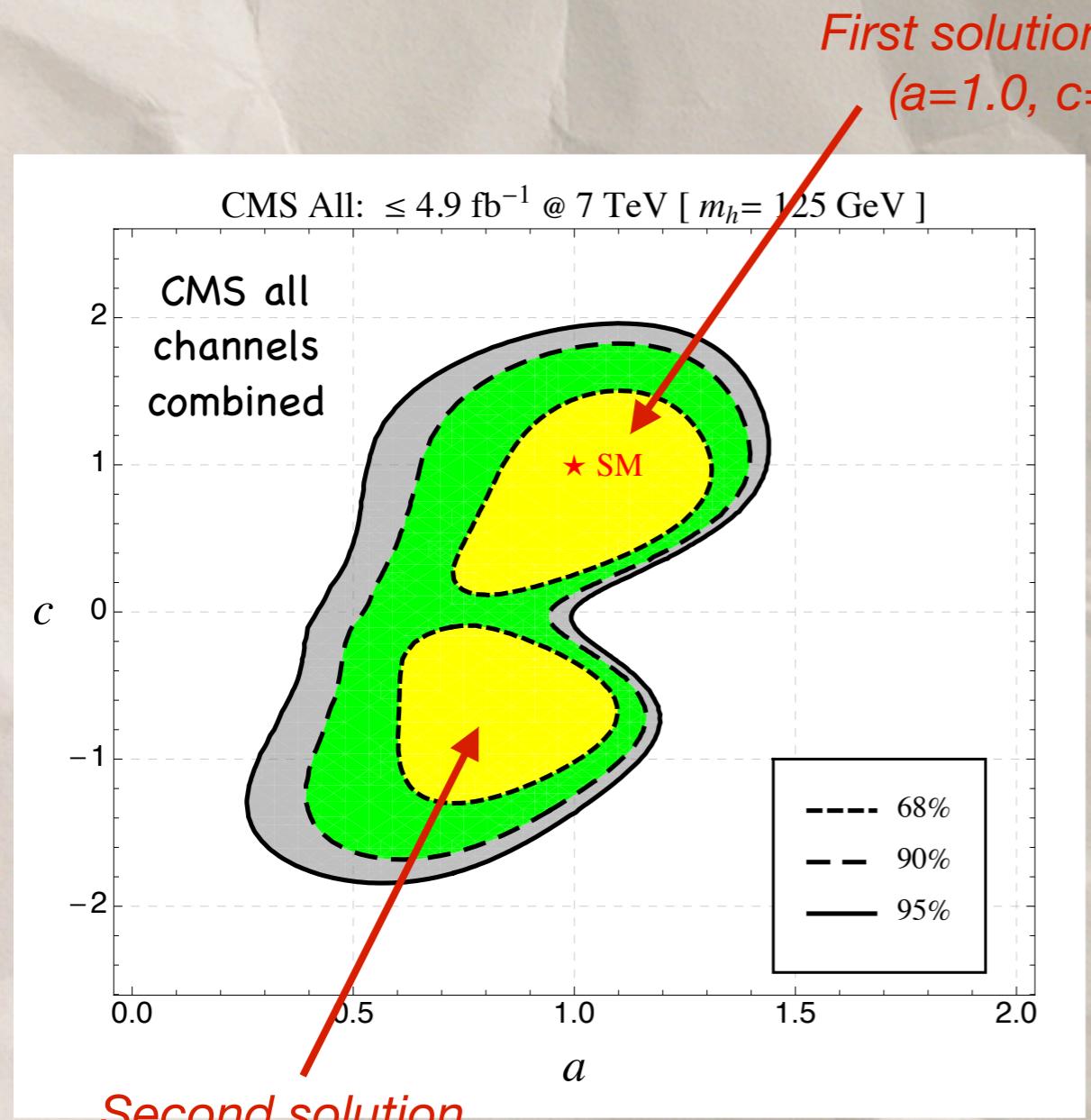


Fit at $m_H=125\text{GeV}$

*First solution SM-like
($a=1.0$, $c=0.75$)*



Fit at $m_H=125\text{GeV}$

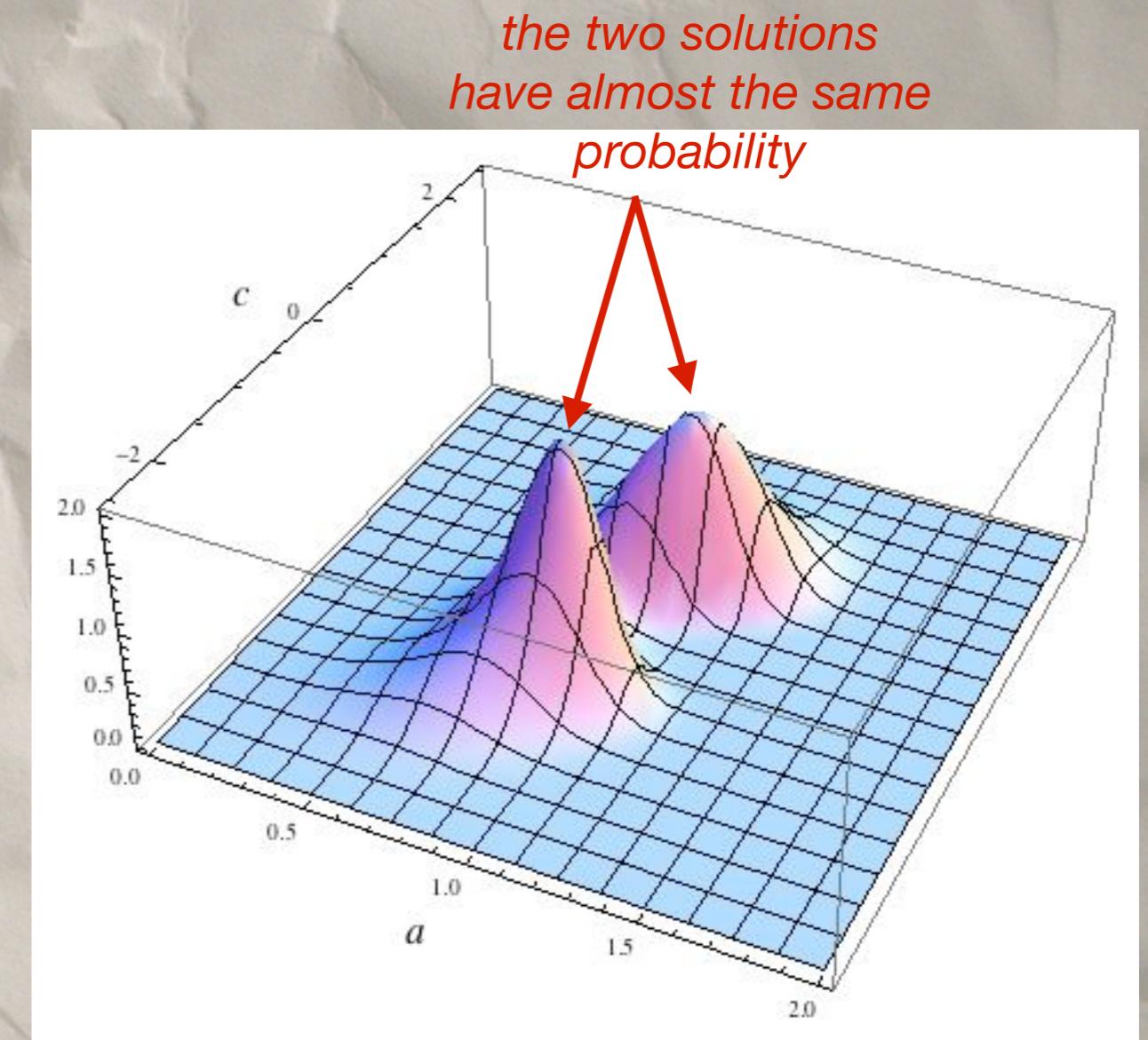
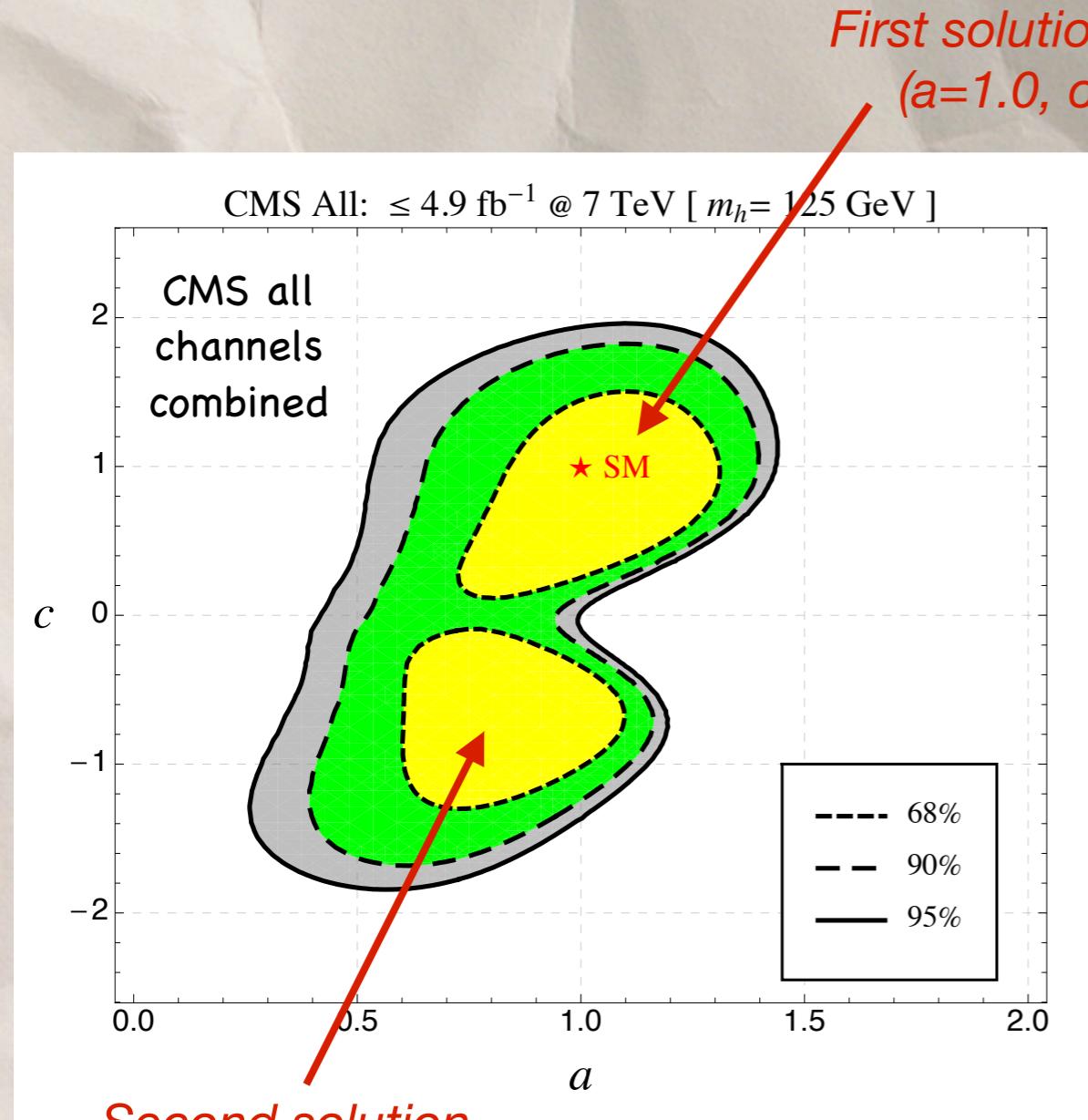


