

Implications of a SM like Higgs for a natural NMSSM with low cutoff

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in collaboration with
T. Gherghetta, B. v. Harling, A. Medina
[to appear soon]

Outline

1 Introduction/Motivation

2 Framework

3 Results

4 Conclusions

Outline

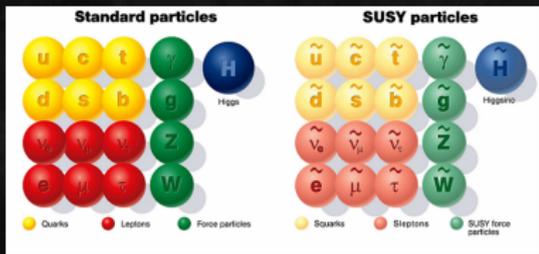
1 Introduction/Motivation

2 Framework

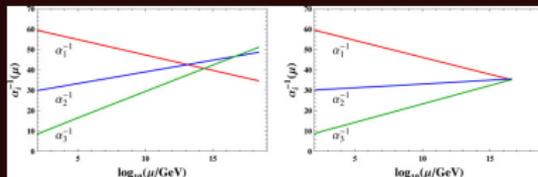
3 Results

4 Conclusions

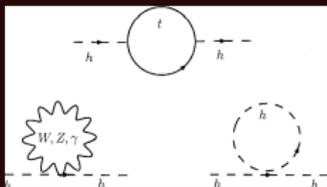
Low-energy Supersymmetry – MSSM



gauge coupling unification



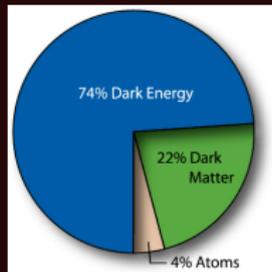
solution to hierarchy problem



$$\Delta m_H^2 = \frac{\lambda_S}{16\pi^2} \left[\Lambda_{UV}^2 - 2m_S^2 \ln \frac{\Lambda_{UV}}{m_S} + \dots \right]$$

$$\Delta m_H^2 = -\frac{|\lambda_f|^2}{16\pi^2} \left[\Lambda_{UV}^2 + \dots \right]$$

DM candidate



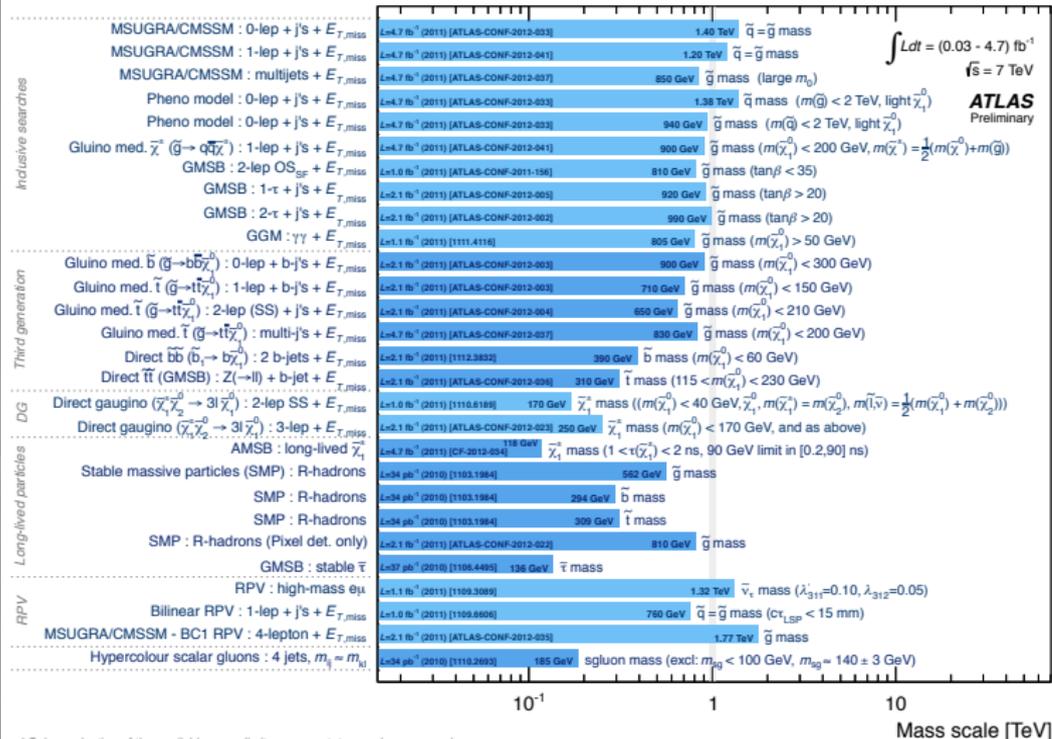
Higgs mass prediction

At tree level

$$m_{h^0} < m_Z |\cos 2\beta|$$

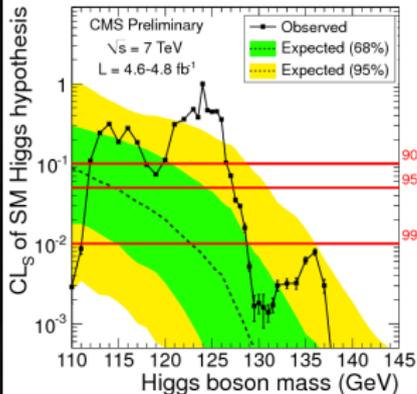
LHC constraints

ATLAS SUSY Searches* - 95% CL Lower Limits (Status: March 2012)



