

Black Hole's Quantum Portrait

Gia Dvali

LMU-MPI, CERN, NYU

with: Cesar Gomez

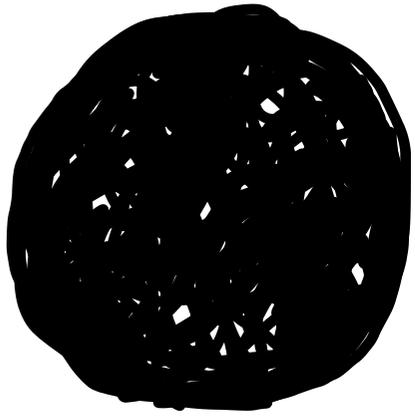
The Big Goal is to understand UV-completion of gravity.

This is impossible without understanding quantum portrait of Black Holes.

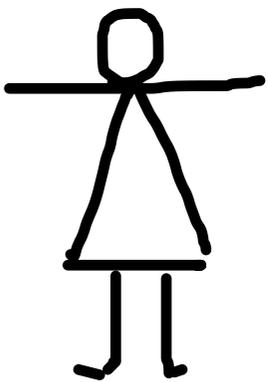
In this talk we shall introduce such a portrait.

Black Hole Mysteries (semi-classically):

- * Absence of hair;
- * Exact thermality of Hawking radiation and negative heat;
- * Bekenstein entropy;
- *



Must be a quantum
field-theoretic substance
at temperature T_H !



But, none work!

Absence of hair and exact
thermality

+

A small logical gap filled
with a seemingly-logical assumption

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* "Folk theorems" about
no global charges (e.g. baryon
and lepton numbers).

* "Information Paradox".

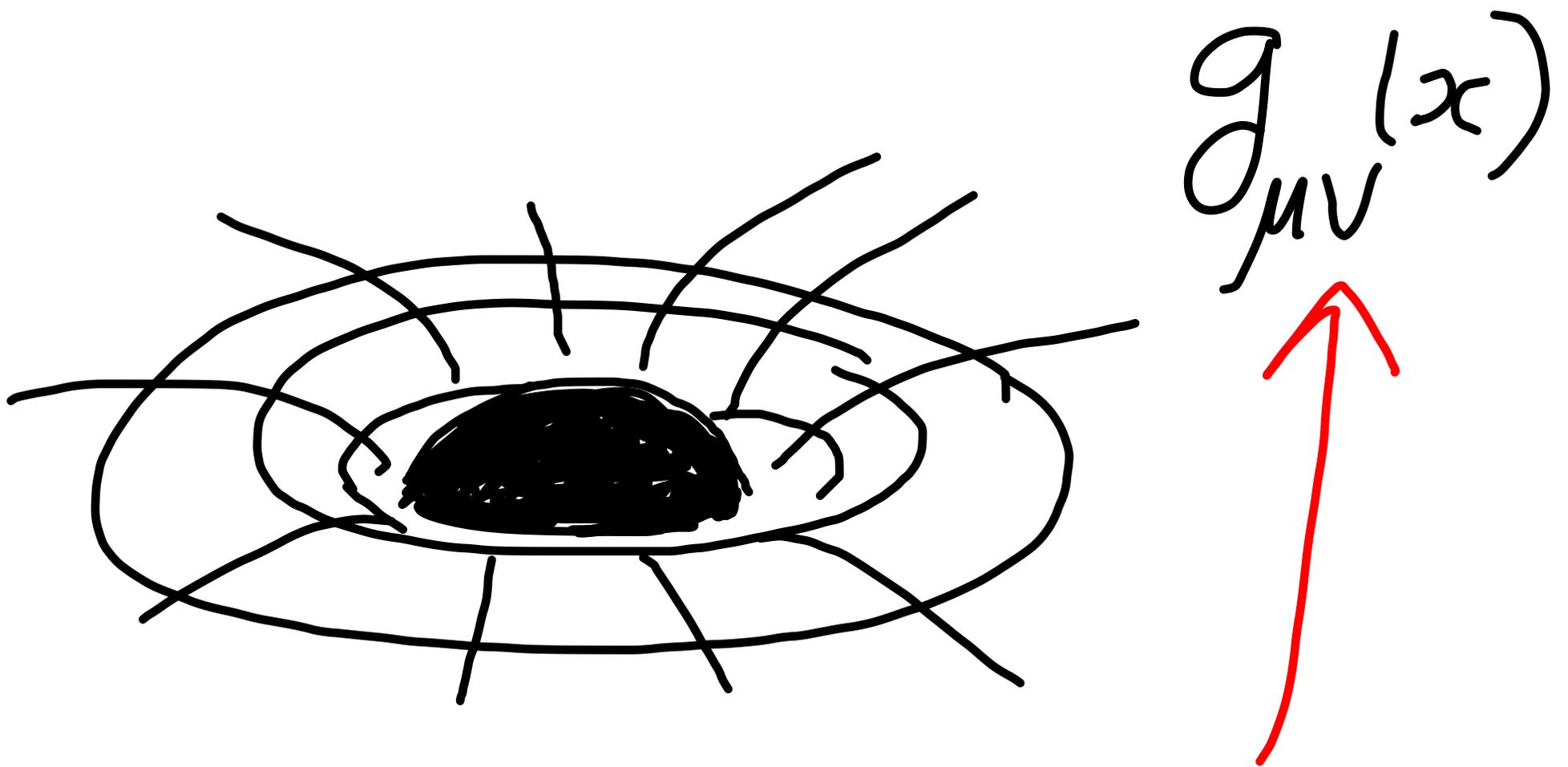
To resolve these "paradoxes," and to close the logical gap, we need a microscopic quantum theory.