Second solution due
to degeneracy in $\gamma\gamma$:

$$a_2 \simeq + a_1 \frac{4.5 a_1 - c_1}{4.5 a_1 + c_1}$$

$$c_2 \simeq - c_1 \frac{4.5 a_1 - c_1}{4.5 a_1 + c_1}$$

Fit at $m_H=125\text{GeV}$

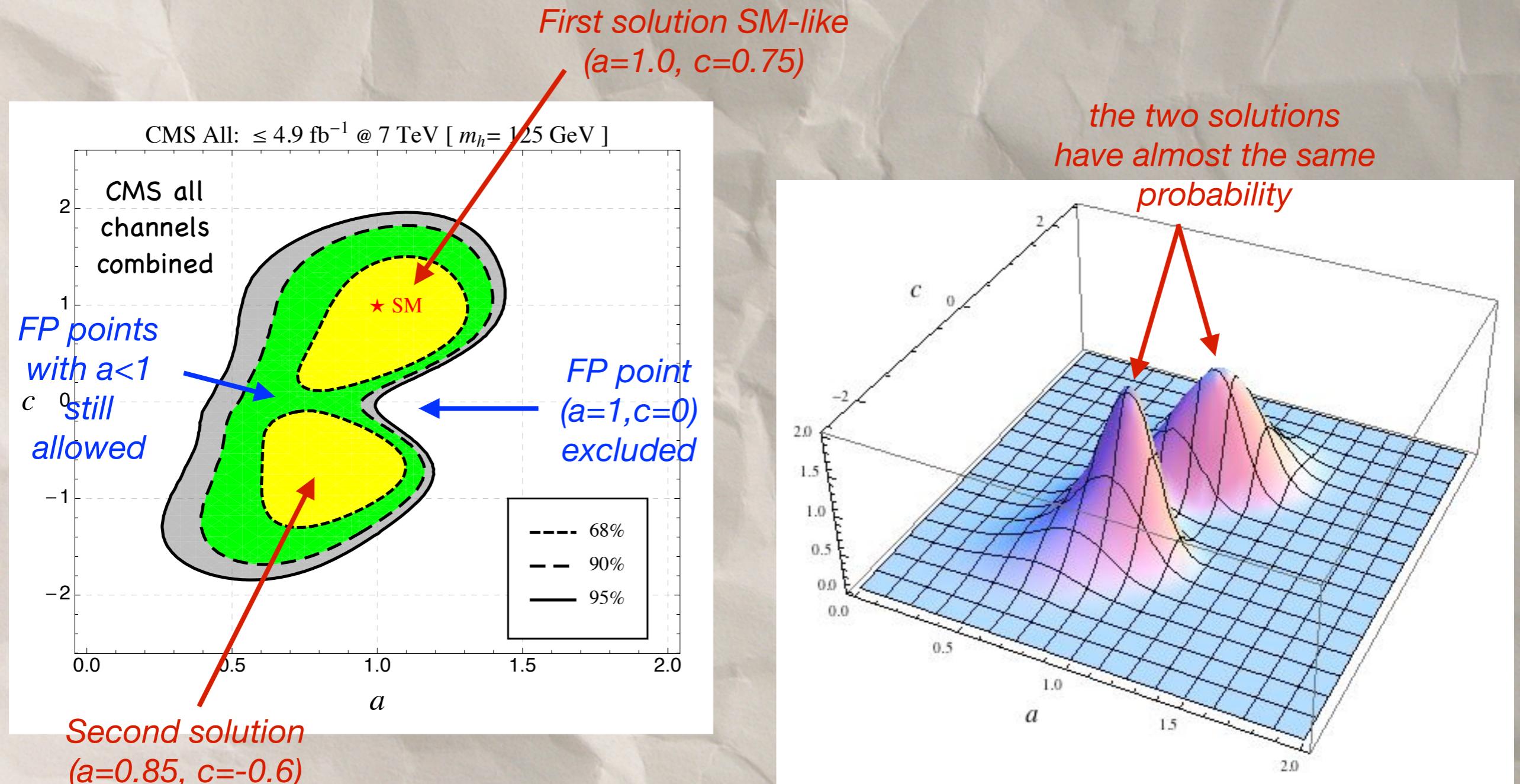


Second solution due to degeneracy in $\gamma\gamma$:

$$a_2 \simeq + a_1 \frac{4.5 a_1 - c_1}{4.5 a_1 + c_1}$$

$$c_2 \simeq - c_1 \frac{4.5 a_1 - c_1}{4.5 a_1 + c_1}$$

Fit at $m_H=125\text{GeV}$



Second solution due to degeneracy in $\gamma\gamma$:

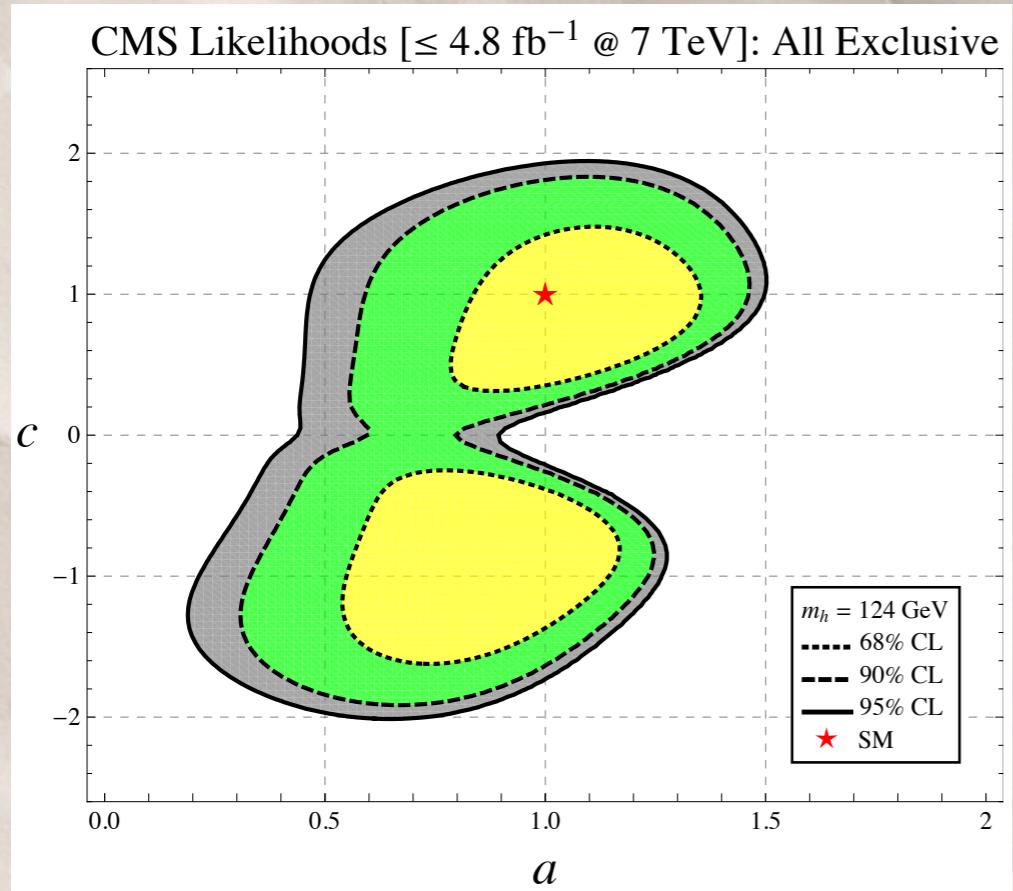
$$a_2 \simeq + a_1 \frac{4.5 a_1 - c_1}{4.5 a_1 + c_1}$$

$$c_2 \simeq - c_1 \frac{4.5 a_1 - c_1}{4.5 a_1 + c_1}$$

The importance of being “exclusive”

(from J. Galloway, talk at Pheno 2012)

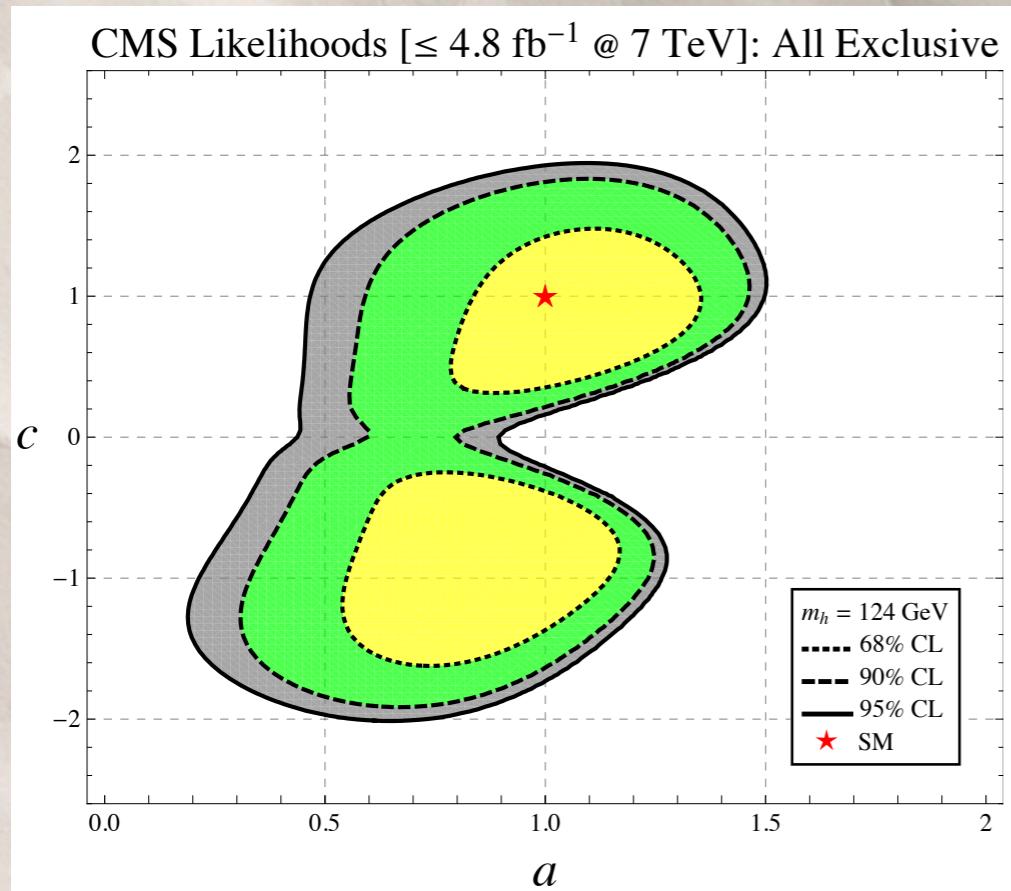
CMS all channels combined:
WW and $\gamma\gamma$ fully EXCLUSIVE



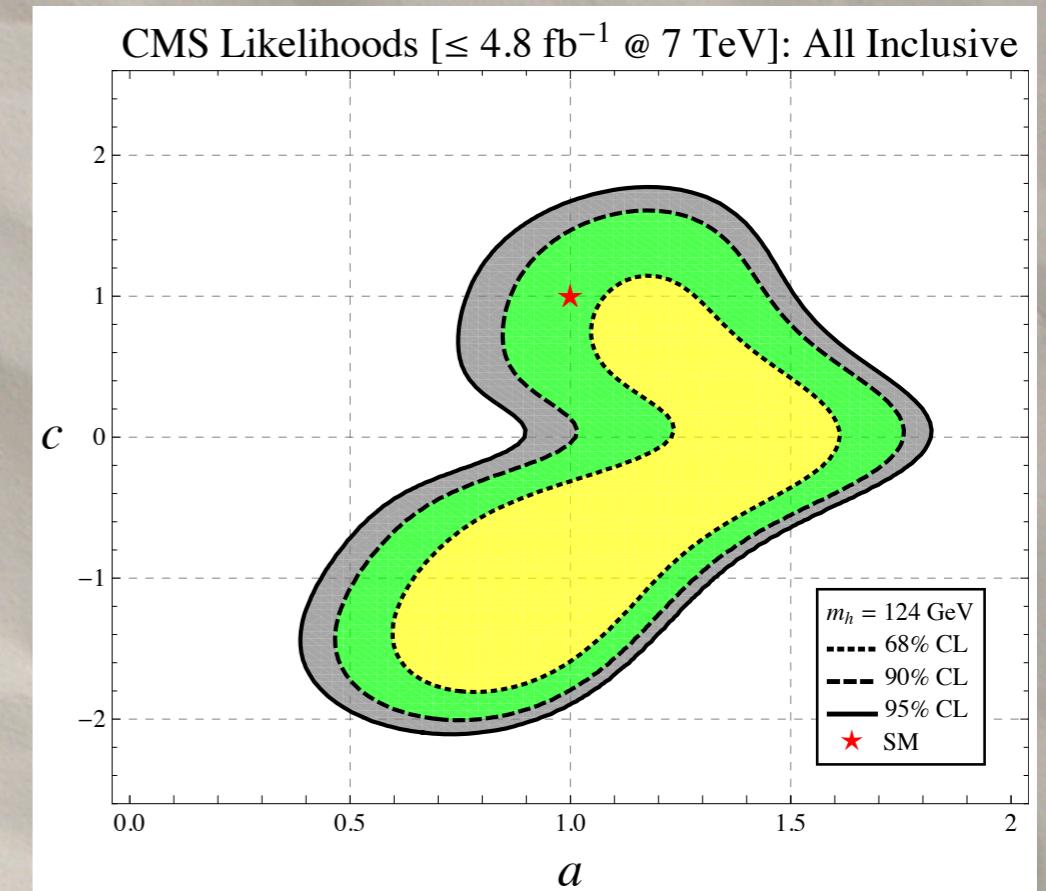
The importance of being “exclusive”

(from J. Galloway, talk at Pheno 2012)