* Only a selection of the available mass limits on new states or phenomena shown

LHC constraints



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Search	Lower Limit	Notes
(2011) [ATLAS-CONF-2012-033]	1.40 TeV	$\tilde{q} = \tilde{g}$ mass
(2011) [ATLAS-CONF-2012-041]	1.20 TeV	$\tilde{q} = \tilde{g}$ mass
(2011) [ATLAS-CONF-2012-037]	850 GeV	\tilde{g} mass (large m_0)
(2011) [ATLAS-CONF-2012-033]	1.38 TeV	\tilde{q} mass ($m(\tilde{g}) < 2 \text{ TeV}$, light $\tilde{\chi}_1^0$)
(2011) [ATLAS-CONF-2012-033]	940 GeV	\tilde{g} mass ($m(\tilde{q}) < 2 \text{ TeV}$, light $\tilde{\chi}_1^0$)
(2011) [ATLAS-CONF-2012-041]	900 GeV	\tilde{g} mass ($m(\tilde{\chi}_1^0) < 200 \text{ GeV}$, $m(\tilde{\chi}_2^0) = \frac{1}{2}(m(\tilde{\chi}_1^0) + m(\tilde{g}))$)
(2011) [ATLAS-CONF-2011-156]	810 GeV	\tilde{g} mass ($\tan\beta < 35$)
(2011) [ATLAS-CONF-2012-003]	920 GeV	\tilde{g} mass ($\tan\beta > 20$)
(2011) [ATLAS-CONF-2012-002]	990 GeV	\tilde{g} mass ($\tan\beta > 20$)
(2011) [1111.4115]	805 GeV	\tilde{g} mass ($m(\tilde{\chi}_1^0) > 50 \text{ GeV}$)
(2011) [ATLAS-CONF-2012-003]	900 GeV	\tilde{g} mass ($m(\tilde{\chi}_1^0) < 300 \text{ GeV}$)
(2011) [ATLAS-CONF-2012-003]	710 GeV	\tilde{g} mass ($m(\tilde{\chi}_1^0) < 150 \text{ GeV}$)
(2011) [ATLAS-CONF-2012-004]	650 GeV	\tilde{g} mass ($m(\tilde{\chi}_1^0) < 210 \text{ GeV}$)
(2011) [ATLAS-CONF-2012-037]	830 GeV	\tilde{g} mass ($m(\tilde{\chi}_1^0) < 200 \text{ GeV}$)
(2011) [1112.3832]	390 GeV	\tilde{b} mass ($m(\tilde{\chi}_1^0) < 60 \text{ GeV}$)
(2011) [ATLAS-CONF-2012-036]	310 GeV	\tilde{t} mass
(2011) [1110.6189]	170 GeV	$\tilde{\chi}_1^+$ mass ($ m $)
(2011) [ATLAS-CONF-2012-023]	250 GeV	$\tilde{\chi}_1^+$ mass
(2011) [JCF-2012-050]	118 GeV	$\tilde{\chi}_1^+$ mass ($1 < \tau(\tilde{\chi}_1^+)$)
(2010) [1103.1984]	562 GeV	\tilde{t} mass
(2010) [1103.1984]	294 GeV	\tilde{b} mass
(2010) [1103.1984]	309 GeV	\tilde{t} mass
(2011) [ATLAS-CONF-2012-032]	-	\tilde{t} mass
(2010) [1108.4495]	136 GeV	$\tilde{\tau}$ mass
(2011) [1109.3095]	-	$\tilde{\tau}$ mass
(2011) [1109.6606]	78	$\tilde{\tau}$ mass
(2011) [ATLAS-CONF-2012-035]	-	$\tilde{\tau}$ mass
(2010) [1110.2693]	185 GeV	sgluon mass

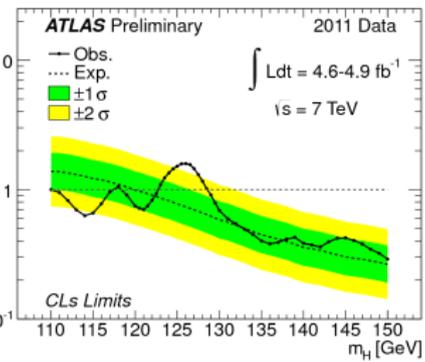
$$\int L dt = (0.03 - 4.7) \text{ fb}^{-1}$$