In this talk we shall provide such a theory and show how it demystifies semi-classical black hole properties.

We shall see:

Black holes do carry hair under global charges (baryonic and leptonic numbers), which can be of 100% astrophysical importance.

Recall:
Schwarzschild black
hole is a solution in GR



Intrinsically - classical
concept!

In quantum field-theory
the building blocks are
particles:

$$a^{\dagger}|0\rangle = |1\rangle$$

There is nothing
else.

Our main concept:

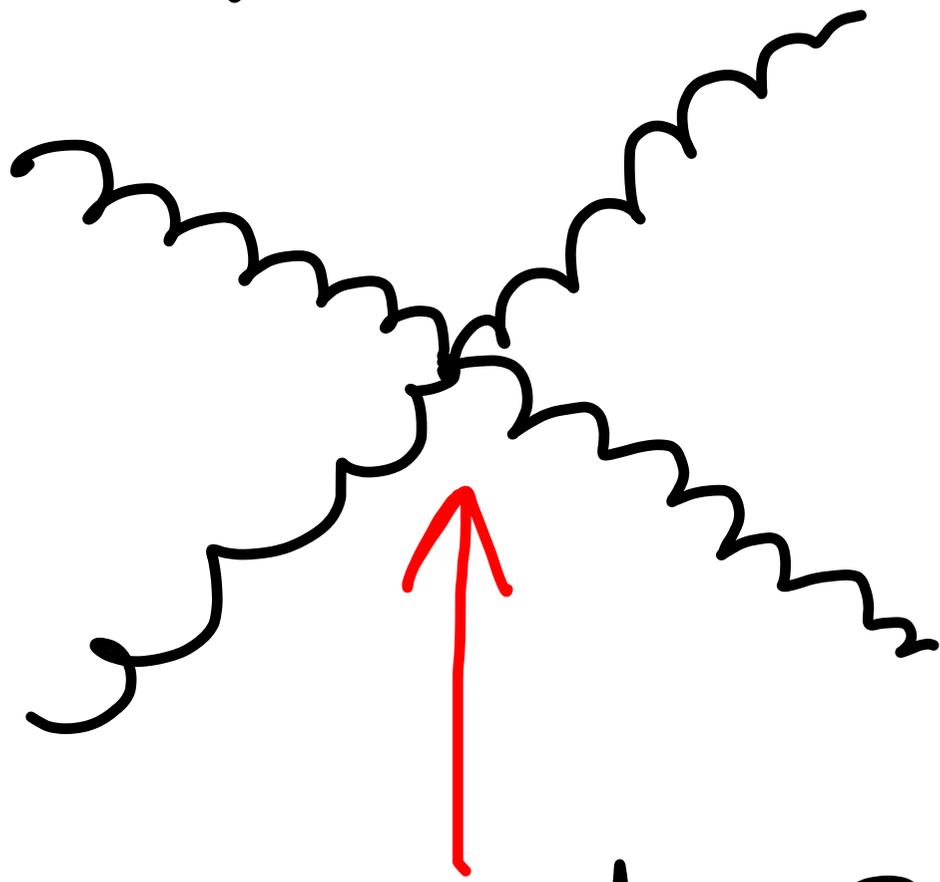
geometry = many quanta

$$g_{\mu\nu} = \sum N_\lambda a_\lambda^\dagger |0\rangle$$

Geometry is a quantum

Bose-condensate of
gravitons!

Gravity is a quantum
theory of a particle
(graviton) of $m = 0$
and Spin = 2



$$\alpha_{gr} \equiv h G_N \lambda^{-2}$$

Quantum entities:
Planck length and Mass

$$L_p^2 \equiv \hbar G_N, \quad M_p \equiv \frac{\hbar}{L_p}$$

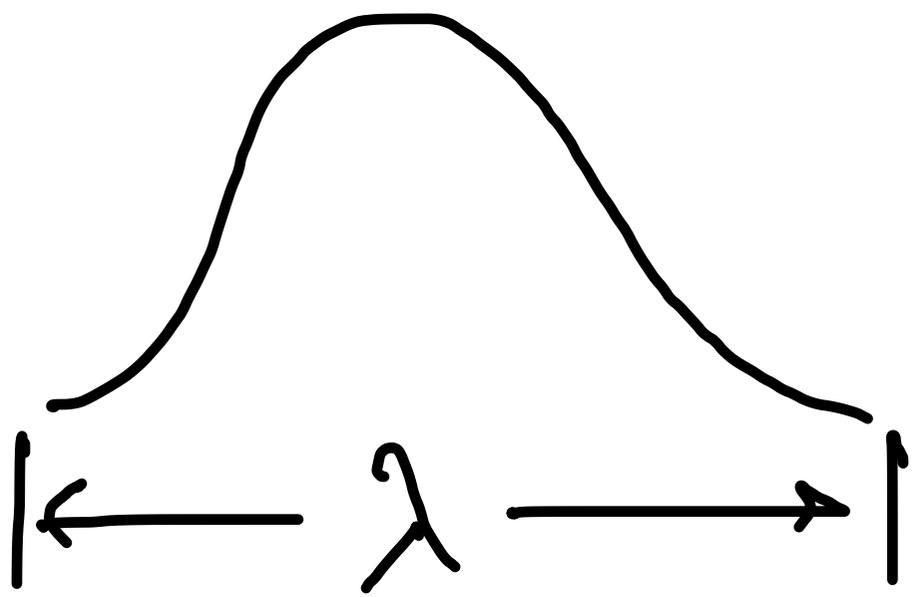
$$\alpha_{gr} = \frac{L_p^2}{\lambda^2}$$

In classical limit ($\hbar \rightarrow 0$)

$$L_p \rightarrow 0$$

$$\alpha_{gr} \rightarrow 0$$

Now, try to form a graviton wave packet.