CMS all channels combined:
WW and $\gamma\gamma$ fully EXCLUSIVE



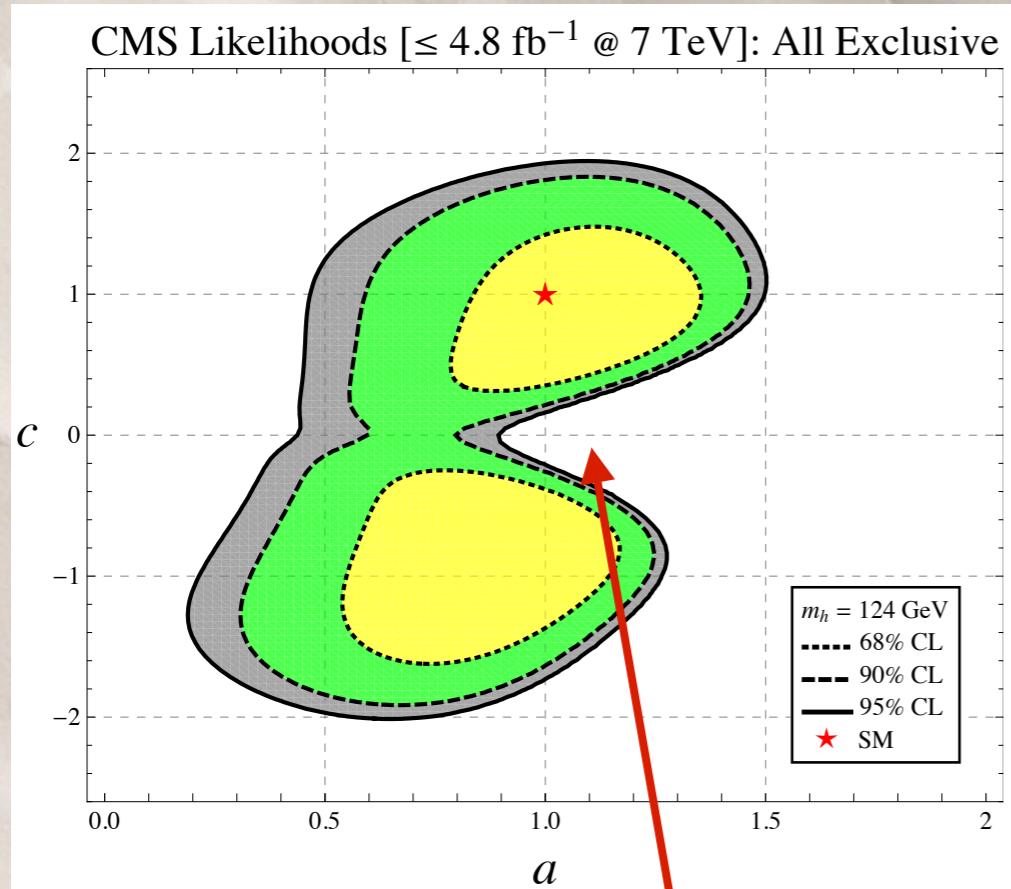
CMS all channels combined:
WW and $\gamma\gamma$ fully INCLUSIVE



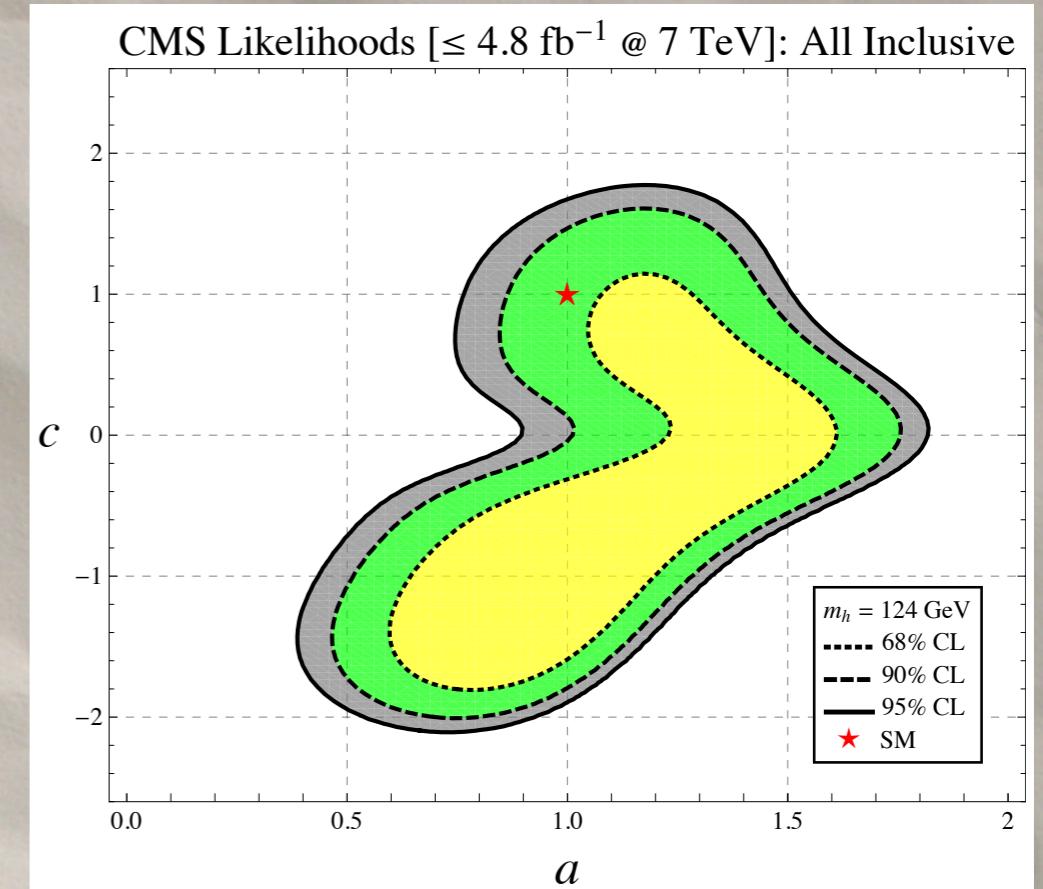
The importance of being “exclusive”

(from J. Galloway, talk at Pheno 2012)

CMS all channels combined:
WW and $\gamma\gamma$ fully EXCLUSIVE



CMS all channels combined:
WW and $\gamma\gamma$ fully INCLUSIVE

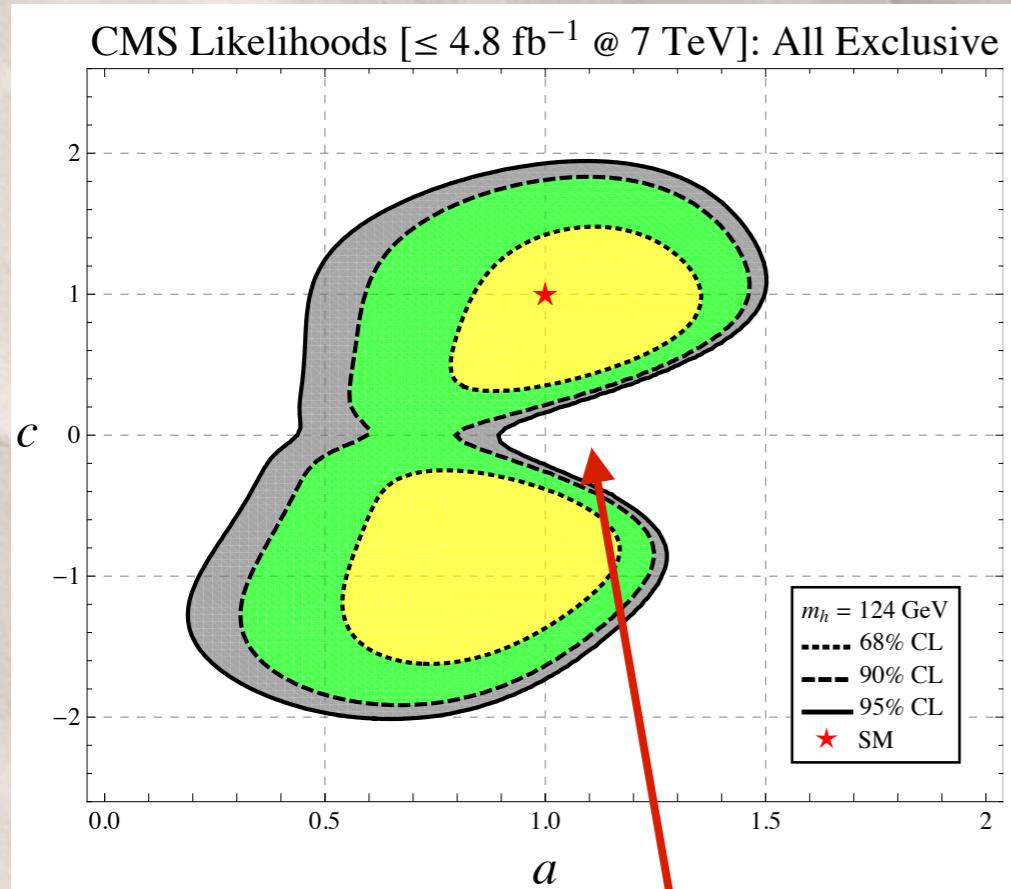


*absence of excess in $WWjj$ and
exclusive $\gamma\gamma$ analysis of CMS
rule out FP region with $a > 1$*

The importance of being “exclusive”

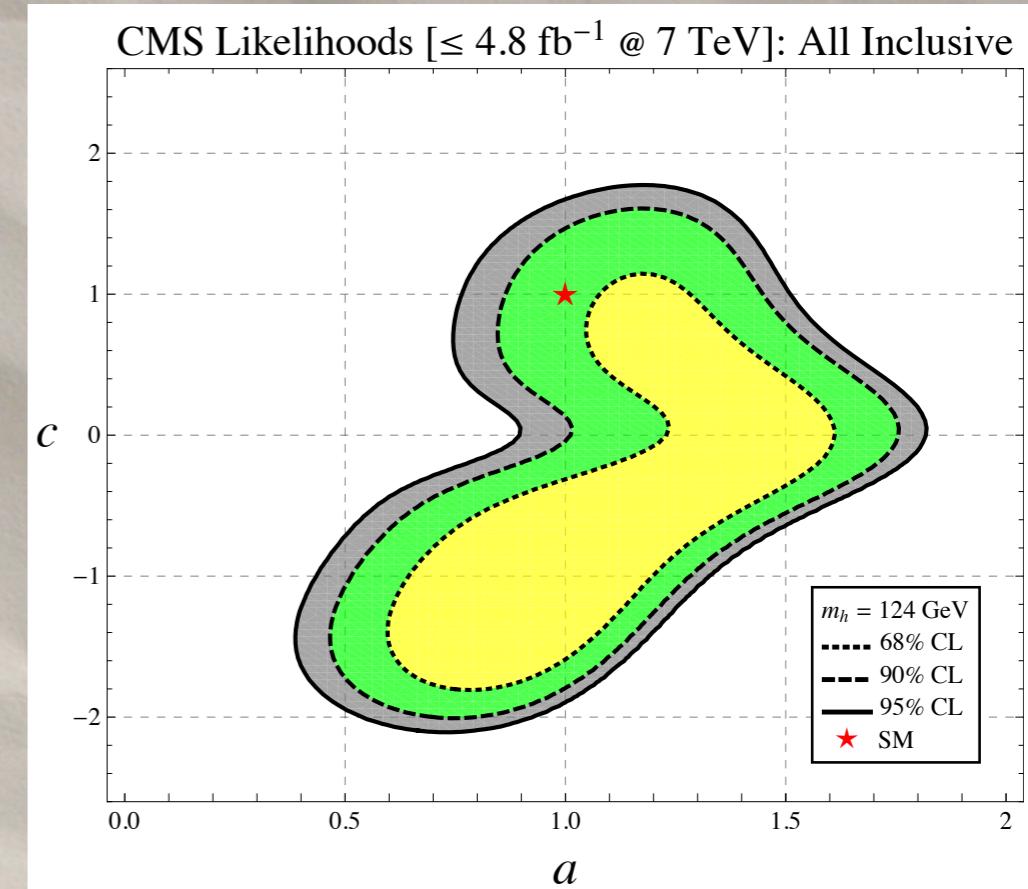
(from J. Galloway, talk at Pheno 2012)

CMS all channels combined:
WW and $\gamma\gamma$ fully EXCLUSIVE



*absence of excess in $WWjj$ and
exclusive $\gamma\gamma$ analysis of CMS
rule out FP region with $a>1$*

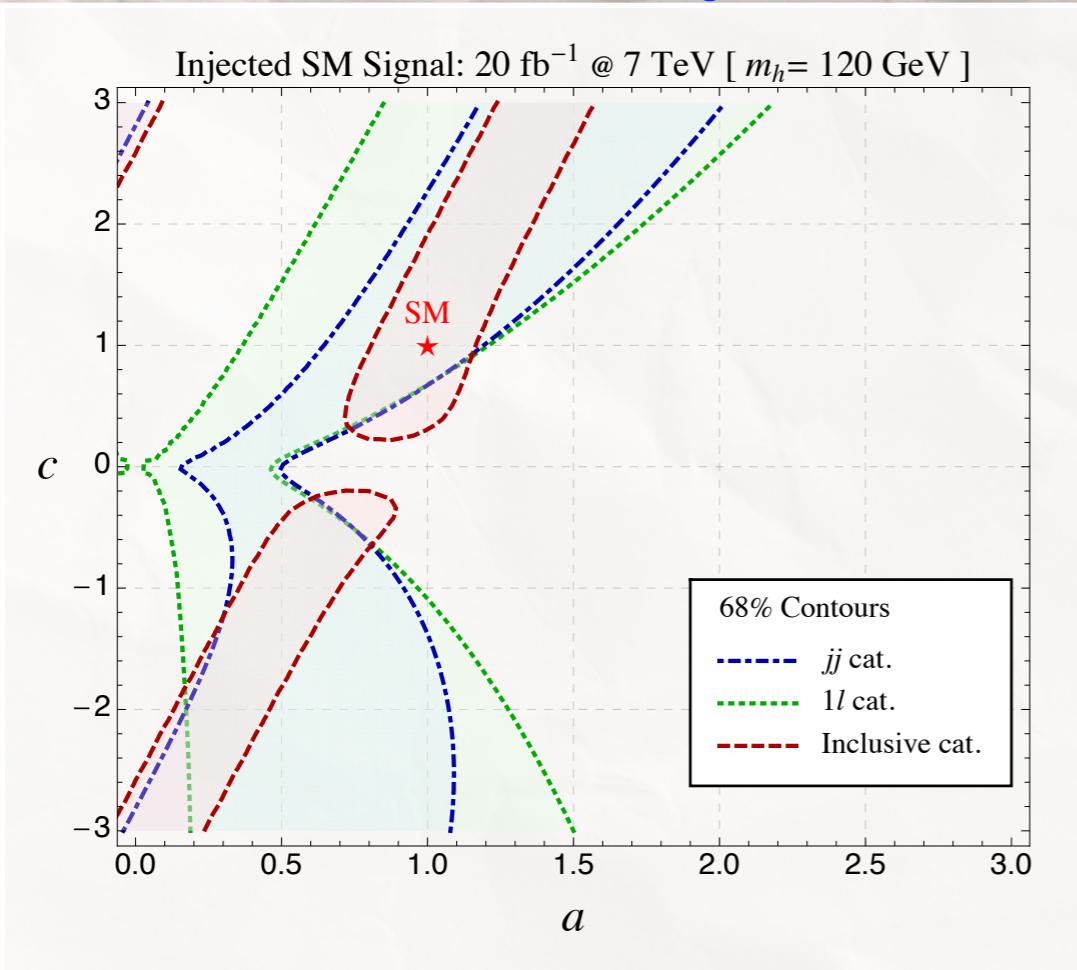
CMS all channels combined:
WW and $\gamma\gamma$ fully INCLUSIVE