$$\sqrt{s} = 7 \text{ TeV}$$

ATLAS Preliminary

- TH: Direct $\tilde{b}\tilde{b}$ ($\tilde{b}_1 \rightarrow \tilde{b}_2^* \gamma$) : 2 b-jets + $E_{T, \text{miss}}$
- TH: Direct $\tilde{t}\tilde{t}$ (GMSB) : $Z(\rightarrow \text{ll}) + \text{b-jet} + E_{T, \text{miss}}$
- DG: Direct gaugino ($\tilde{\chi}_2^0 \tilde{\chi}_2^0 \rightarrow 3l \tilde{\chi}_1^0$) : 2-lep SS + $E_{T, \text{miss}}$
- DG: Direct gaugino ($\tilde{\chi}_2^0 \tilde{\chi}_2^0 \rightarrow 3l \tilde{\chi}_1^0$) : 3-lep + $E_{T, \text{miss}}$
- DG: AMSB : long-lived $\tilde{\chi}_1^0$
- Long-lived particles: Stable massive particles (SMP) : R-hadrons
- SMP : R-hadrons
- SMP : R-hadrons
- SMP : R-hadrons (Pixel det. only)
- RPV: GMSB : stable $\tilde{\tau}$
- RPV: high-mass $\tilde{\mu}_1$
- RPV: Bilinear RPV : 1-lep + $\tilde{\chi}_1^0$ + $E_{T, \text{miss}}$
- MSUGRA/CMSSM - BC1 RPV : 4-lepton + $E_{T, \text{miss}}$
- Hypercolour scalar gluons : 4 jets, $m_{\tilde{g}} = m_{\tilde{u}}$

95% CL Limit on $\sigma/\sigma_{\text{SM}}$



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Status of low-energy supersymmetry

- strong limits on squark and gluino masses (mostly in CMSSM/simplified models)
- relatively large Higgs mass
- large loop correction to Higgs mass needed
- fine tuning
- ⇒ Simplest SUSY models in bad shape



Barbieri-Giudice fine-tuning measure

[Barbieri/Giudice (1988)]

$$\Sigma_{\xi}^{\vee} \equiv \left| \frac{d \log m_Z^2}{d \log \xi} \right| = \left| \frac{\xi}{m_Z^2} \frac{dm_Z^2}{d\xi} \right| < \Sigma,$$

measures the sensitivity of the weak scale (or m_Z) to the dimensionful parameters ξ of the theory

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How viable is natural
low-energy SUSY?

Saving SUSY



Drop one of the assumptions in the (C)MSSM

- give up on small fine-tuning
- landscape, anthropic principle, ...

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- constraints relaxed, but generation of large Higgs mass still requires large stop contribution, although fine-tuning relaxed for light stops

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- proton decay, LSP no longer stable → no dark matter candidate

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- Extend MSSM to modify Higgs mass prediction, more specifically raise Higgs mass
- e.g. at tree level in NMSSM: introduce gauge singlet S

Literature: NMSSM and 125 GeV

- "A Natural SUSY Higgs Near 126 GeV"
L. J. Hall, D. Pinner and J. T. Ruderman.
arXiv:1112.2703 [hep-ph]
JHEP 1204, 131 (2012)
- "Higgs bosons near 125 GeV in the NMSSM with constraints at the GUT scale"
U. Ellwanger and C. Hugonie.
arXiv:1203.5048 [hep-ph]
- "The 125 GeV Higgs in the NMSSM in light of LHC results and astrophysics constraints"
D. A. Vasquez, G. Belanger, C. Boehm, J. Da Silva, P. Richardson and C. Wymant.
arXiv:1203.3446 [hep-ph]
- "A SM-like Higgs near 125 GeV in low energy SUSY: a comparative study for MSSM and NMSSM"
J. Cao, Z. Heng, J. M. Yang, Y. Zhang and J. Zhu.
arXiv:1202.5821 [hep-ph]
JHEP 1203, 086 (2012)
- "NMSSM Higgs Benchmarks Near 125 GeV"
S. F. King, M. Muhlleitner and R. Nevzorov.
arXiv:1201.2671 [hep-ph]
Nucl. Phys. B 860, 207 (2012)
- "The Constrained NMSSM and Higgs near 125 GeV"
J. F. Gunion, Y. Jiang and S. Kraml.
arXiv:1201.0982 [hep-ph]
Phys. Lett. B 710, 454 (2012)
- "A Higgs boson near 125 GeV with enhanced di-photon signal in the NMSSM"
U. Ellwanger.
arXiv:1112.3548 [hep-ph]
JHEP 1203, 044 (2012)
- "The fine-tuning of the generalised NMSSM"
G. G. Ross and K. Schmidt-Hoberg.
arXiv:1108.1284 [hep-ph] "The generalised NMSSM at one loop: fine tuning and phenomenology"
G. G. Ross, K. Schmidt-Hoberg and F. Staub.
arXiv:1205.1509 [hep-ph]
- ...