For $\lambda \gg L_p$

$$\alpha_{gr} \ll 1$$

A typical Hartree situation:

Each graviton sees a collective potential.

Collective binding
potential for $r \sim \lambda$

$$V = -N \alpha_g \frac{\hbar}{\lambda}$$

and kinetic energy

$$E_k = \frac{\hbar}{\lambda}$$

The boundstate condition

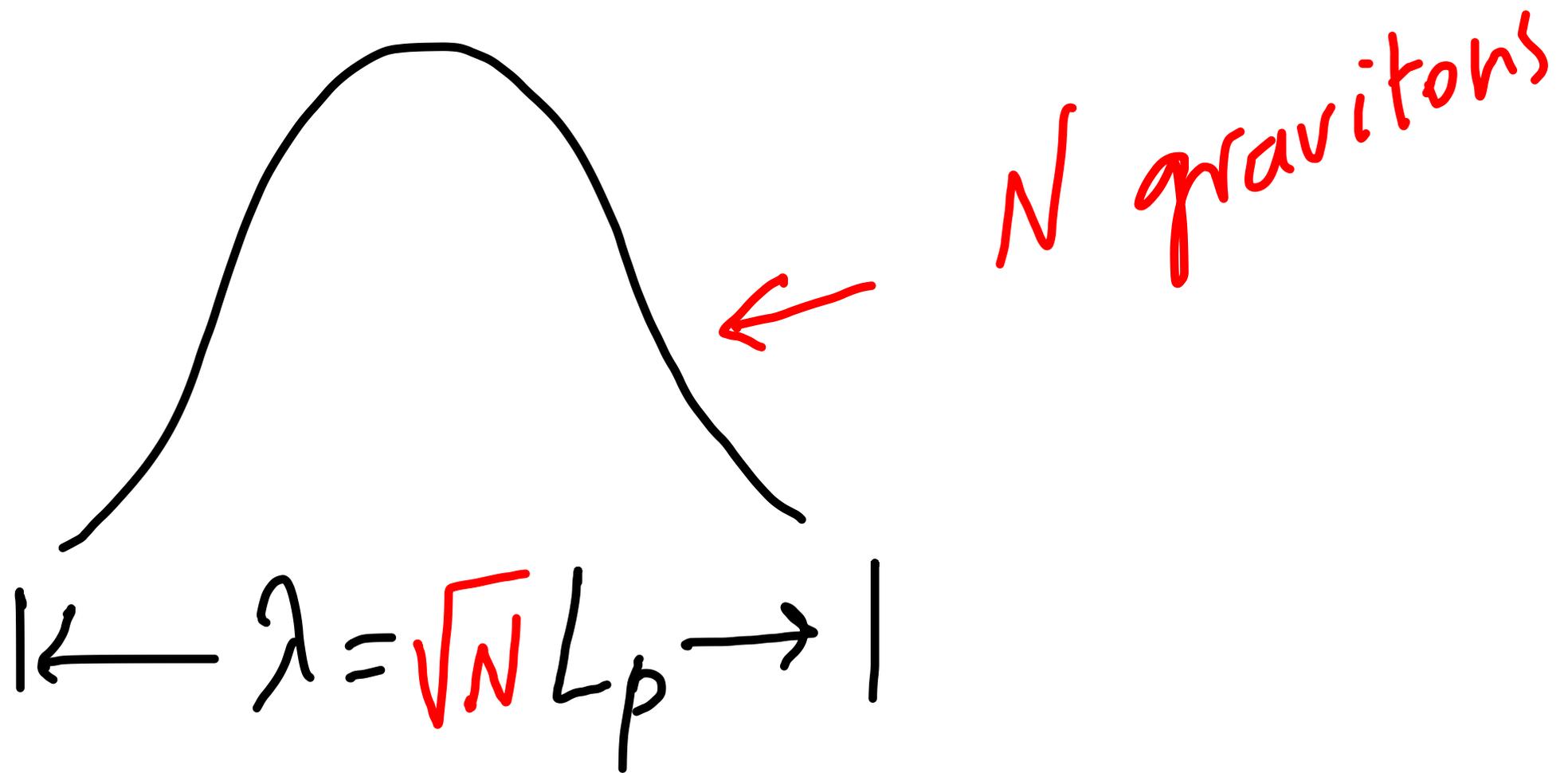
$$E_k + V = 0$$



$$(1 - N \alpha_{gr}) \frac{\hbar}{\lambda} = 0$$

A self-sustained boundstate is formed for

$$\alpha_{gr} = \frac{1}{N} !$$



This self-sustained
boundstate is a black
hole

$$\lambda = \sqrt{N} L_p, \quad \alpha_{gr} = \frac{1}{N}$$

Black hole quantum physics is remarkably simple, with a single parameter N :