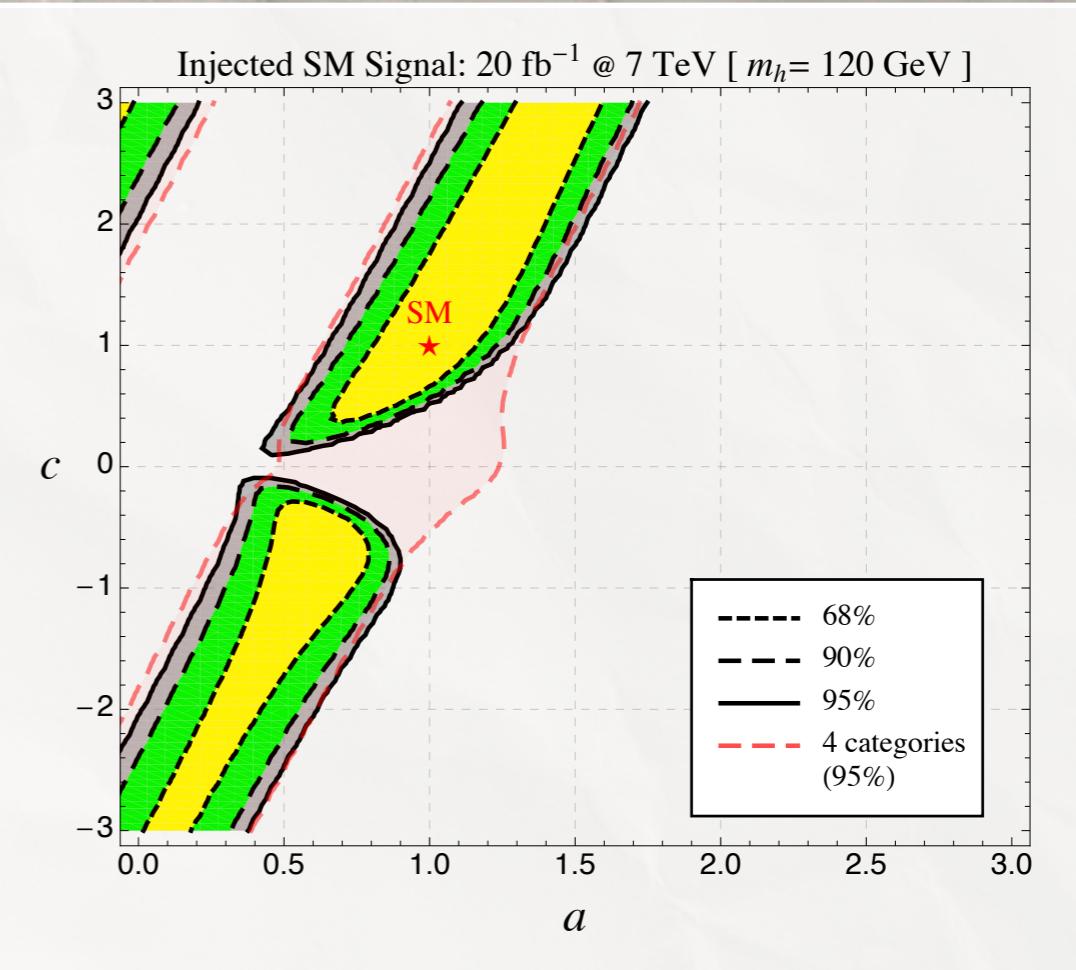
Moral: Exclusive analyses probe individual Higgs productions and are much more powerful than inclusive ones

Exclusive vs Inclusive: the $h \rightarrow \gamma\gamma$ channel

Projection at 20 fb^{-1} (SM injected)
Breakdown of exclusive categories



Projection at 20 fb^{-1} (SM injected)
Fully exclusive vs inclusive analysis



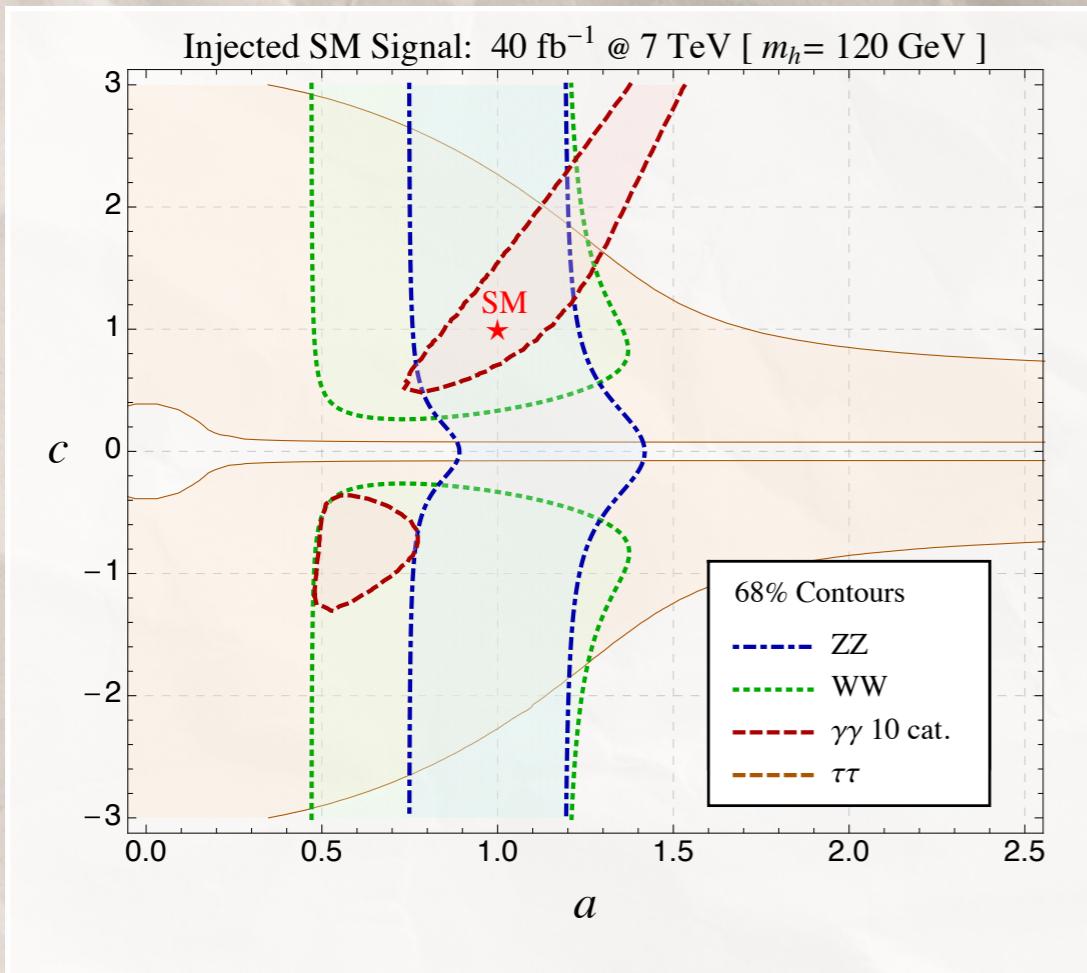
$$\mu_{jj} \sim \mu_{1l} \sim a^2 \frac{(4.5a - c)^2}{c^2}$$

$$\mu_{incl} \sim (c^2 + \zeta a^2) \frac{(4.5a - c)^2}{c^2}$$

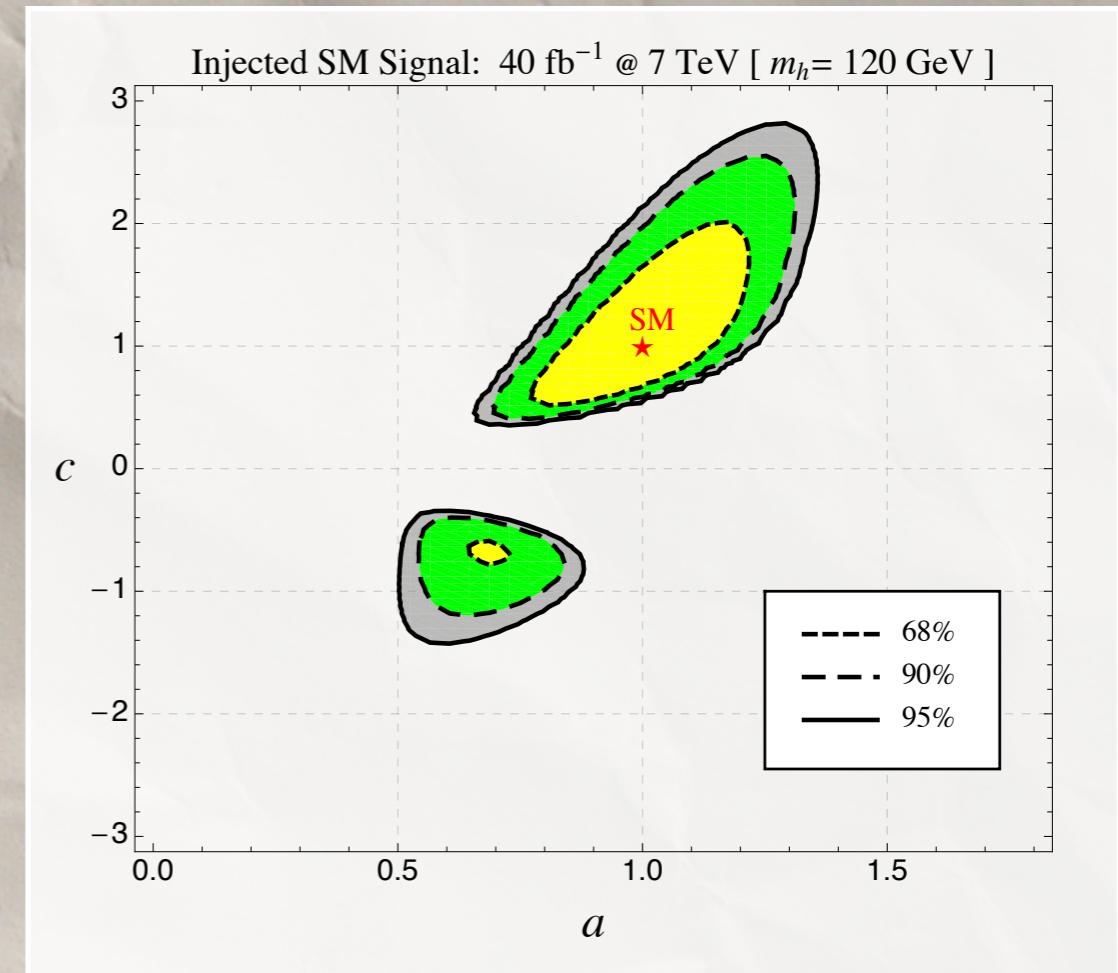
results from:
Azatov, R.C., Del Re, Galloway, Grassi, Rahatlou,
arXiv:1204.4817

Shall we break the degeneracy ?