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NMSSM

Superpotential

$$W_{NMSSM} = \lambda S H_d H_u + \frac{\kappa}{3} S^3$$

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Soft breaking terms

$$V_{soft} = m_{H_d}^2 |H_d|^2 + m_{H_u}^2 |H_u|^2 + m_S^2 |S|^2 - (a_\lambda S H_d H_u + \frac{a_\kappa}{3} S^3 + h.c.)$$

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Bound on lightest Higgs mass m_h

$$m_h^2 \leq m_Z^2 \cos^2 2\beta + \lambda^2 v^2 \sin^2 2\beta \rightarrow \begin{cases} m_Z^2 \cos^2 2\beta & \text{large } \tan \beta \\ \lambda^2 v^2 \sin^2 2\beta & \text{small } \tan \beta \end{cases}$$

- No gain for large $\tan \beta$ compared to MSSM

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- No gain for large $\tan \beta$ compared to MSSM
 - small $\tan \beta \Rightarrow$ large $\lambda \gtrsim 0.7$ or additional stop loop contribution
- \rightarrow Landau pole below GUT scale \rightarrow low cutoff required

NMSSM with low cutoff $\Lambda = 10 \text{ TeV}$

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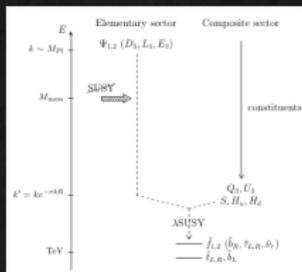
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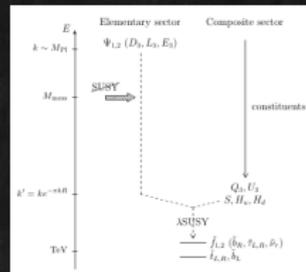
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Our Study

- Aim: Find the "Golden Region" of small fine-tuning (better 10%), i.e. the natural region of parameter space in the NMSSM
- assuming "SM-like" Higgs with $m_h \approx 125 \text{ GeV}$ [(124 – 126) GeV]

Our Study

Assumptions

- Cutoff $\Lambda = 10 \text{ TeV}$
- no CP violation
- First two generations of squarks and all sleptons decoupled
- impose LEP bounds on particle masses
- Neutralino LSP (EW gauginos not decoupled)
- no invisible Higgs decays

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Analysis

- Markov-Chain Monte-Carlo (MCMC)
- with modified version of NMHDECAY [Ellwanger, Gunion, Hugonie]
- No Landau pole below the cutoff scale $\Lambda = 10$ TeV
- Electroweak precision test: S and T taking into account
 - Neutralino-Chargino sector
 - Higgs sector
 - T in stop-sbottom sector (S is small [Barbieri, Hall, Nomura, Rychkov])

Analysis

- Input values are specified at

$$M_{SUSY}^2 = \frac{m_{\tilde{t}_1}^2 + m_{\tilde{t}_2}^2 + m_{\tilde{b}_1}^2 + m_{\tilde{b}_2}^2}{4}$$

- range of input values:

$$0 \text{ TeV} < M_1, M_2 < 3 \text{ TeV}$$

$$700 \text{ GeV} < M_3 < 3 \text{ TeV}$$

$$0 \text{ TeV} < m_{Q_3}, m_{u_3}, m_{d_3} < 3 \text{ TeV}$$

$$-3 \text{ TeV} < A_\lambda, A_\kappa, A_t, A_b < 3 \text{ TeV}$$

$$\lambda > 0$$

$$a_\kappa = \kappa A_\kappa \geq 0$$

$$\tan \beta > 0.5$$

Fine-Tuning

Loop corrected Minimisation Conditions

$$v_u(Q) \left(m_{H_u}^2 + \mu_{\text{eff}}^2 + \lambda^2 v_d(Q)^2 + \frac{g_1^2 + g_2^2}{4} (v_u(Q)^2 - v_d(Q)^2) \right) = v_d(Q) \mu_{\text{eff}} B_{\text{eff}} - \frac{1}{2} \frac{\partial \Delta V_{\text{eff}}}{\partial v_u(Q)}$$

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$$s(Q) \left(m_S^2 + \kappa A_\kappa s(Q) + 2\kappa^2 s^2(Q) + \lambda^2 (v_u^2(Q) + v_d^2(Q) - 2\lambda \kappa v_u(Q) v_d(Q)) \right) = \lambda A_\lambda v_u(Q) v_d(Q) - \frac{1}{2} \frac{\partial \Delta V_{\text{eff}}}{\partial s(Q)}$$

with $\mu_{\text{eff}} = \lambda s(Q)$, $B_{\text{eff}} = A_\lambda + \kappa s(Q)$ and

$$v_u(Q) = v_u / \sqrt{Z_{H_u}}$$

$$v_d(Q) = v_d / \sqrt{Z_{H_d}}$$

$$s(Q) = s / \sqrt{Z_s}$$

- Three minimisation conditions result in linear system of equations for $\frac{dv^2}{d\xi}$, $\frac{d \tan \beta}{d\xi}$ and $\frac{ds}{d\xi}$.
- which determines the usual fine-tuning: $\Sigma_\xi^v \equiv \left| \frac{\xi^2}{v^2} \frac{dv^2}{d\xi^2} \right| < \Sigma$

Fine-Tuning 2

- Are the parameters at the SUSY scale fine-tuned?