$$M = \sqrt{N}, \quad \lambda = \sqrt{N},$$

$$\alpha_{gr} = \frac{1}{N}$$

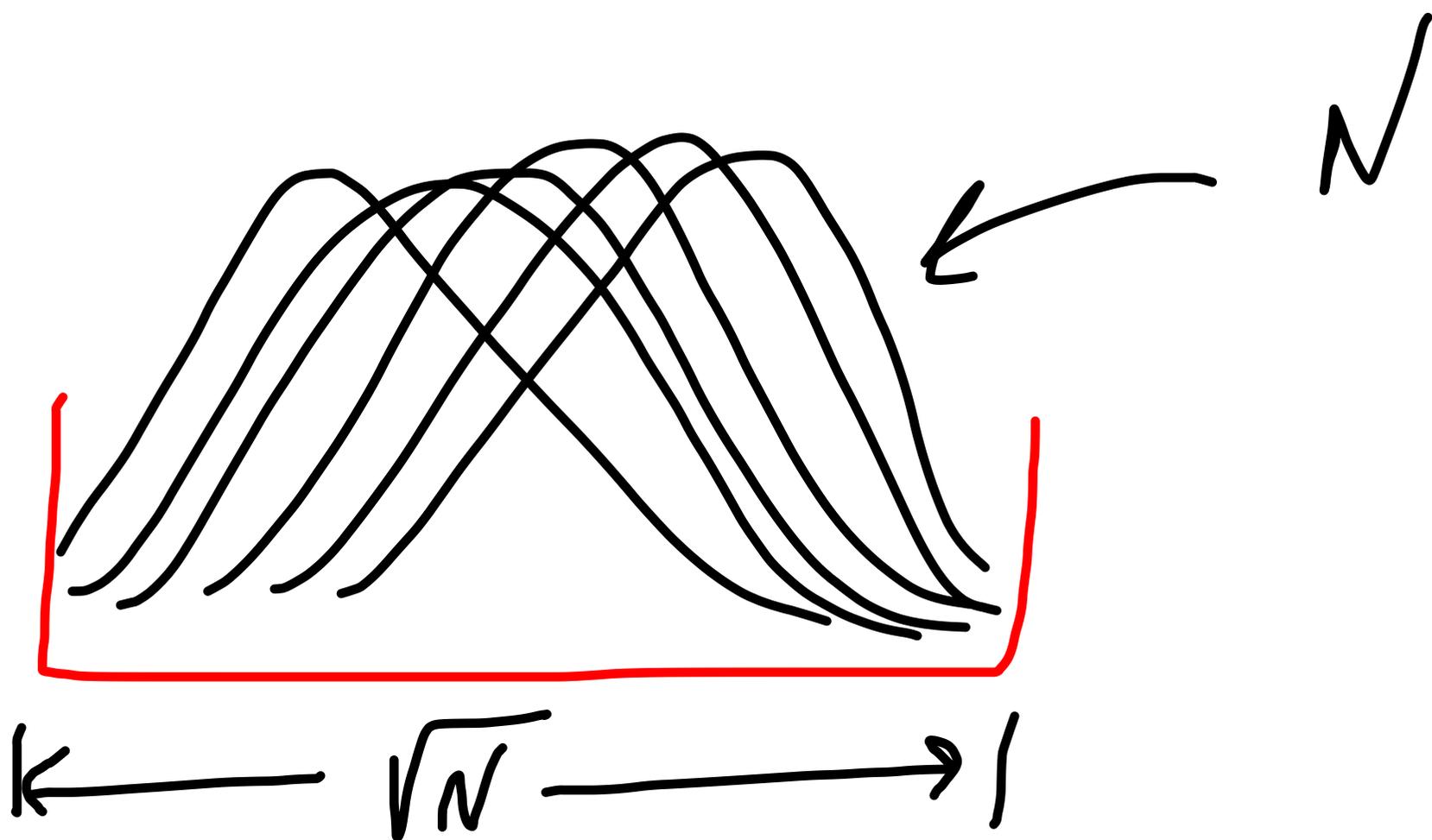
It is a large- N physics (in 't Hooft's sense) and is a result of maximal overpacking.