Projection at 40 fb^{-1} (SM injected)
Breakdown of individual channels



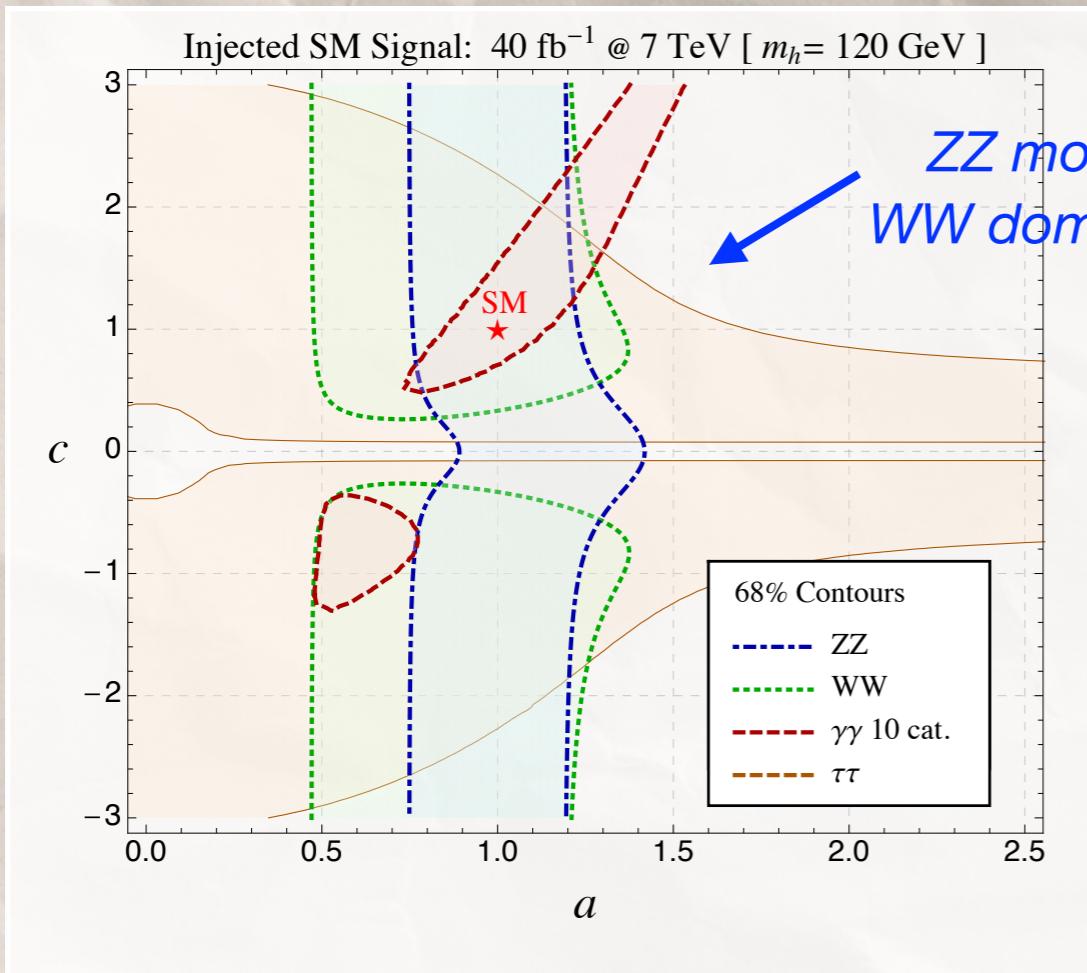
Projection at 40 fb^{-1} (SM injected)
All channels combined



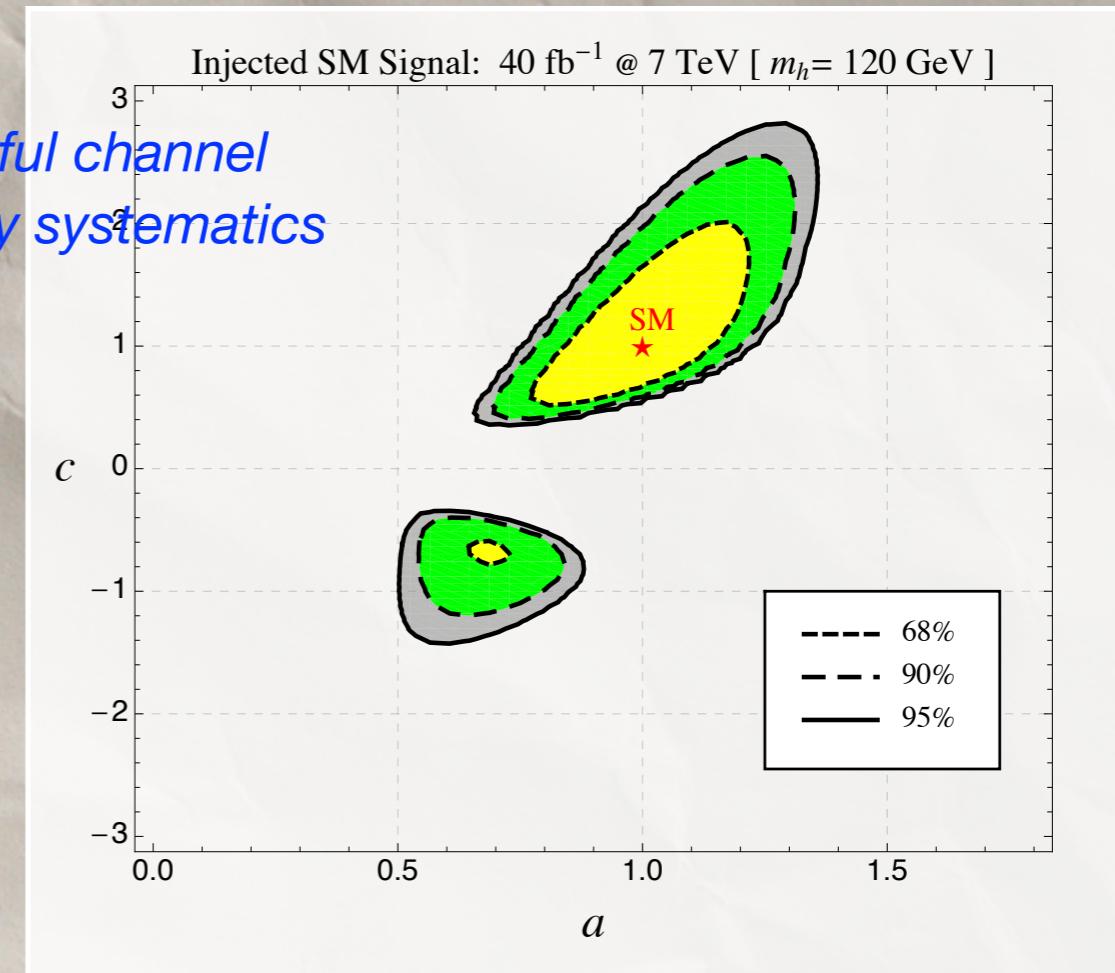
results from:
Azatov, R.C., Del Re, Galloway, Grassi, Rahatlou,
arXiv:1204.4817

Shall we break the degeneracy ?

Projection at 40 fb^{-1} (SM injected)
Breakdown of individual channels



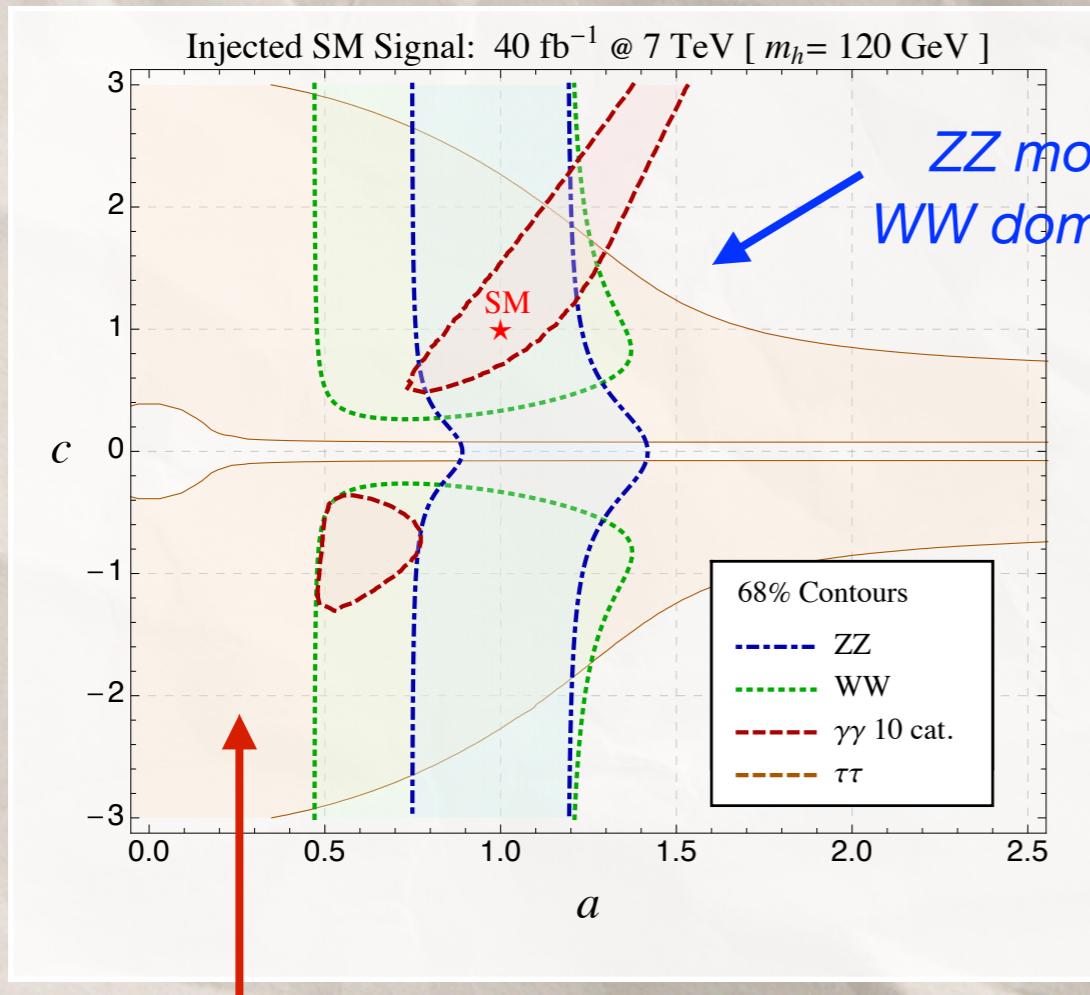
Projection at 40 fb^{-1} (SM injected)
All channels combined



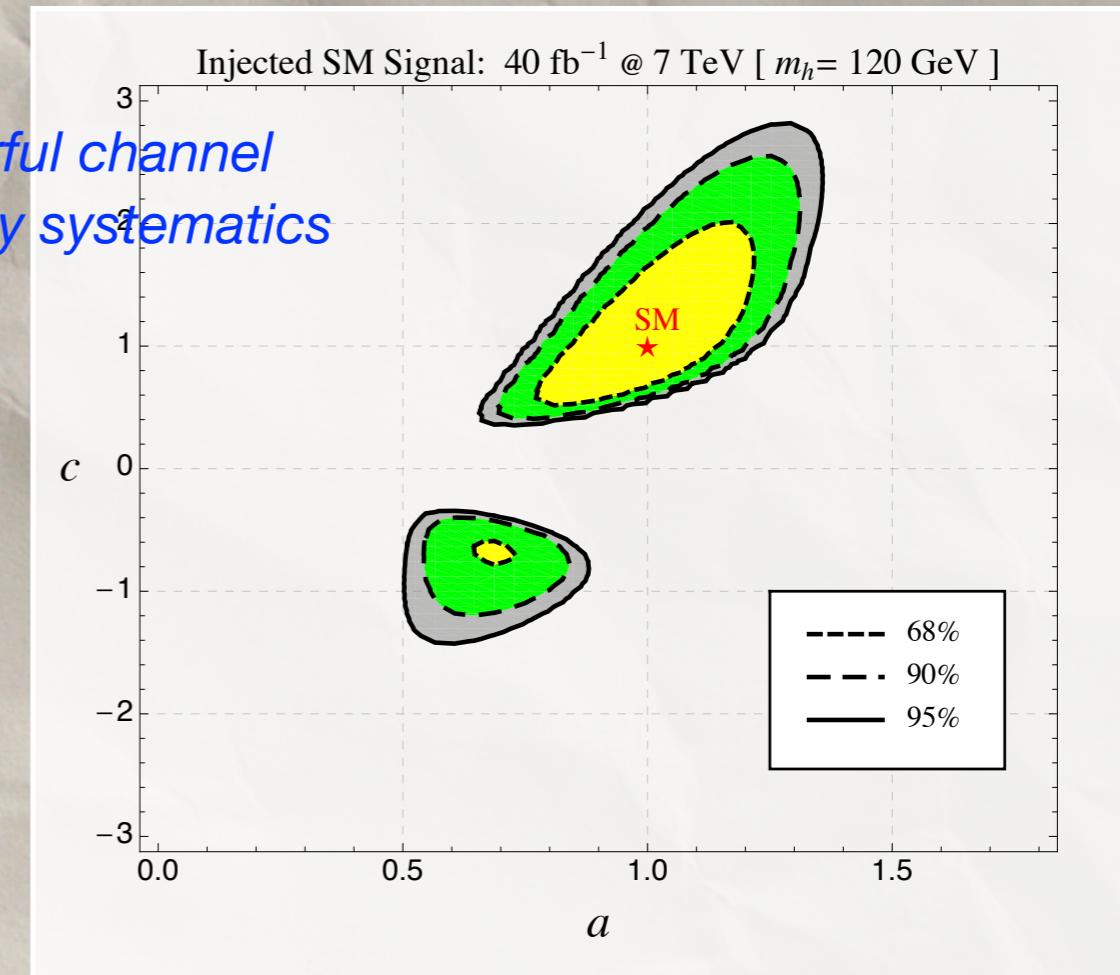
results from:
Azatov, R.C., Del Re, Galloway, Grassi, Rahatlou,
arXiv:1204.4817

Shall we break the degeneracy ?