Fine-Tuning 2

Quantum Corrections between SUSY scale and cutoff Λ

For example stop contribution to m_{H_u} :

$$\delta m_{H_u}^2 \approx \frac{3}{16\pi^2} \left(m_{Q_3}^2 + m_{u_3}^2 + A_t^2 \right) \log \left[\frac{\Lambda^2}{M_{SUSY}^2} \right]$$

- Are the parameters at the SUSY scale fine-tuned?
- Consider fine-tuning measure for each dimensionful coupling

$$\text{e.g. } \sum \frac{m_{H_u}^2}{m_{Q_3}^2} \equiv \frac{m_{Q_3}^2}{m_{H_u}^2} \frac{d\delta m_{H_u}^2}{dm_{Q_3}^2}$$

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$$\text{e.g. } \Sigma_{m_{Q_3}^2}^{m_{H_u}^2} \equiv \frac{m_{Q_3}^2}{m_{H_u}^2} \frac{d\delta m_{H_u}^2}{dm_{Q_3}^2}$$

- Fine-tuning of EW scale with respect to $m_{Q_3}^2$:

$$\left| \Sigma_{m_{Q_3}^2}^v \right| \approx \left| \Sigma_{m_{Q_3}^2}^{v,MC} + \sum_i \frac{m_{Q_3}^2}{v^2} \frac{dv^2}{d\xi_i} \frac{d\delta\xi_i}{dm_{Q_3}^2} \right| = \left| \Sigma_{m_{Q_3}^2}^{v,MC} + \sum_i \Sigma_{m_{Q_3}^2}^{\xi_i} \Sigma_{\xi_i}^{v,MC} \right| < \Sigma$$

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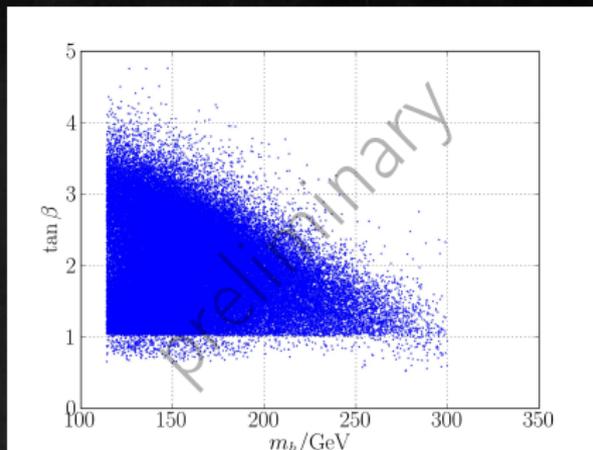
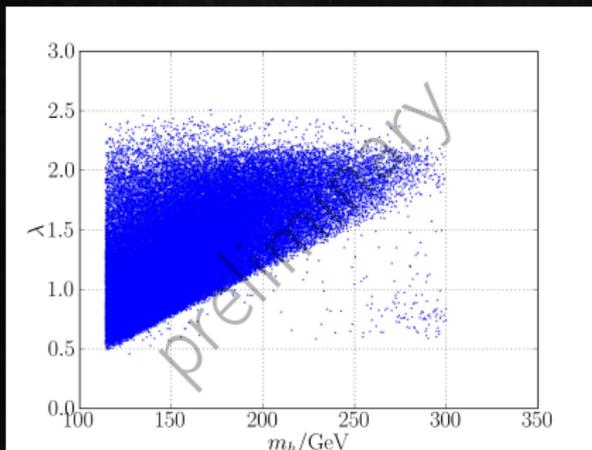
- Hence, the overall fine-tuning of the EW scale is given by

$$\Sigma^v \equiv \max_i \left| \Sigma_{\xi_i}^v \right|$$

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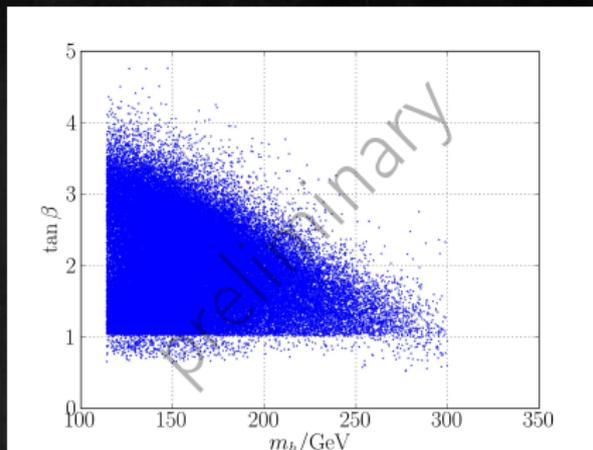
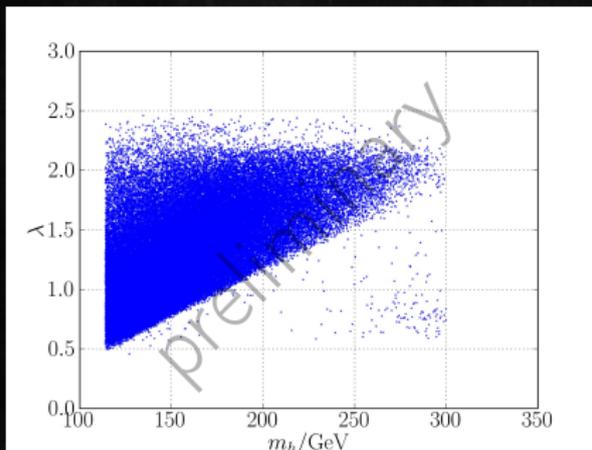
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- Higgs defined as CP even scalar with largest H_u component (not necessarily lightest CP even scalar)