Black hole is a
most overpacked quantum
system of nature

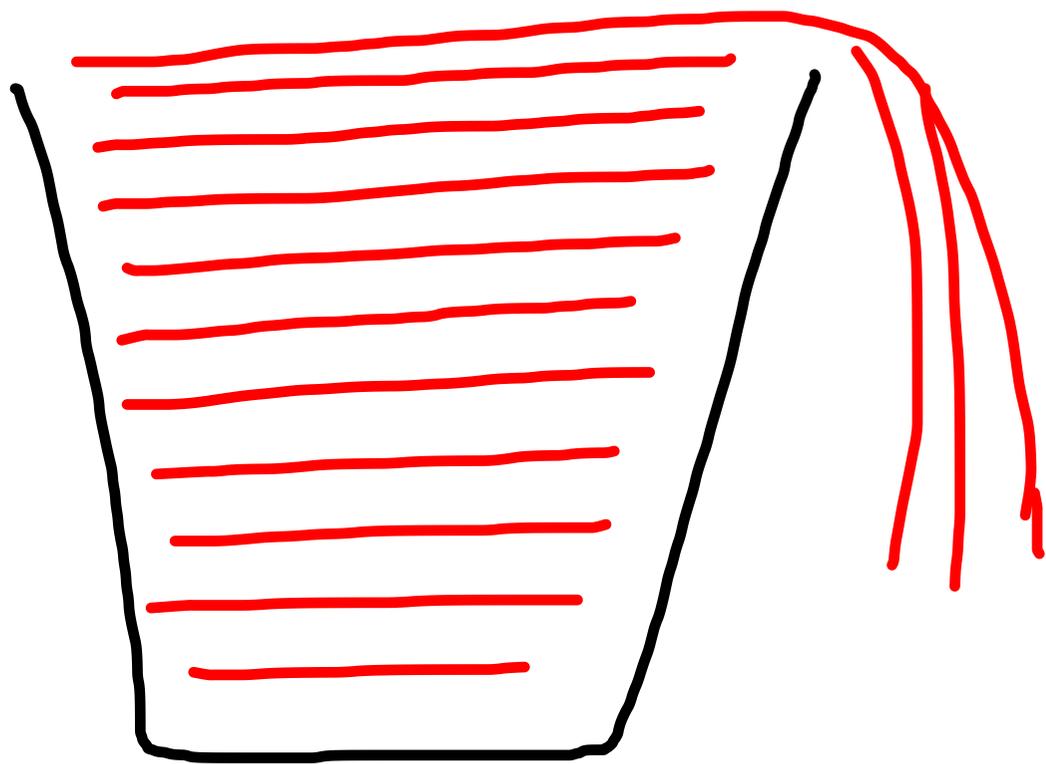
and

because of this it is

maximally simple

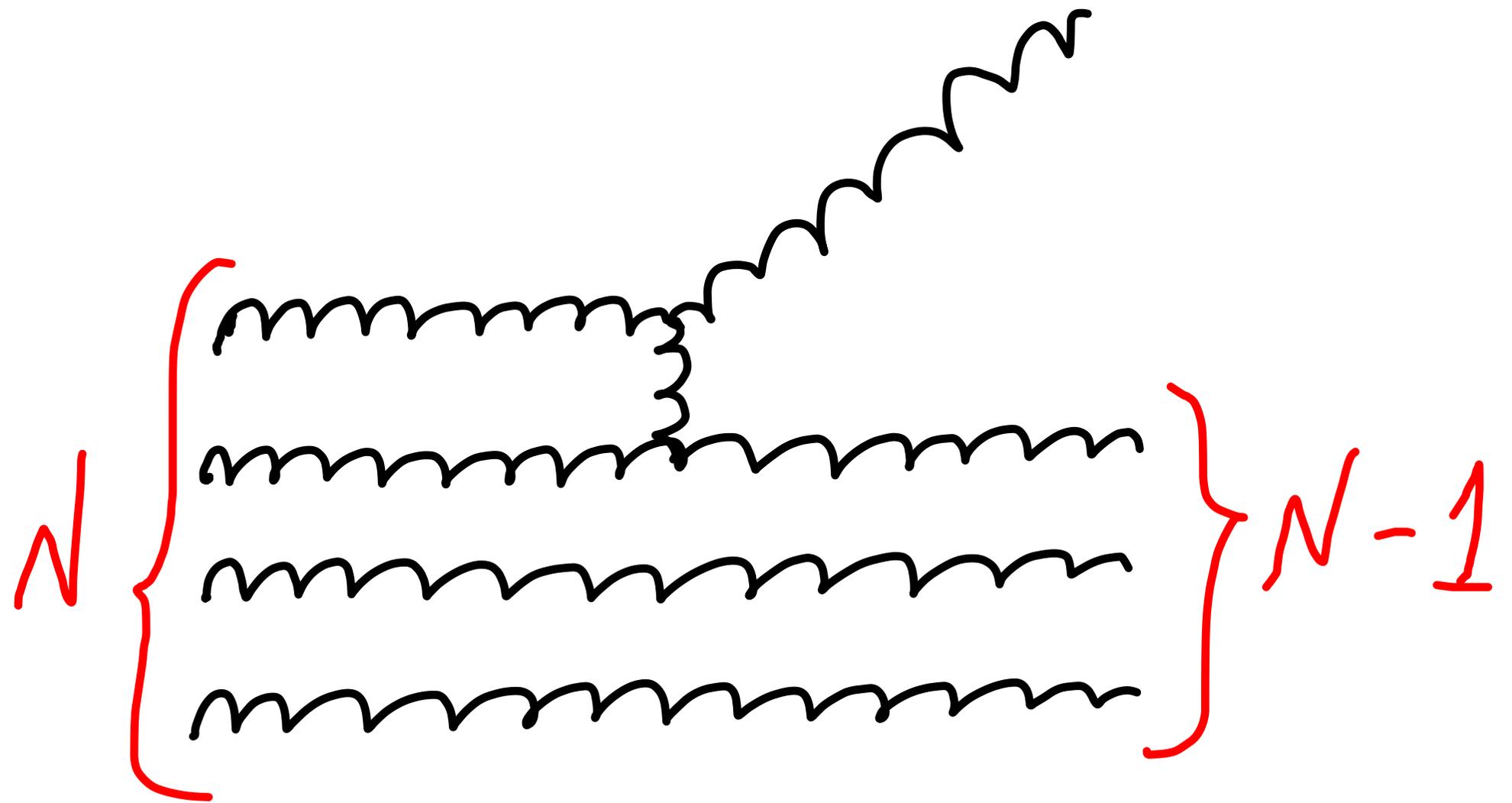


This self-sustained
Bose-condensate exists
for any N and for any
 N it is leaky



The condensate
depletes self-similarly

$$N \rightarrow N-1$$



depletion law

$$\dot{N} = -\frac{1}{\sqrt{N} L_P} + \mathcal{O}\left(\frac{1}{N^{3/2}}\right)$$