Projection at 40 fb^{-1} (SM injected)
Breakdown of individual channels



Projection at 40 fb^{-1} (SM injected)
All channels combined



$\tau\tau$ channel currently not very powerful

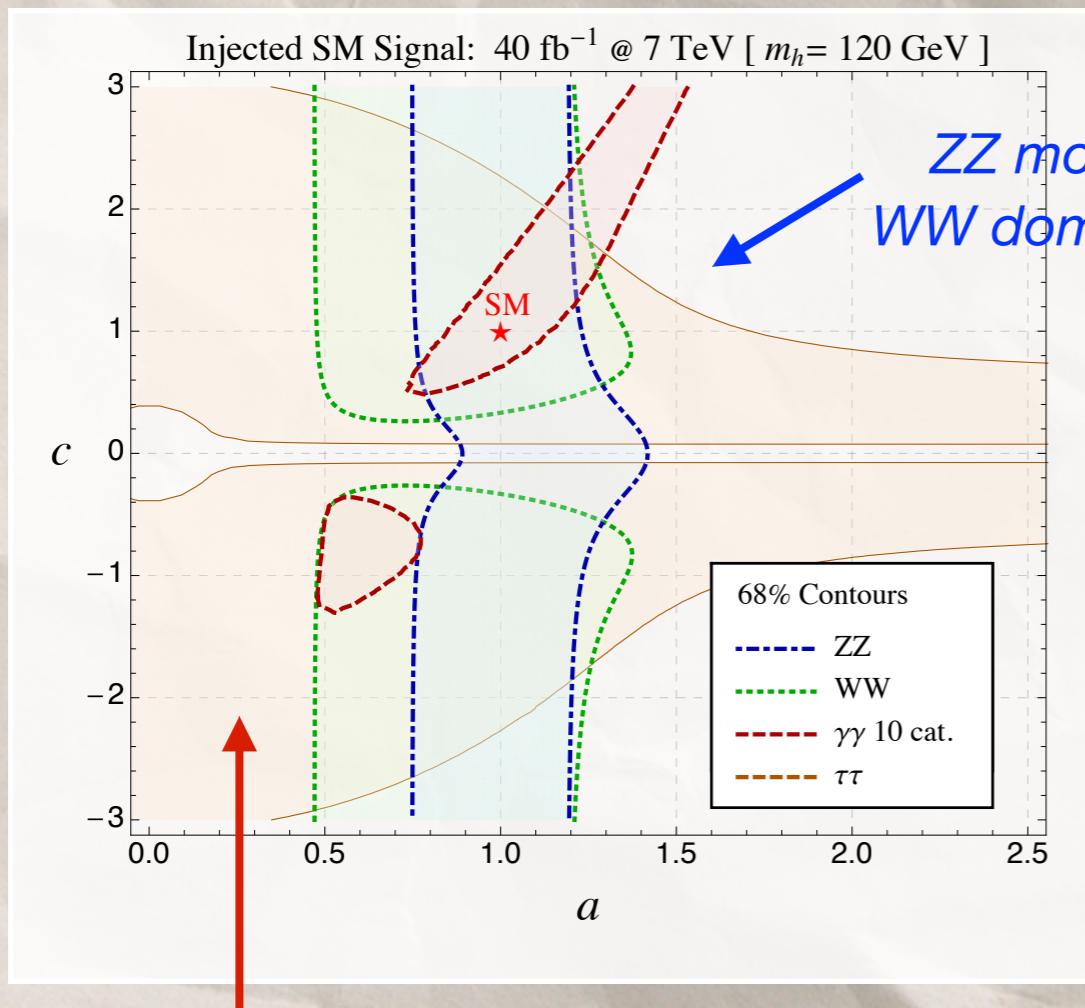
Background underestimated in old
studies on Higgs couplings

results from:

Azatov, R.C., Del Re, Galloway, Grassi, Rahatlou,
arXiv:1204.4817

Shall we break the degeneracy ?

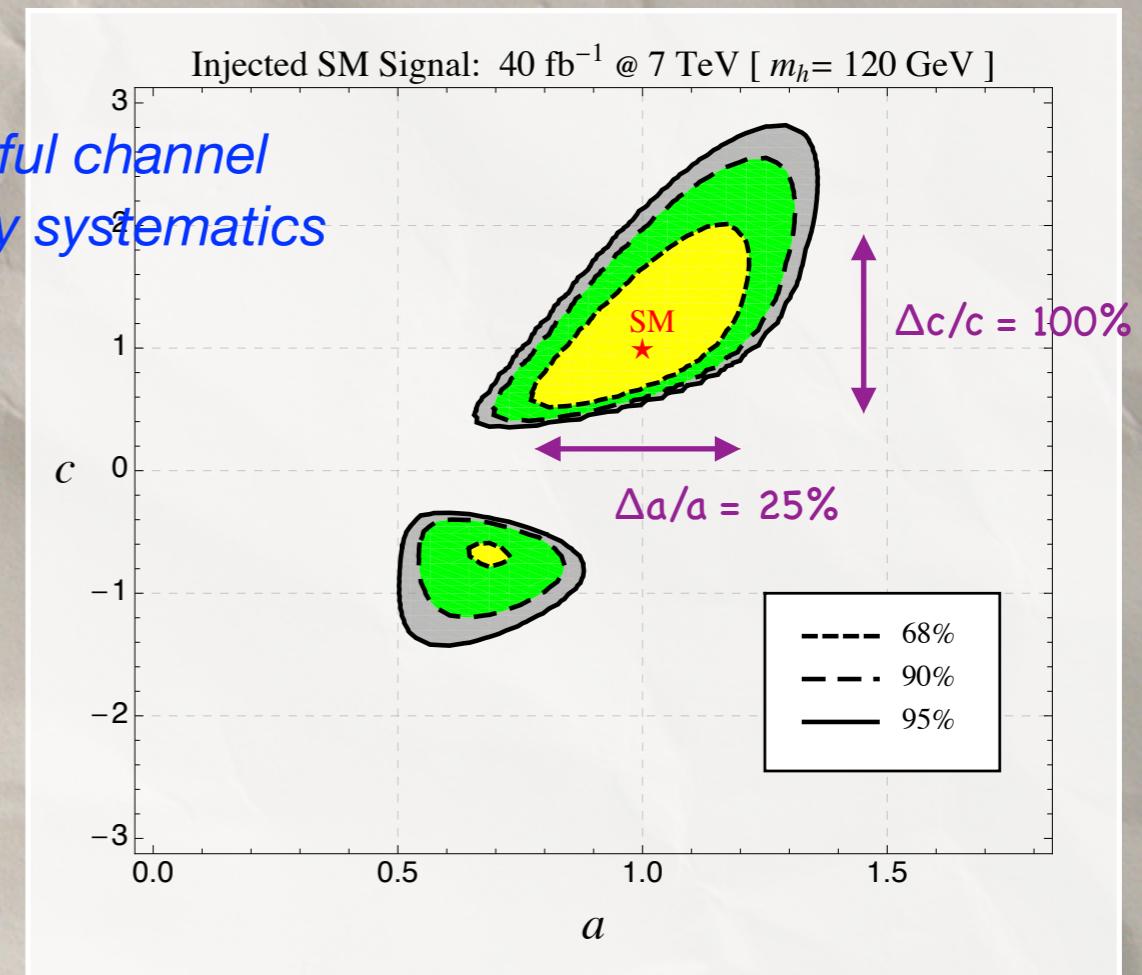
Projection at 40 fb^{-1} (SM injected)
Breakdown of individual channels



$\tau\tau$ channel currently not very powerful

Background underestimated in old
studies on Higgs couplings

Projection at 40 fb^{-1} (SM injected)
All channels combined



results from:

Azatov, R.C., Del Re, Galloway, Grassi, Rahatlou,
arXiv:1204.4817

CONCLUSIONS

CONCLUSIONS

- ▶ Results of Higgs searches can/should be presented (also) in a model-independent way

CONCLUSIONS

- ▶ Results of Higgs searches can/should be presented (also) in a model-independent way
- ▶ Theory predictions can be made in a fully-consistent and systematic way by means of Chiral Lagrangian (bottom-up approach)

CONCLUSIONS

- ▶ Results of Higgs searches can/should be presented (also) in a model-independent way
- ▶ Theory predictions can be made in a fully-consistent and systematic way by means of Chiral Lagrangian (bottom-up approach)
- ▶ First fit to (a,c) based on LHC data shows: i) SM is in good shape; ii) there is a second, degenerate solution

CONCLUSIONS

- ▶ Results of Higgs searches can/should be presented (also) in a model-independent way
- ▶ Theory predictions can be made in a fully-consistent and systematic way by means of Chiral Lagrangian (bottom-up approach)
- ▶ First fit to (a,c) based on LHC data shows: i) SM is in good shape; ii) there is a second, degenerate solution
- ▶ Exclusive vs Inclusive analyses much more powerful in a model-independent approach

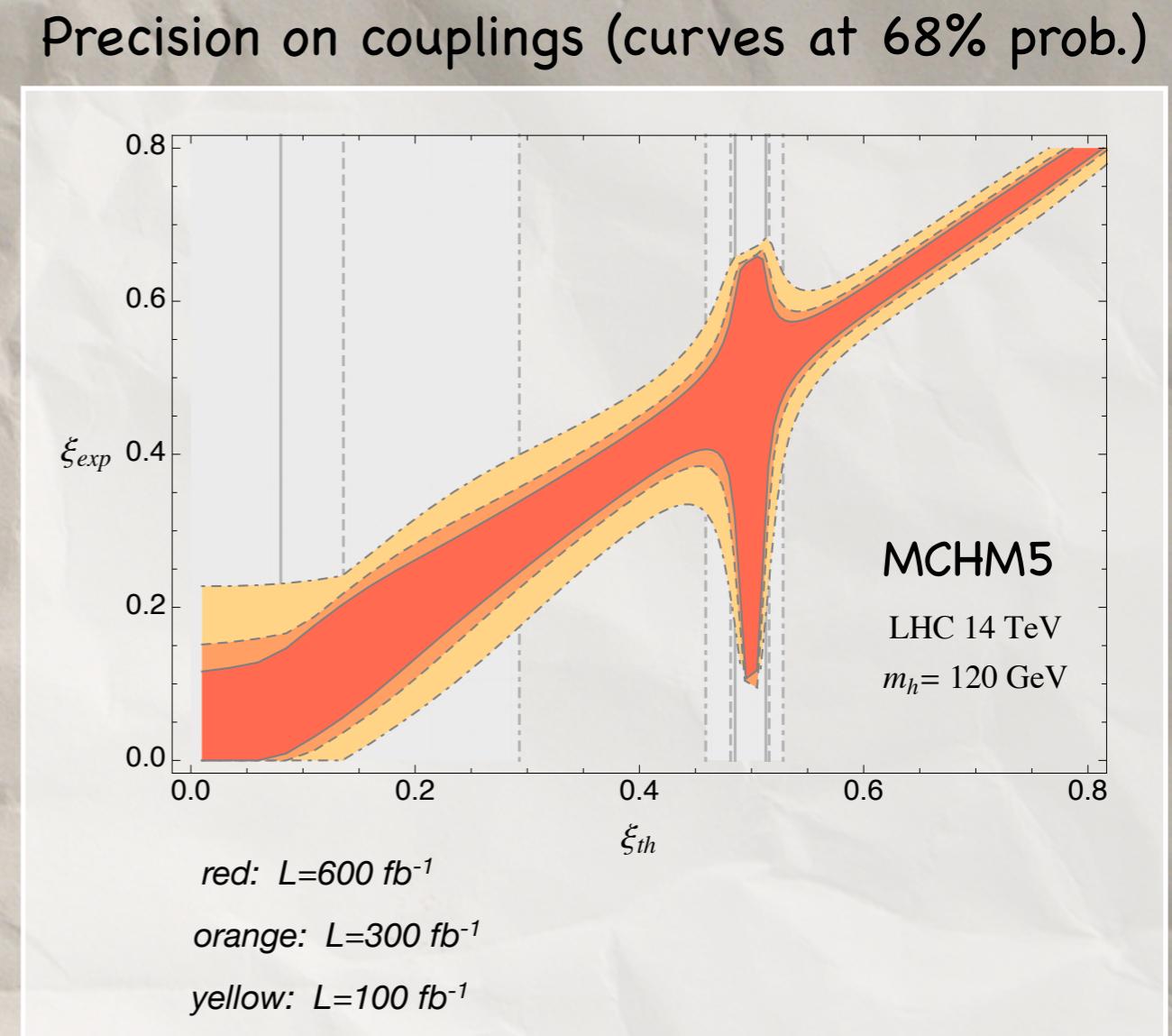
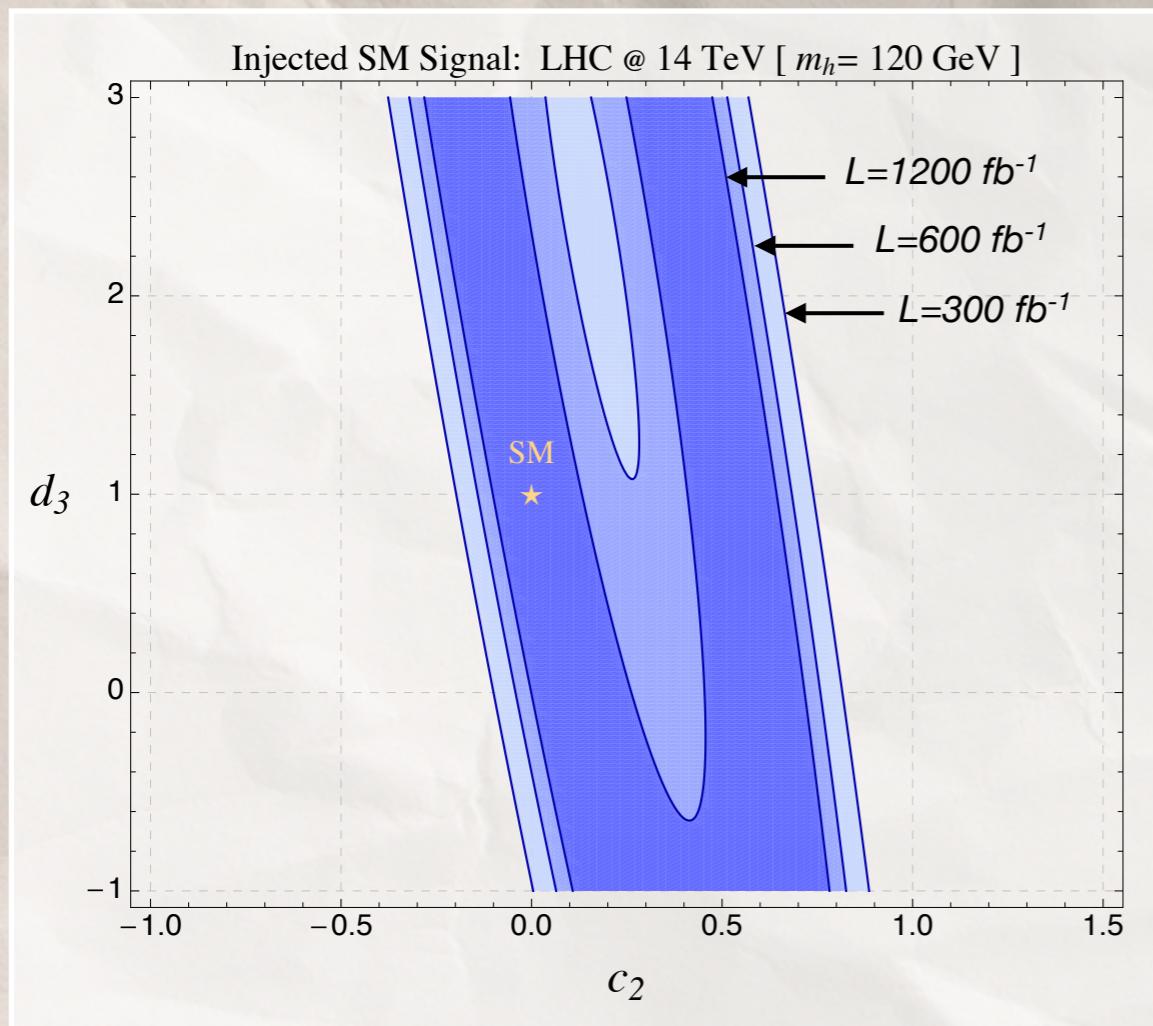
CONCLUSIONS