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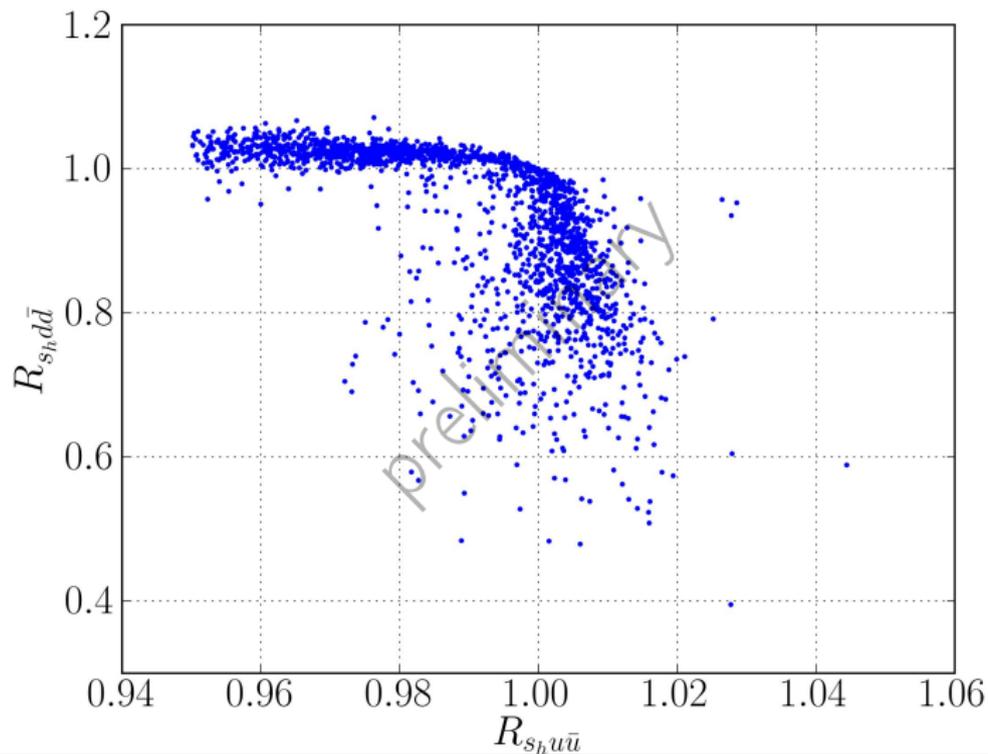


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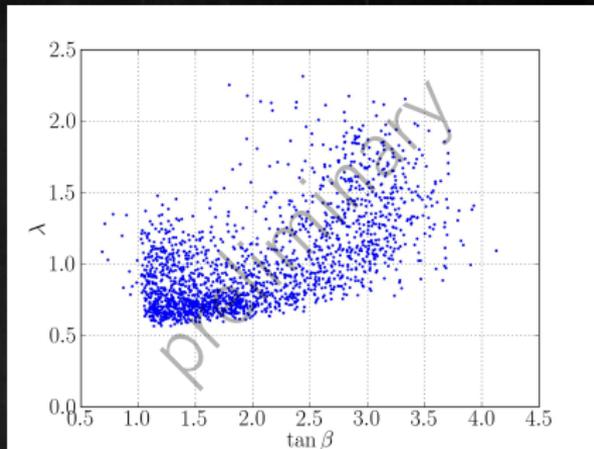
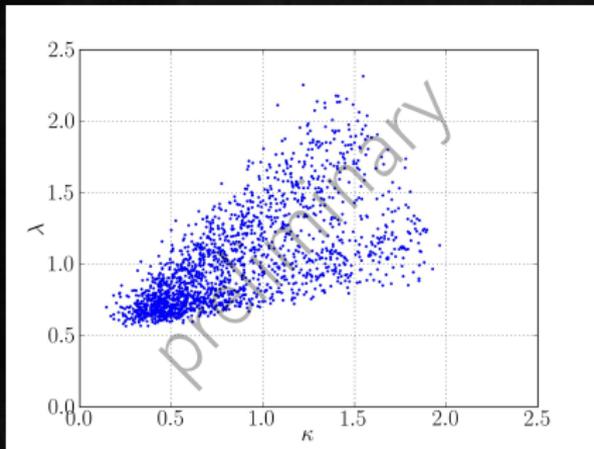
In the following

- $124 \text{ GeV} < m_h < 126 \text{ GeV}$
- "SM-like" Higgs: $|R_{ZZh} - 1| < 0.05$ and $|R_{u\bar{u}h} - 1| < 0.05$

Higgs couplings



$\lambda - \kappa$



- $0.6 \lesssim \lambda \lesssim 2.3$
- $\lambda \sim \kappa$
- running of κ stronger than $\lambda \Rightarrow \max \kappa < \max \lambda$
- $\tan \beta$ bounded due to contribution to T parameter