$$\dot{N} = -\frac{1}{\sqrt{N} L_p}$$

Defining $T \equiv \frac{\hbar}{\sqrt{N} L_p}$,

in the semi-classical limit

$$N \rightarrow \infty, L_p \rightarrow 0, \sqrt{N} L_p = \text{fixed}$$

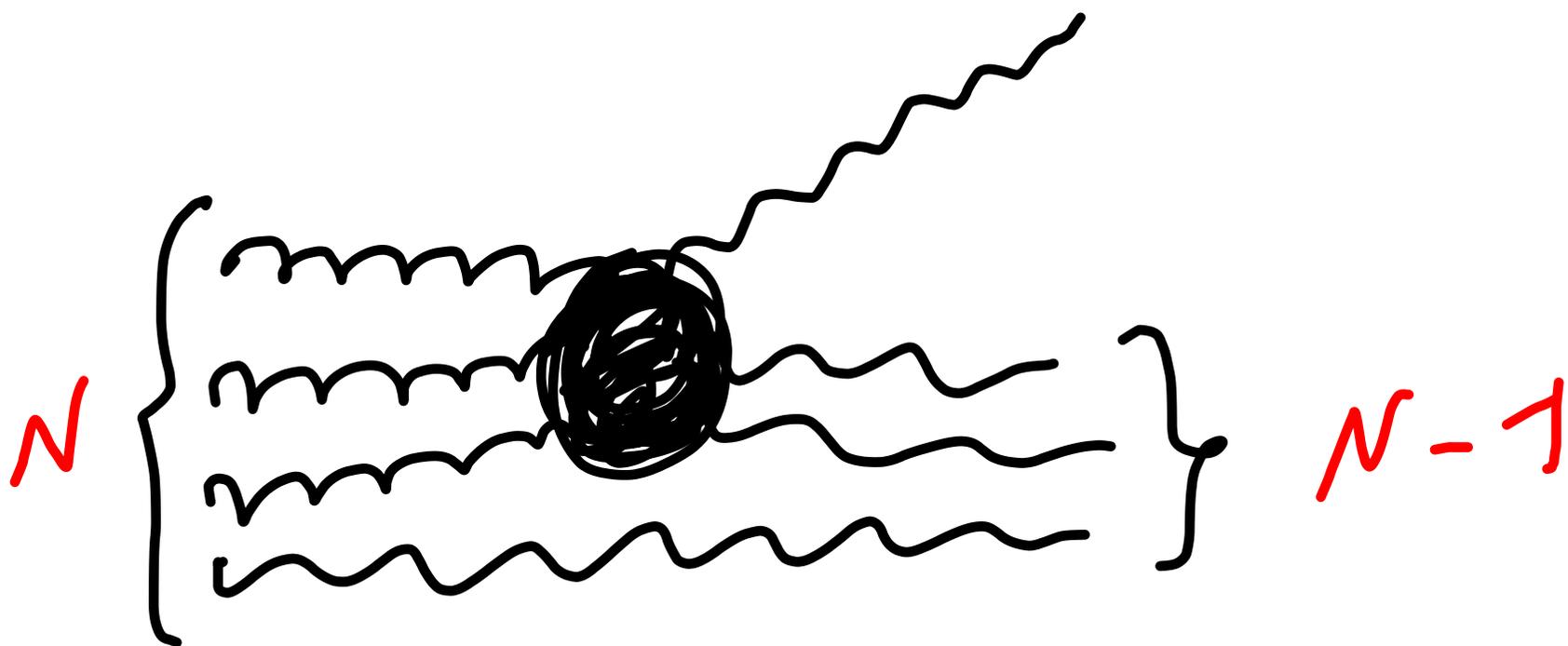
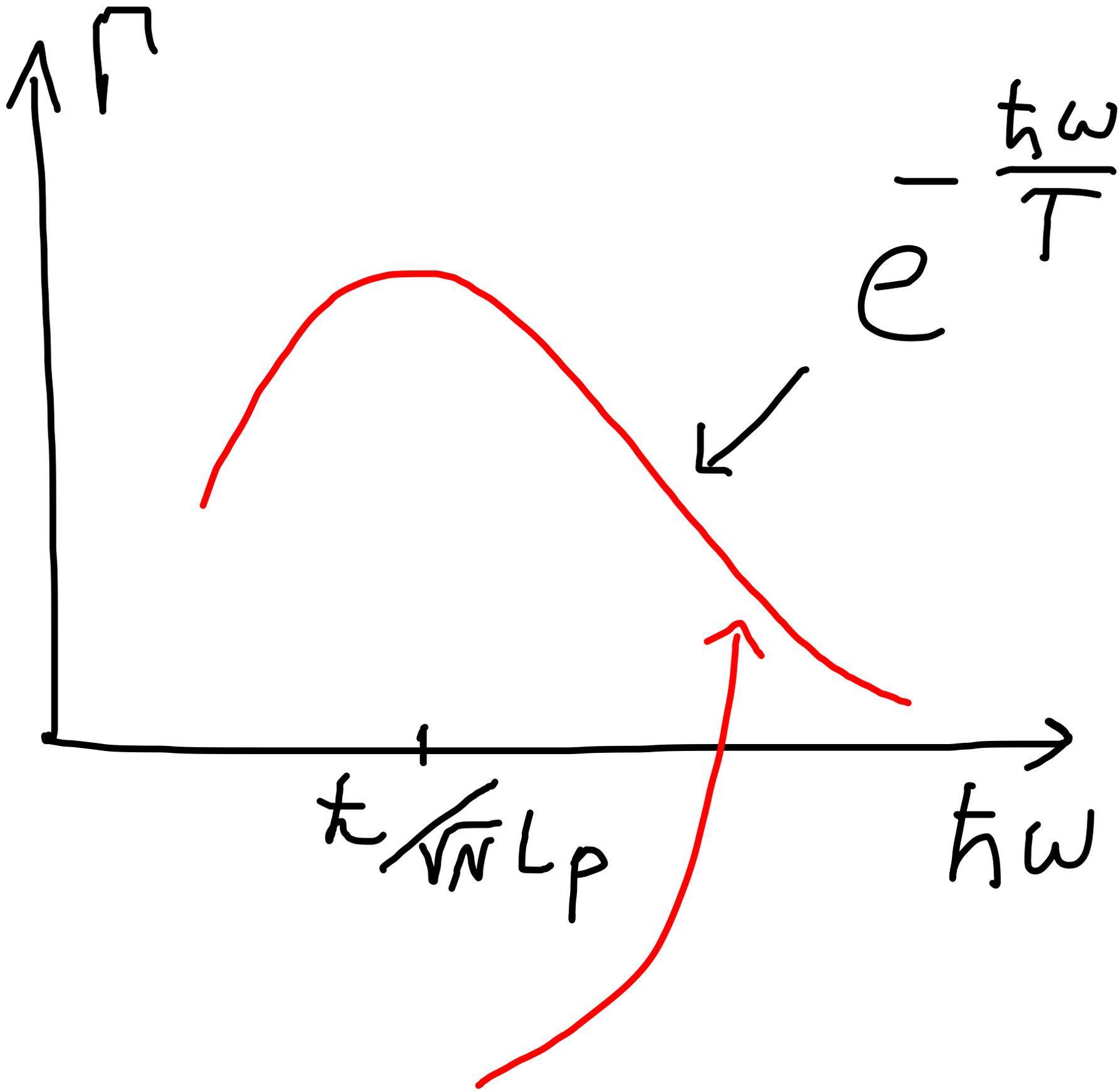
We get Stefan-Boltzmann law for Hawking evaporation