- ▶ Results of Higgs searches can/should be presented (also) in a model-independent way
- ▶ Theory predictions can be made in a fully-consistent and systematic way by means of Chiral Lagrangian (bottom-up approach)
- ▶ First fit to (a,c) based on LHC data shows: i) SM is in good shape; ii) there is a second, degenerate solution
- ▶ Exclusive vs Inclusive analyses much more powerful in a model-independent approach
- ▶ Double Higgs production via gluon fusion ($gg \rightarrow hh$) strongly sensitive on $tthh$ coupling (c_2)

EXTRA SLIDES

Double Higgs production via gluon fusion

Precision on couplings (curves at 68% prob.)

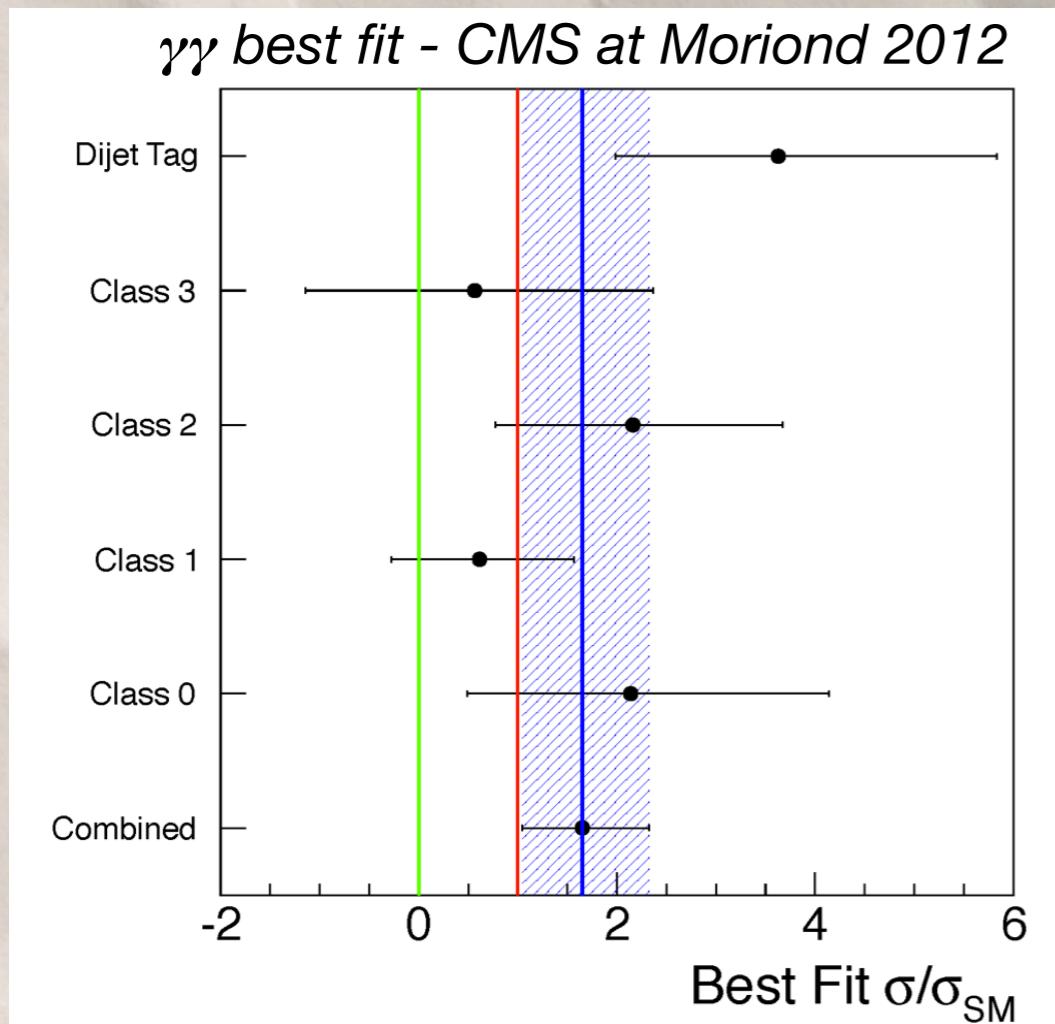


results from:

R.C., Ghezzi, Moretti, Panico, Piccinini, Wulzer,
arXiv:1205.5444

Ex: with $L=300 \text{ fb}^{-1}$ $\Delta \xi / \xi = 30\%$ for $\xi=0.2$

Exclusive vs Inclusive: the $h \rightarrow \gamma\gamma$ channel



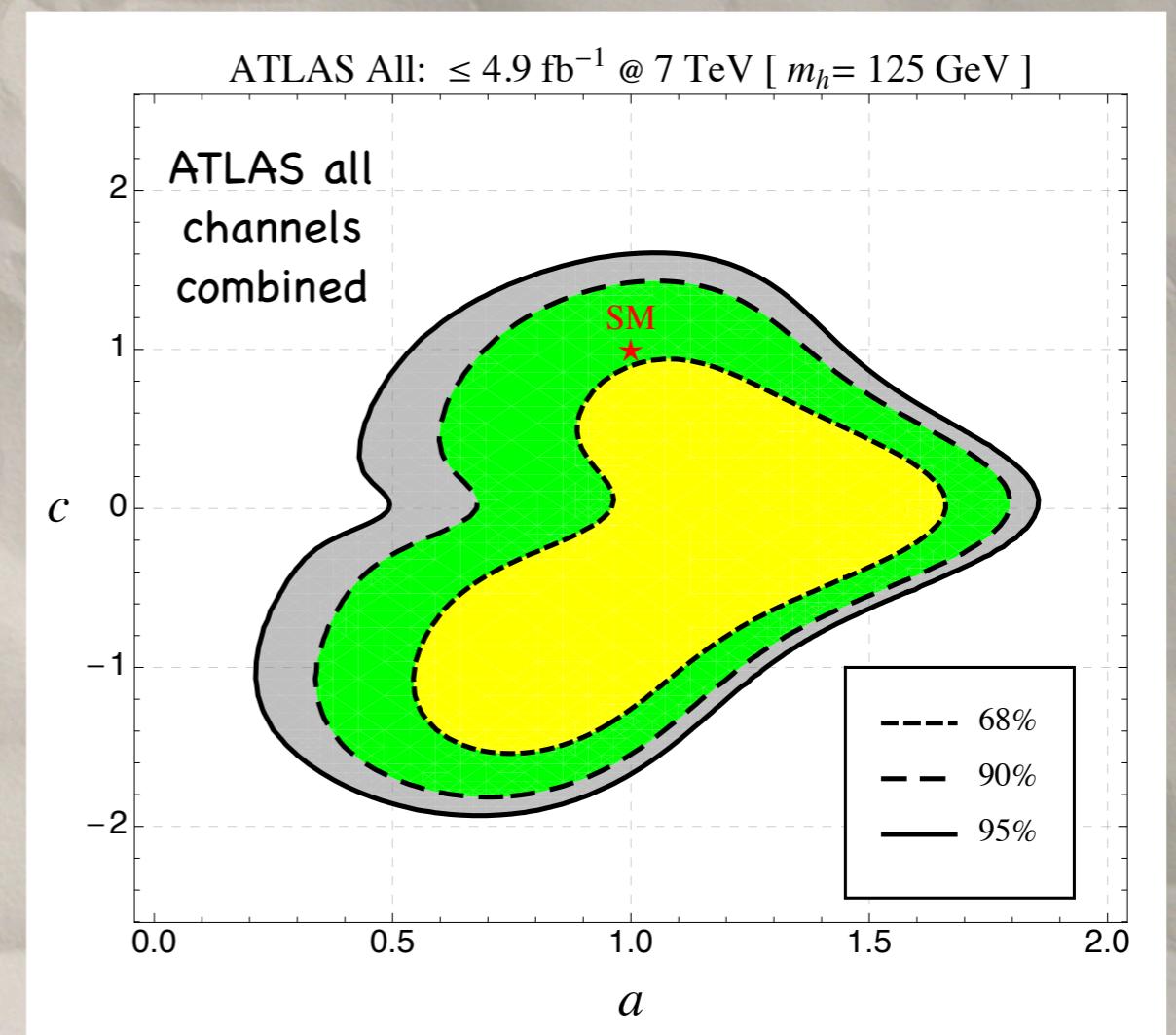
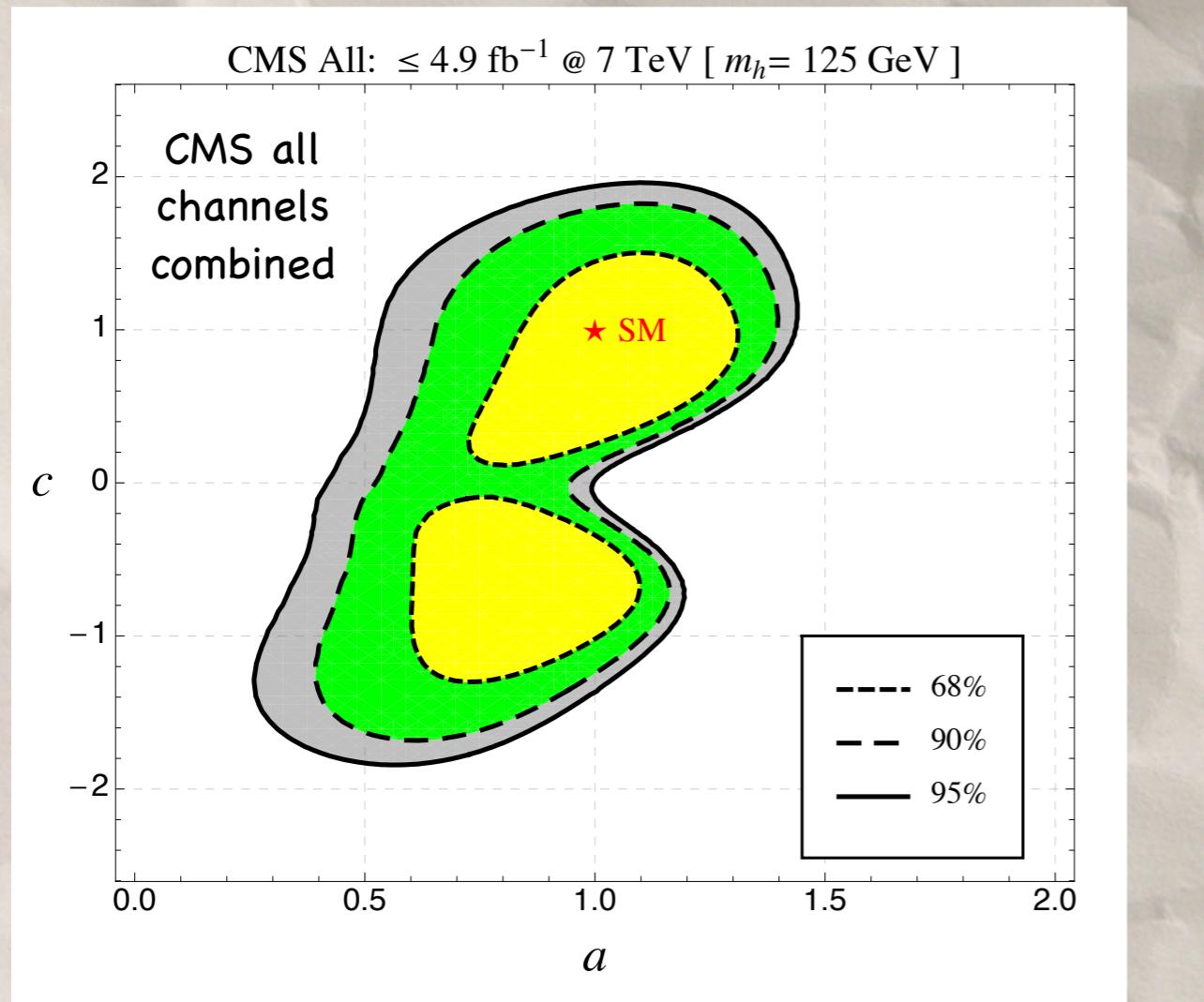
First solution:
 $(a=1.0, c=0.75)$