Neutralinos and T parameter

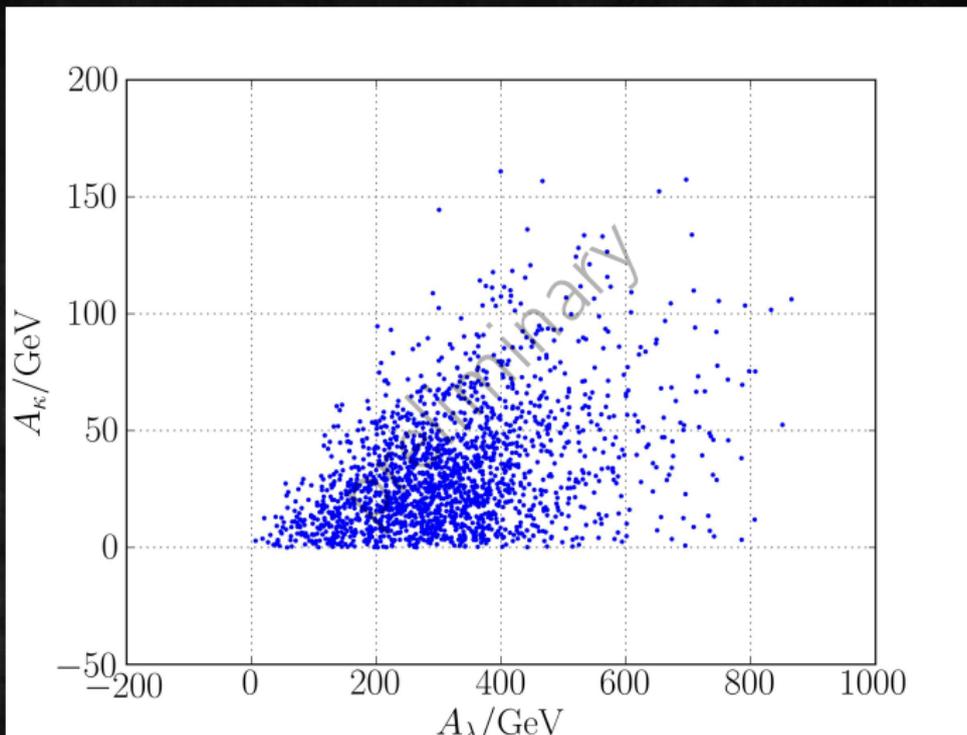
Neutralinos

$$\begin{pmatrix} M_1 & 0 & -\cos\beta \sin\theta_W m_Z & \sin\beta \sin\theta_W m_Z & 0 \\ \cdot & M_2 & \cos\beta \cos\theta_W m_Z & -\sin\beta \cos\theta_W m_Z & 0 \\ \cdot & \cdot & 0 & -\mu & -\lambda v \sin\beta \\ \cdot & \cdot & -\mu & 0 & -\lambda v \cos\beta \\ \cdot & \cdot & \cdot & \cdot & -2\frac{\kappa}{\lambda}\mu \end{pmatrix}$$

in gauge-eigenbasis $\psi^0 = (\tilde{B}, \tilde{W}^3, \tilde{H}_d^0, \tilde{H}_u^0, \tilde{S})$

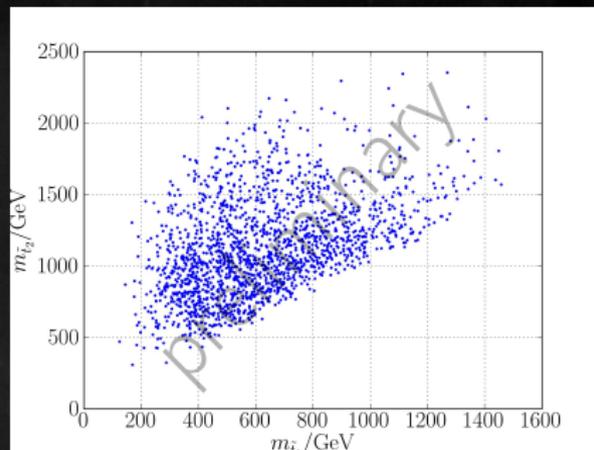
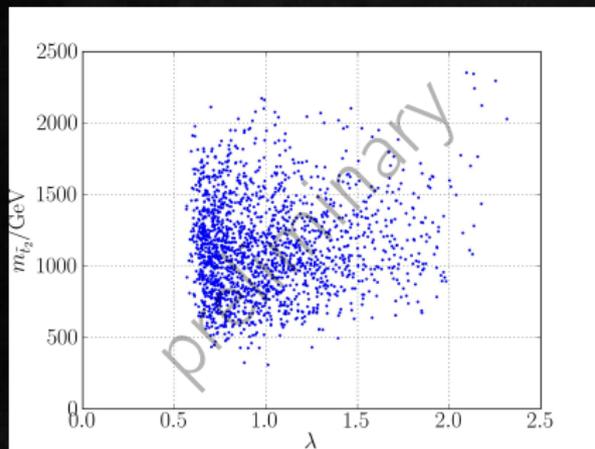
- $\tan\beta \lesssim 5$ limited by stop contribution to T [Barbieri, Hall, Nomura, Rychkov]
- Higgsino-Singlino mixing restricted