$$\dot{M} = -\frac{T^2}{\hbar}$$

We discover that thermality is an "optical illusion".

Spectrum is thermal because of the self-similarity of depletion, not because the source is hot.

The graviton condensate is cold!

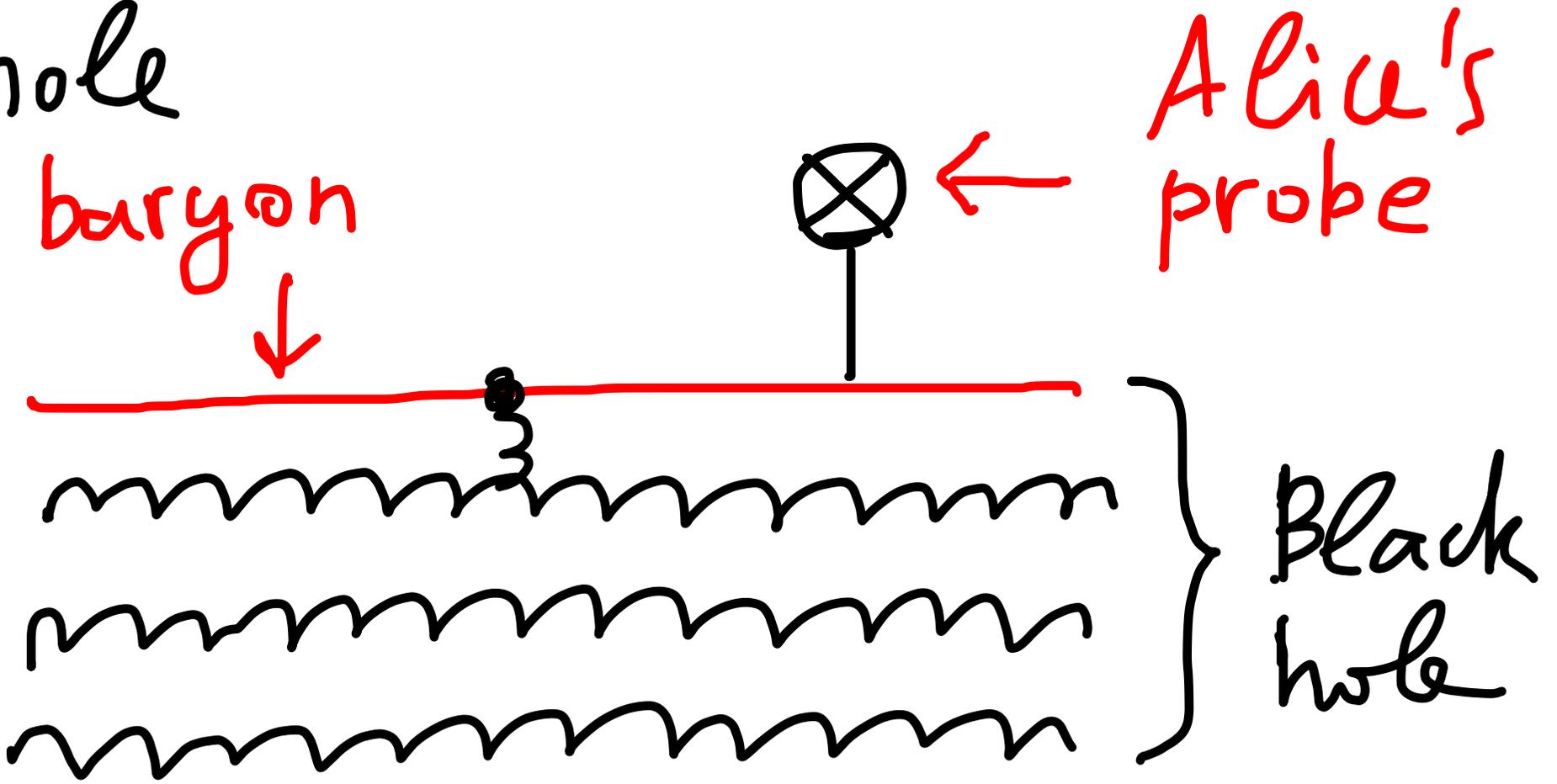


Another (false) artifact of semi-classical limit is the absence of hair.

In reality black holes carry a detectable hair as

$$\frac{N_B}{N} - \text{effect}$$

How Alice detect a
 paryonic hair of a black
 hole



$$\text{hair} = \frac{1}{\sqrt{N} L_p} \left(\frac{N_B}{N} \right)$$

For Astrophysical black holes (that carry large baryonic or leptonic charges) the hair can be an observable effect.