Second solution
 $(a=0.85, c=-0.6)$

$\mu(\gamma\gamma jj) \sim 2$
$\mu(\gamma\gamma incl) \sim 1$
$\mu(WW2j) \sim 1.2$
$\mu(WW0j, WW1j) \sim 1.7$
$\mu(\gamma\gamma jj) \sim 2.4$
$\mu(\gamma\gamma incl) \sim 1.3$
$\mu(WW2j) \sim 1.1$
$\mu(WW0j, WW1j) \sim 0.7$

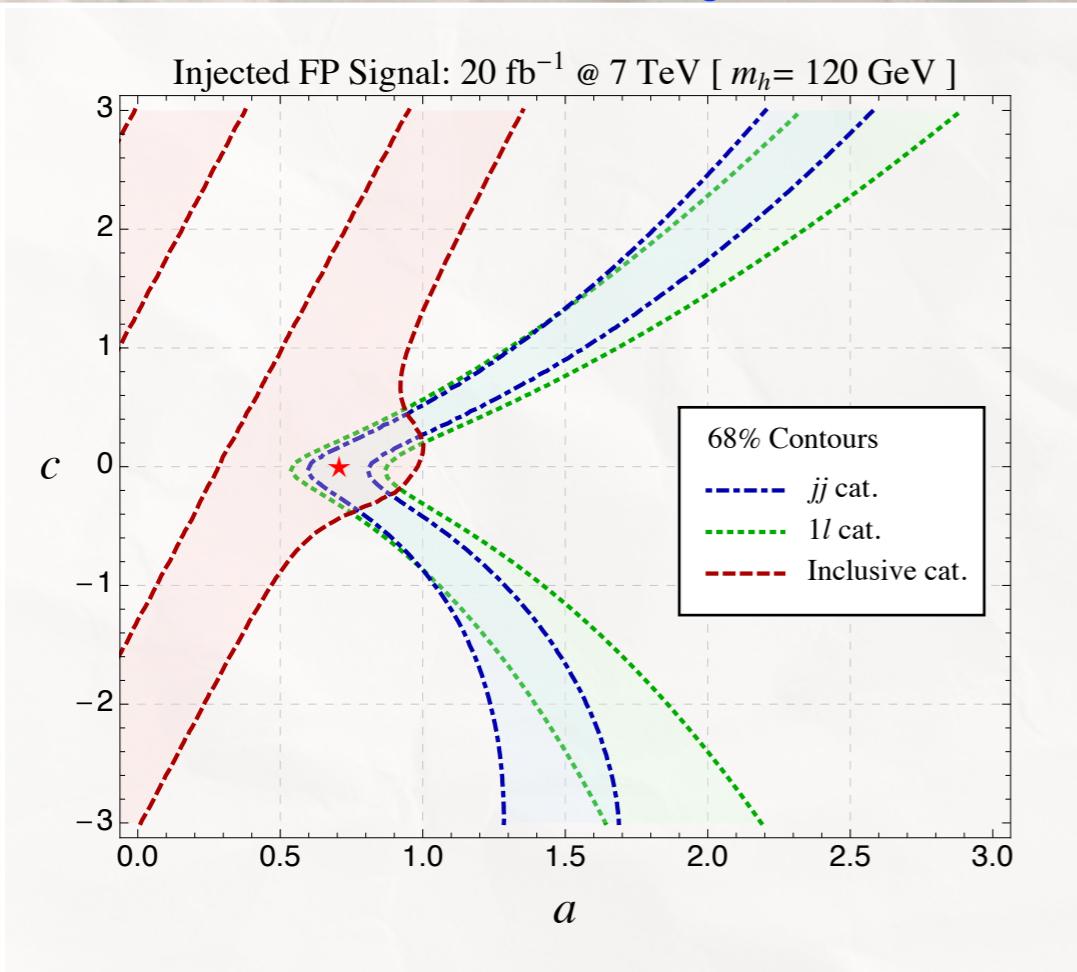
$$\mu(i) \equiv \frac{\sigma \times BR(i)}{[\sigma \times BR(i)]_{SM}}$$

Fit at $m_H=125\text{GeV}$

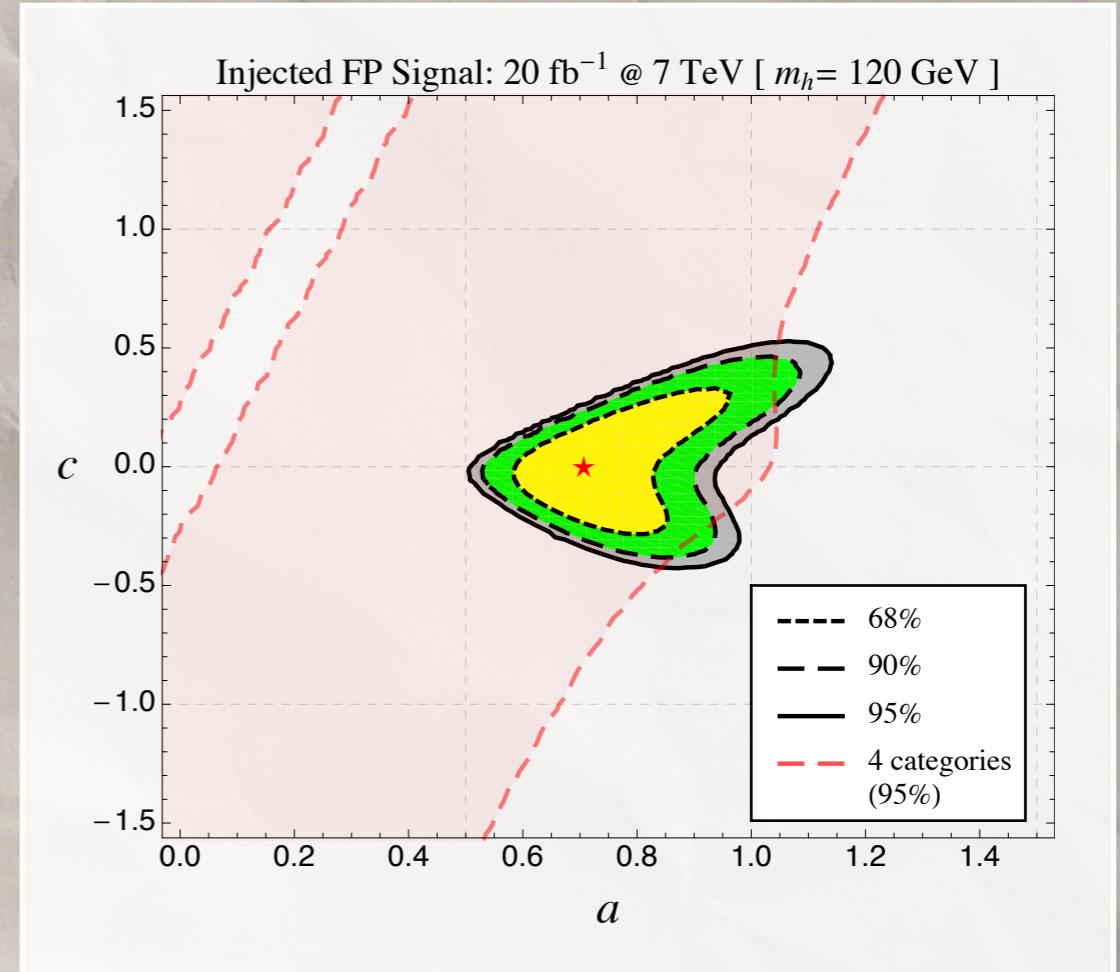


Exclusive vs Inclusive: the $h \rightarrow \gamma\gamma$ channel

FP signal ($a=1/\sqrt{2}$, $c=0$) injected
Breakdown of exclusive categories



FP signal ($a=1/\sqrt{2}$, $c=0$) injected
Fully exclusive vs inclusive analysis

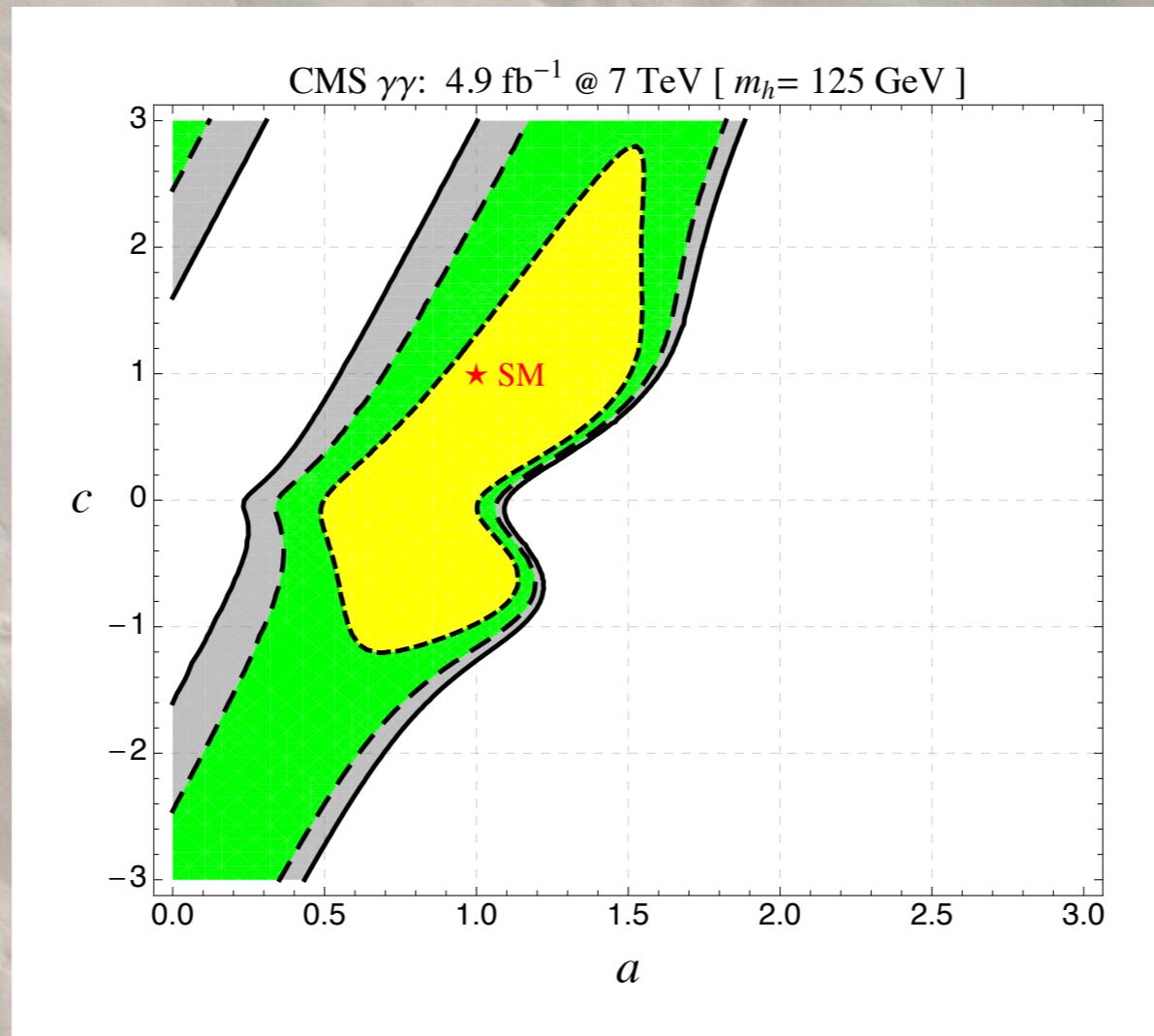


results from:
Azatov, R.C., Del Re, Galloway, Grassi, Rahatlou,
arXiv:1204.4817

FP point ($a=1/\sqrt{2}$, $c=0$) still allowed by CMS
combined limits for $123 \text{ GeV} < m_H < 130 \text{ GeV}$

Exclusive vs Inclusive: the $h \rightarrow \gamma\gamma$ channel

Fit of $\gamma\gamma$ based on CMS data



courtesy of A. Azatov

Shall we break the degeneracy ?