$$A_\lambda - A_\kappa$$



- small A terms: $A_\kappa \lesssim A_\lambda < 1 \text{ TeV}$

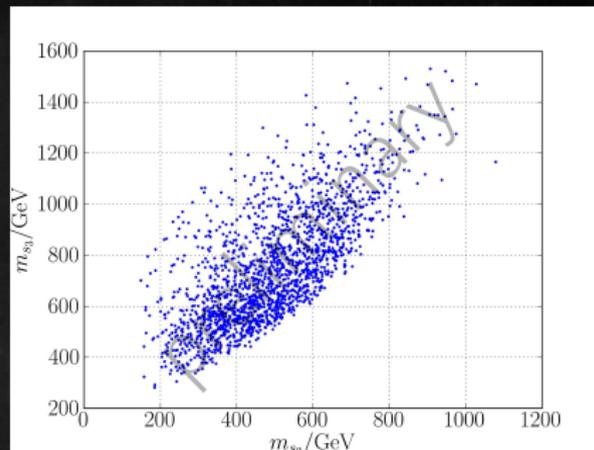
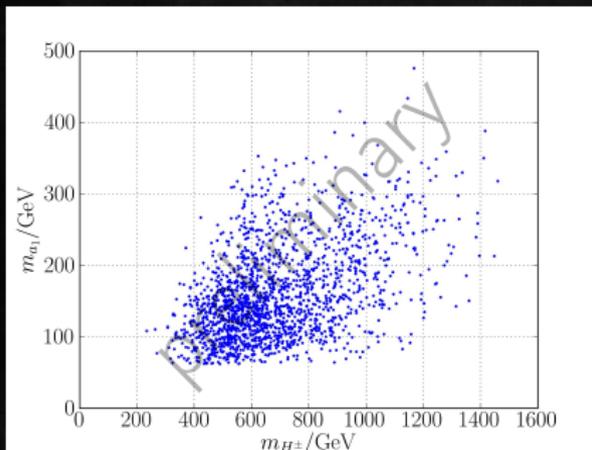
stop masses



- $m_{\tilde{t}_2} \lesssim 2.4 \text{ TeV}$
- natural region with no light stops
- consistent with results obtained for

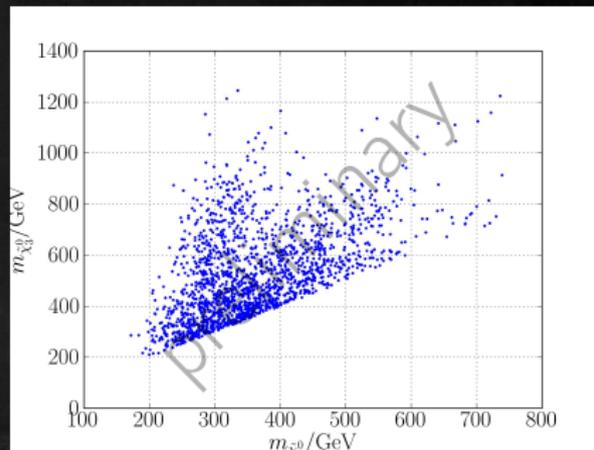
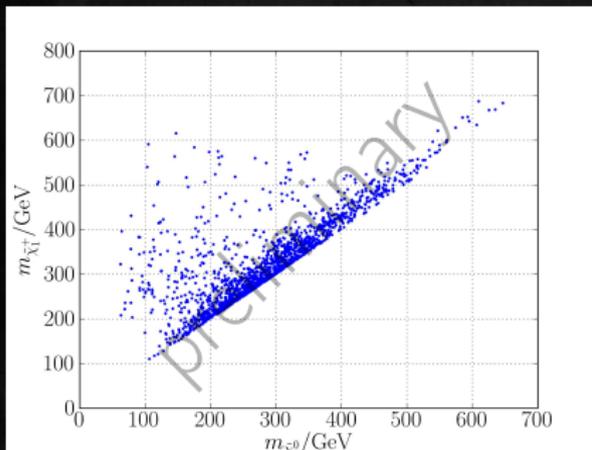
$$W_{NMSSM} = \lambda S H_u H_d + \hat{\mu} H_u H_d + \frac{M}{2} S^2 \text{ [Hall, Pinner, Ruderman (2011)]}$$

What is light? – Higgs sector



- one CP odd Higgs < 500 GeV
- at least one CP even Higgs $\lesssim 1$ TeV
- $b \rightarrow s\gamma$ will constrain light charged Higgs mass, but it is already mostly above 400 GeV

What is light? – Neutralino-Chargino sector



- several light neutralinos (< 1 TeV)
- lightest chargino $m_{\tilde{\chi}_1^\pm} < 700$ GeV

Outline

- 1 Introduction/Motivation
- 2 Framework
- 3 Results
- 4 Conclusions

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- stop masses can be large $\sim 2 \text{ TeV}$, while maintaining small fine-tuning
- Possibly no coloured particles below 1 TeV
- ⇒ Testing the whole parameter space of natural SUSY at the LHC not easy

Thank you very much for your attention.