Outlook

Black hole's quantum portrait is a microscopic framework which allows to address questions that in the conventional treatment cannot even be formulated.

It demystifies the known semi-classical puzzles in black hole physics.

Among many potential applications is

Cosmology:

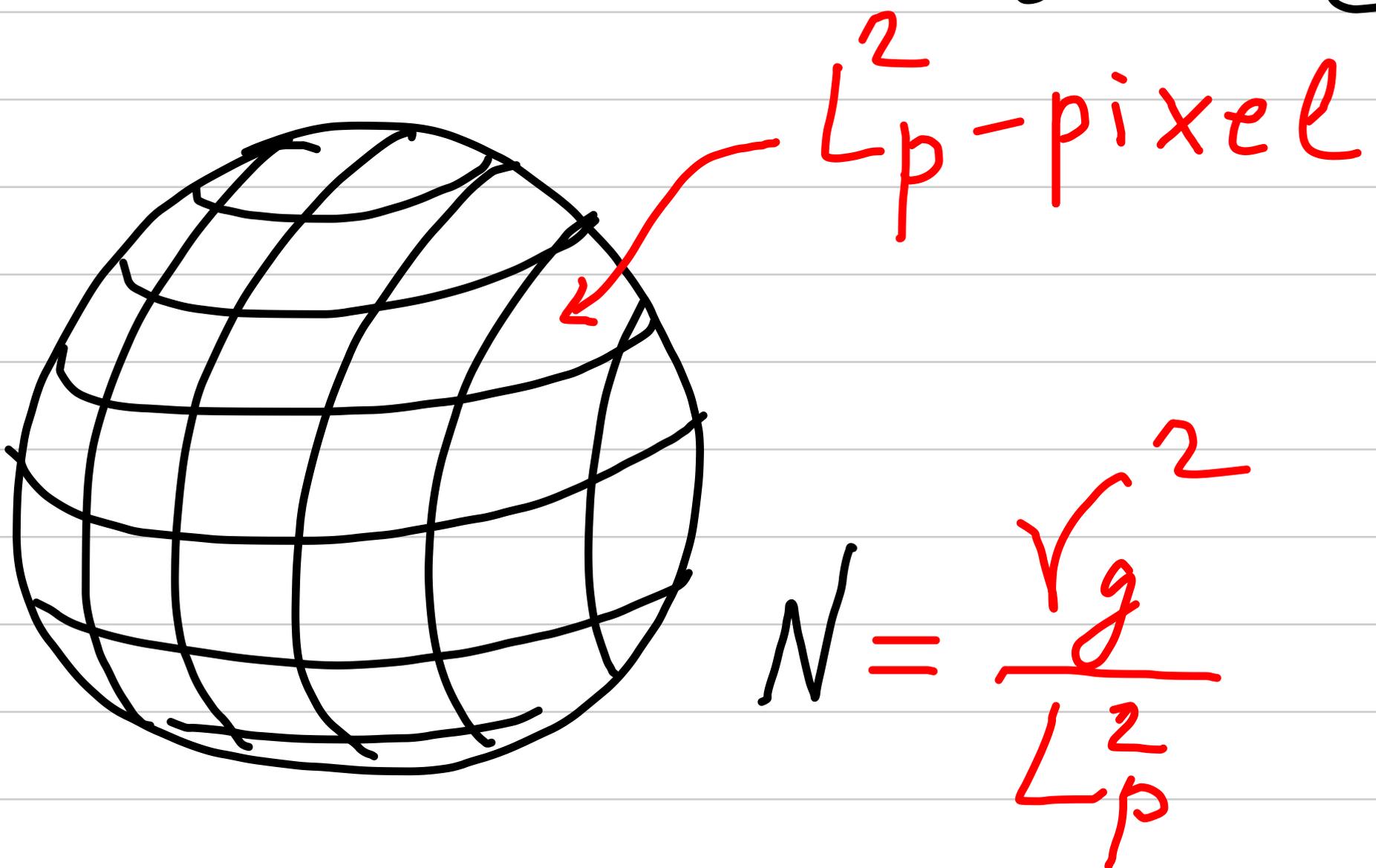
The Universe is the largest black hole we know.

It's a graviton condensate with

$$N \sim 10^{120}$$

We are learning that
overpacked systems
get oversimplified.

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Origin of holography



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